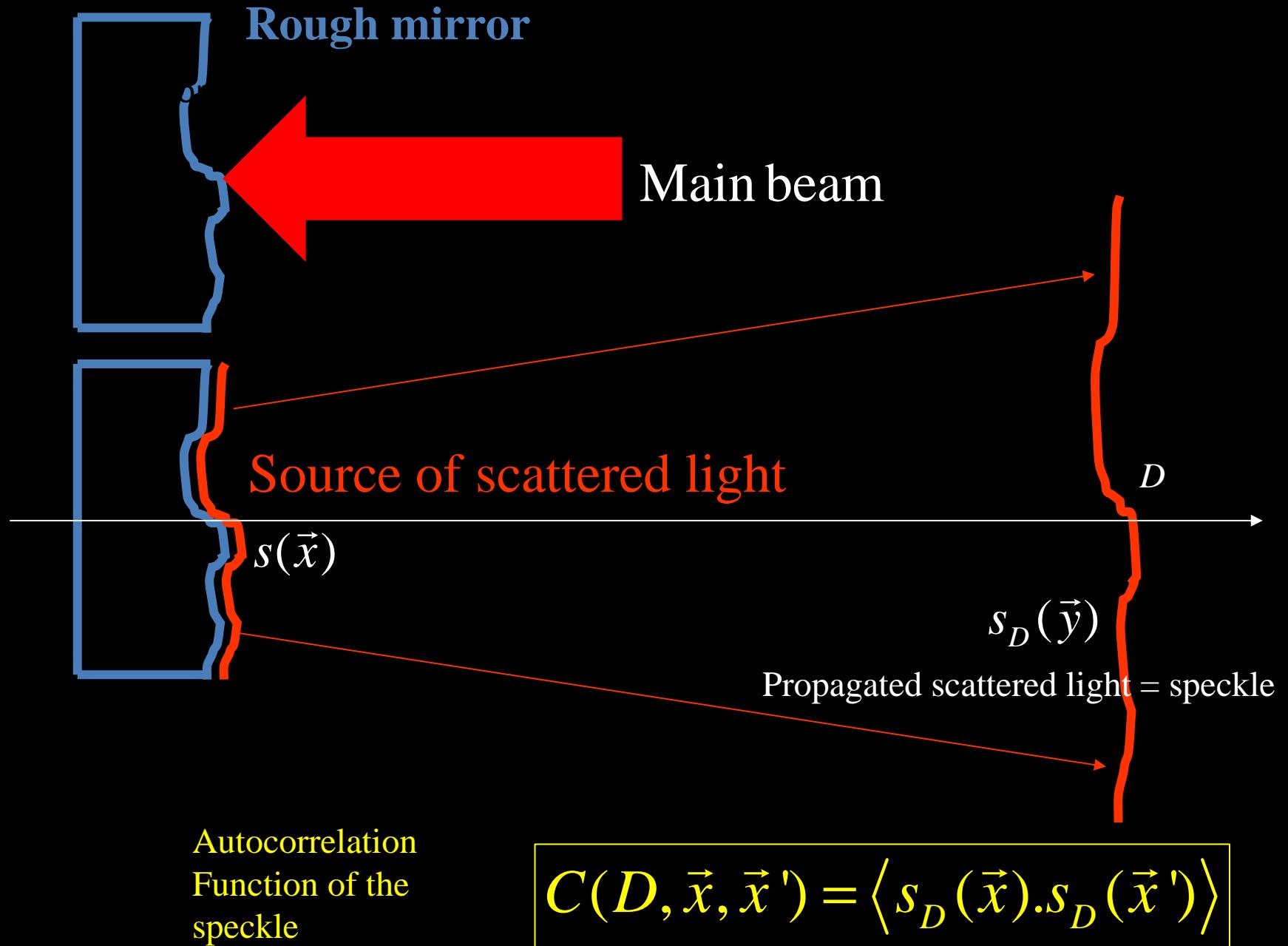




# Baffles and cryostat : Scattered light noise

Jean-Yves VINET



## Hypotheses :

Distribution of scattered light vs scattering angle :

$$p(\theta) \approx \frac{\kappa}{\theta^2} \quad \text{For angles much larger than } \theta_g = \frac{\lambda}{\pi w_0}$$

LSD of seismic motion :

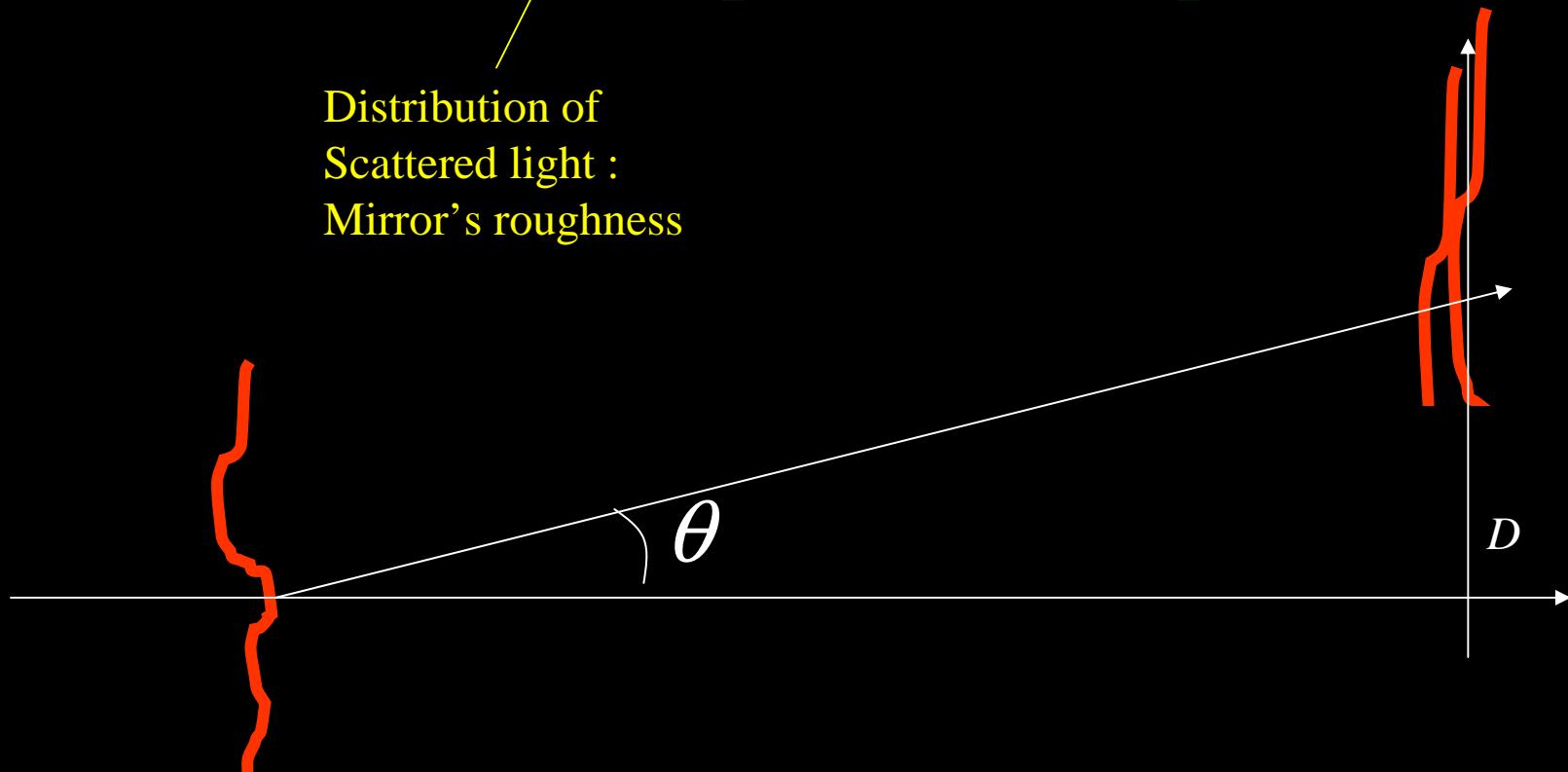
$$\delta x(f) \approx 10^{-8} \text{ m}/\sqrt{\text{Hz}} \left[ \frac{10 \text{ Hz}}{f} \right]^2$$

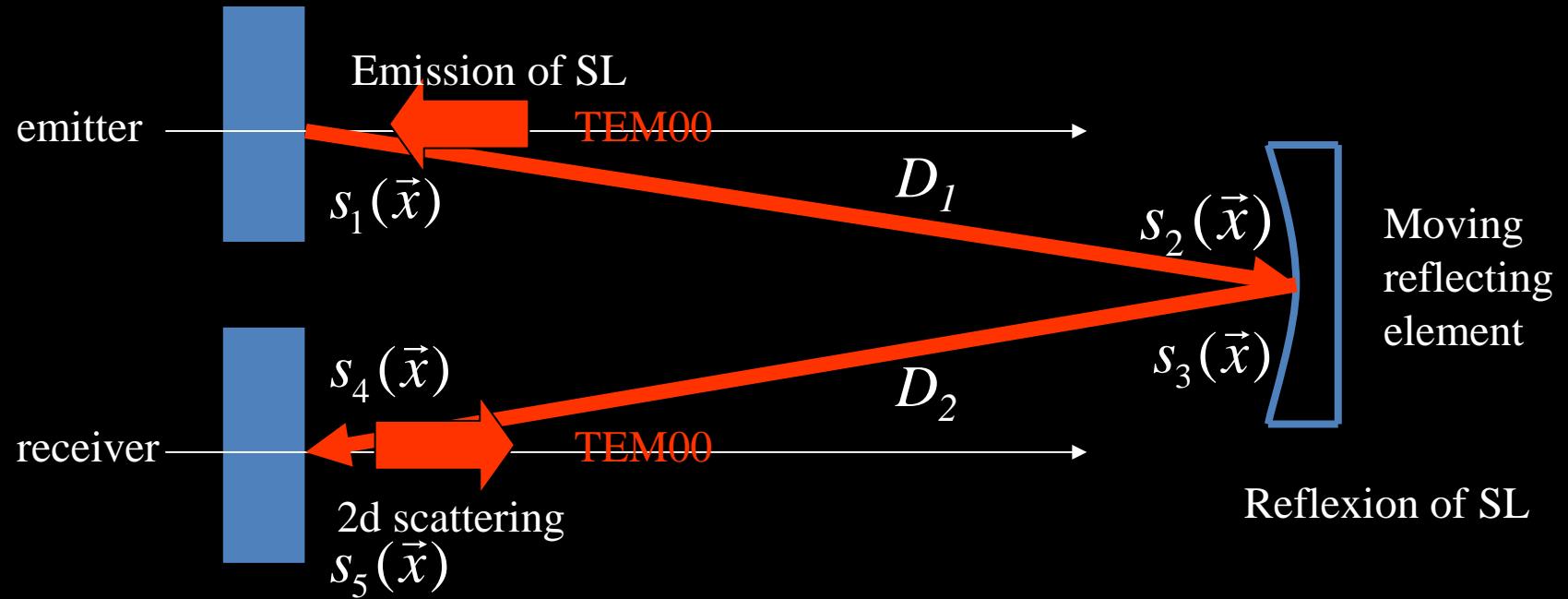
Integrated  
Scatt. losses

$$C(D, \vec{x}, \vec{x}') = \frac{\varepsilon}{2\pi D^2} p(\theta) \exp\left[-\frac{1}{2}\left(kw_0 \frac{\vec{x}-\vec{x}'}{2D}\right)^2\right] \exp\left[ik \frac{x^2 - x'^2}{2D}\right]$$

Memory of the stored mode

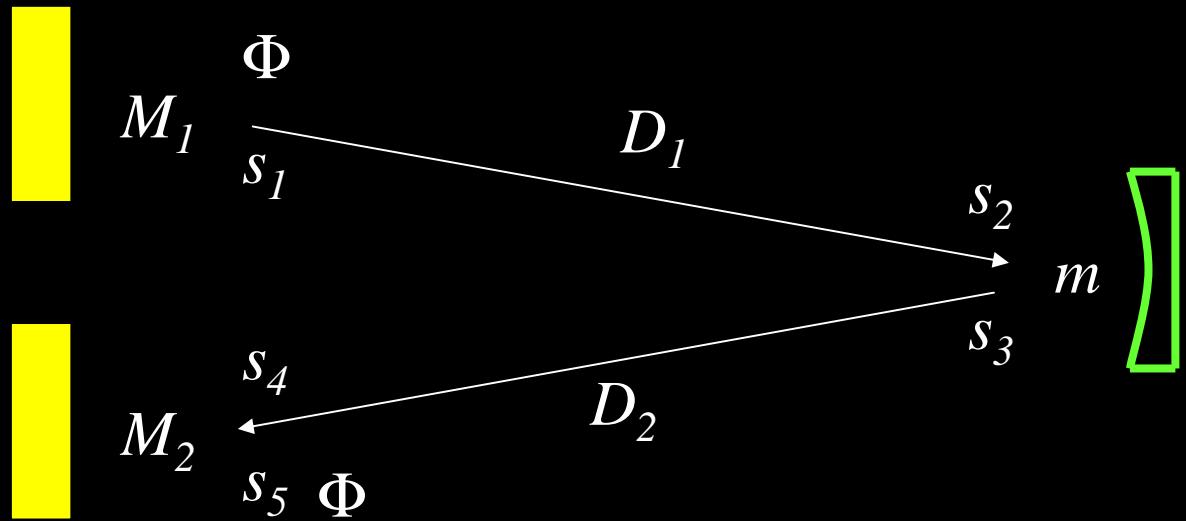
Distribution of  
Scattered light :  
Mirror's roughness





Spurious (noisy) phase modulation caused by the coupling :

$$\gamma = \langle TEM\ 00, s_5 \rangle$$



$$\gamma = \langle \Phi, s_5 \rangle = \langle \Phi, M_2 P_2(m.s_3) \rangle = \langle \Phi, M_2 P_2[m.P_1(M_1\Phi)] \rangle$$

Coupling coefficient with TEM00 :

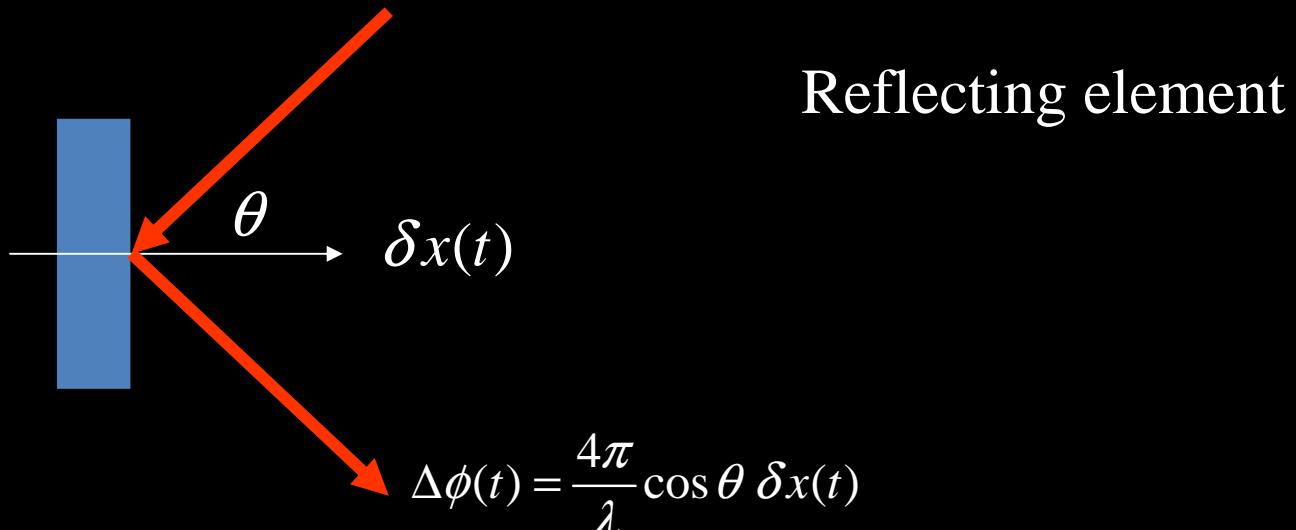
$$\gamma = e^{i\Delta(t)} \int m(\vec{x}) s_{D_1}(\vec{x}) s_{D_2}(\vec{x}) d\vec{x}$$

Expectation value :

$$\langle \mathcal{W}^* \rangle = \int m(\vec{x}) m(\vec{x}')^* \left\langle s_{D_1}(\vec{x}) s_{D_1}(\vec{x}')^* \right\rangle \left\langle s_{D_2}(\vec{x}) s_{D_2}(\vec{x}')^* \right\rangle d\vec{x} d\vec{x}'$$

In terms of the speckle autocorrelation :

$$\Gamma \equiv \langle \mathcal{W}^* \rangle = \int m(\vec{x}) m(\vec{x}')^* \cdot \textcolor{red}{C}(D_1, \vec{x}, \vec{x}'). \textcolor{blue}{C}(D_2, \vec{x}, \vec{x}') d\vec{x} d\vec{x}'$$



If  $\delta x \ll \lambda$

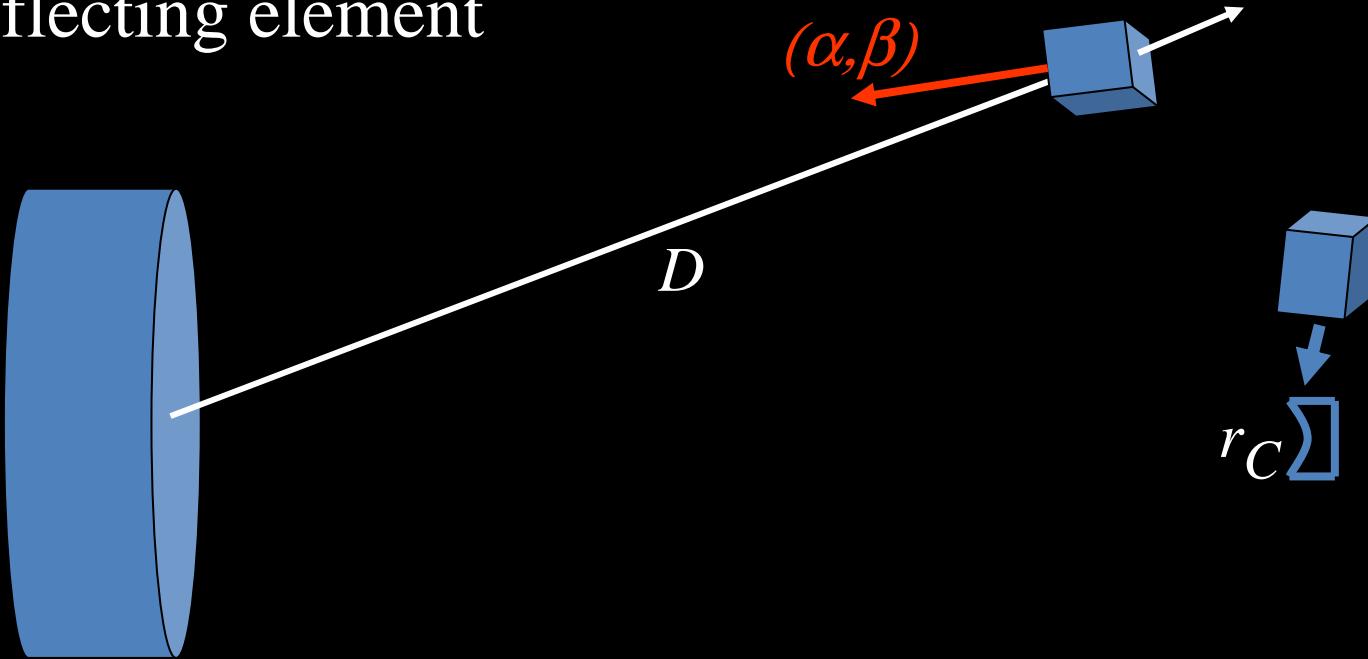
PSD of SL-noise :

$$h(f) = \frac{\Gamma^{1/2}}{\sqrt{2}} \cos \theta \frac{\delta x(f)}{L}$$

## Possible cases of coupling

- 1) Direct reflection from elementary spots  
 $IM \rightarrow IM$  and  $EM \rightarrow EM$
- 1) Transmission  $EM \rightarrow IM$  and  $IM \rightarrow EM$  via inner baffle edge
- 3) Transmission  $EM \rightarrow IM$  and  $IM \rightarrow EM$  via spots on the inner edge

Direct reflection  
By reflecting element



Relevant parameters :

- distance from mirror
- Misalignment  $(\alpha, \beta)$
- Mismatching  $r_C/D$
- Surface of the element

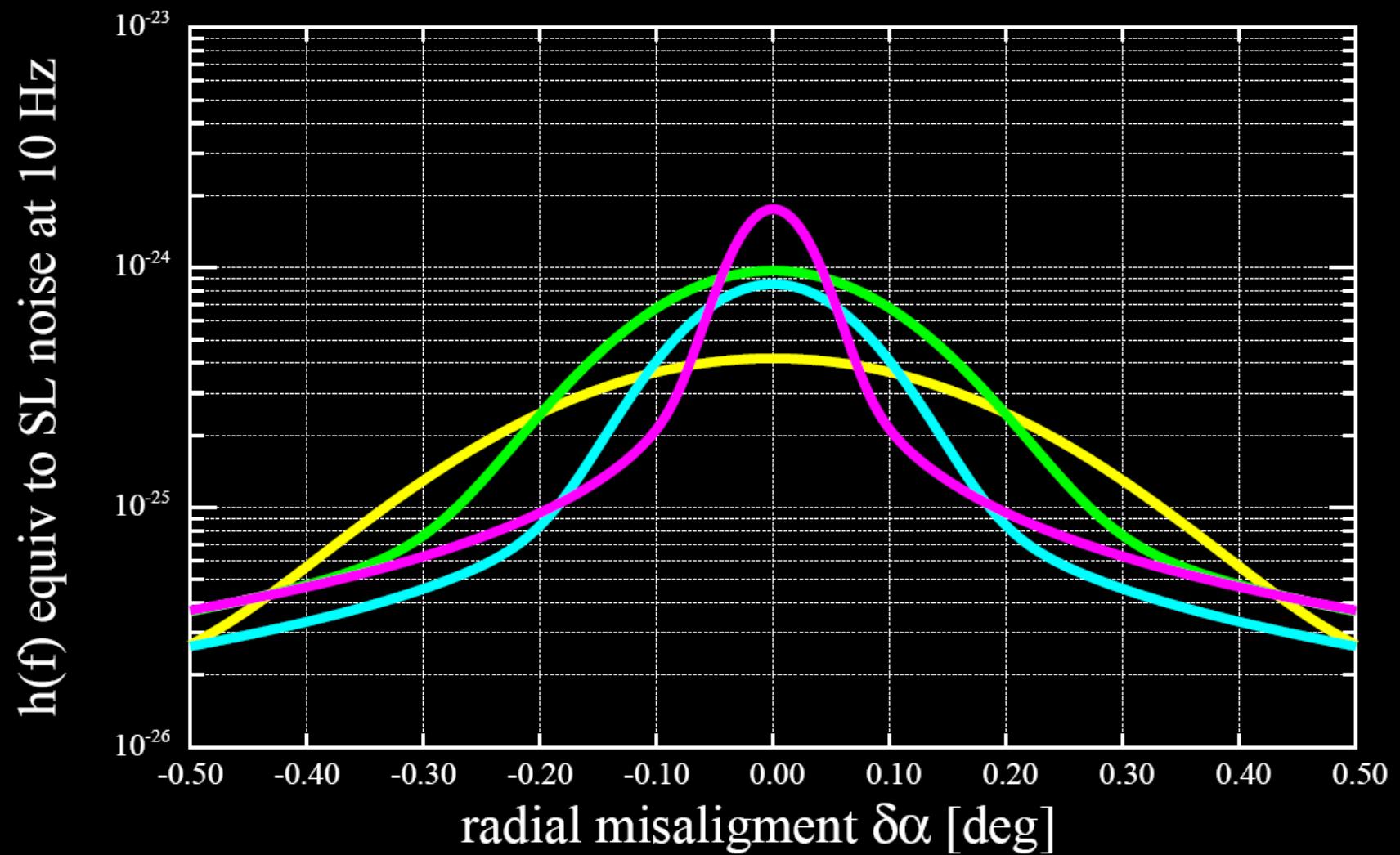
$$\begin{aligned}\beta &= 0 \\ r_C &= D \\ S &= 10^{-6} \text{ m}^2\end{aligned}$$

$$D=0.9 \text{ m}, \Delta\alpha=0.46^\circ$$

$$D=2.2 \text{ m}, \Delta\alpha=0.19^\circ$$

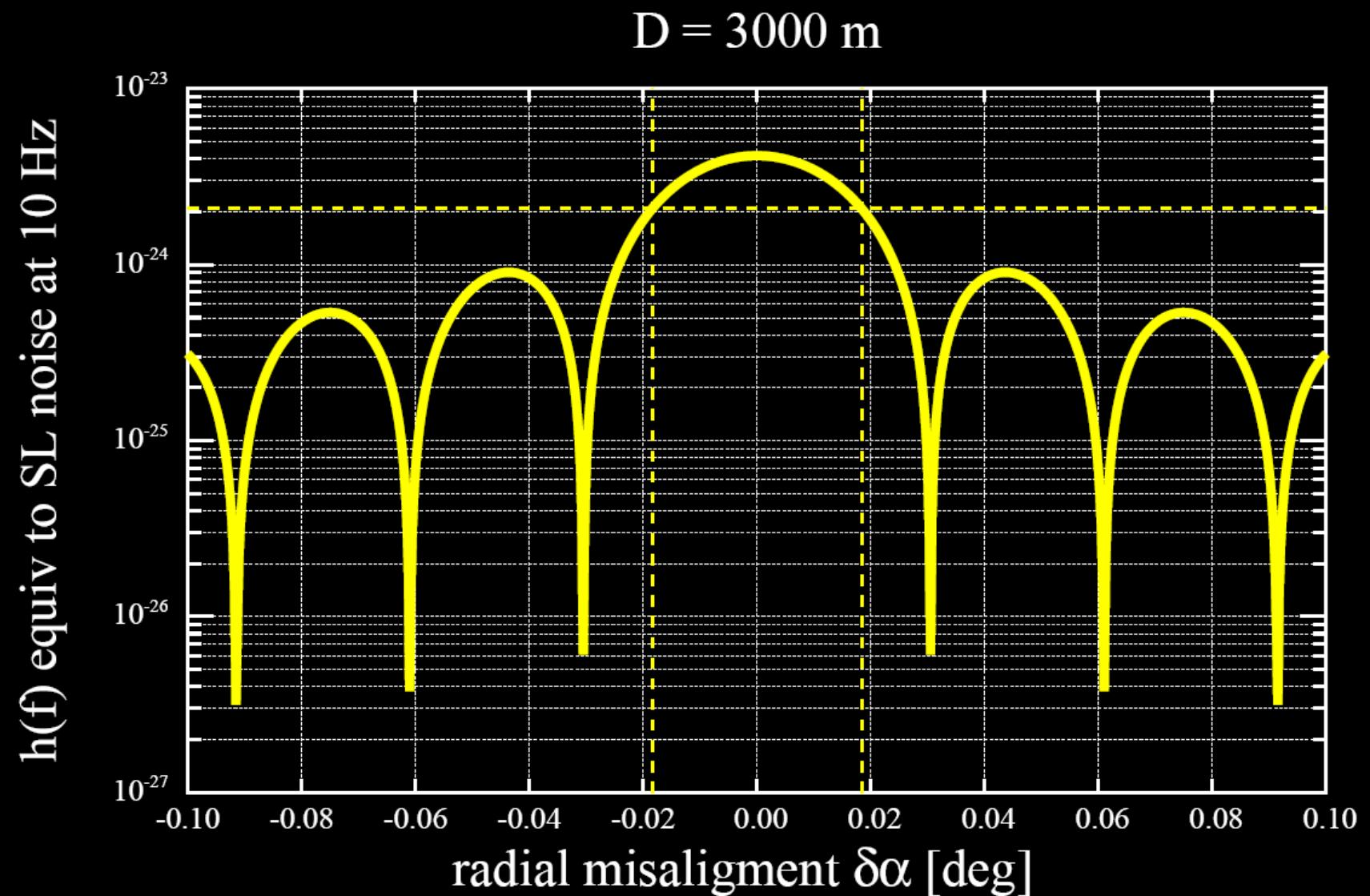
$$D=1.5 \text{ m}, \Delta\alpha=0.28^\circ$$

$$D=5 \text{ m}, \Delta\alpha=0.09^\circ$$

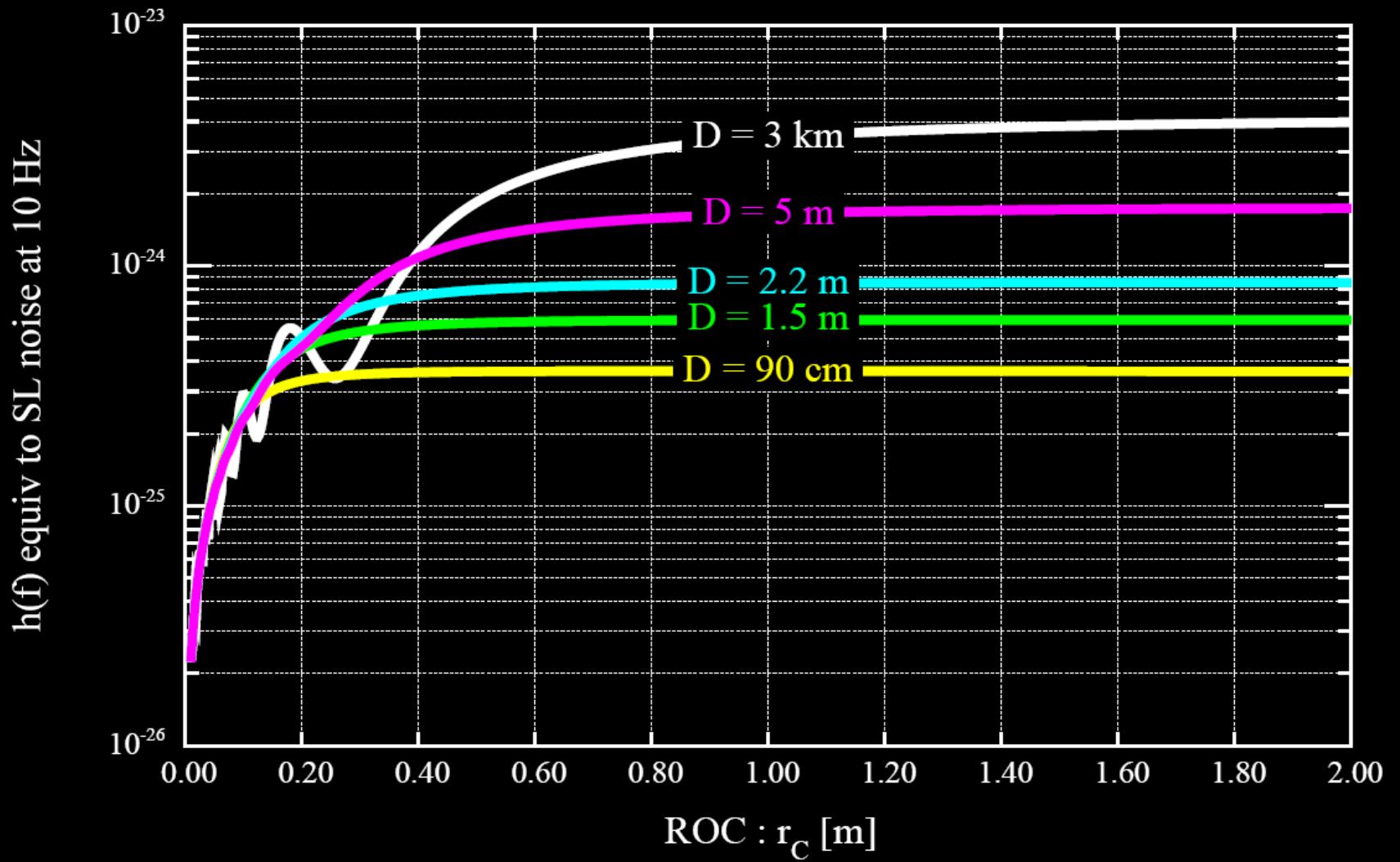


Case of input zone baffle → End mirror

$D=3000$  m,  $\Delta\alpha=0.04^\circ$



# ROC of the reflecting element



Conclusions for small reflecting defects :

- Unsensitive to ROC of the zone
- Sensitive to the distance
- Sensitive to the pointing error

## Summary

$$D = 0.9 \text{ m} : h(f) = 3.65 \cdot 10^{-25} \left[ \frac{10 \text{ Hz}}{f} \right]^2 \text{ Hz}^{-1/2}, \Delta\alpha = 0.46^\circ$$

$$D = 1.5 \text{ m} : h(f) = 5.96 \cdot 10^{-25} \left[ \frac{10 \text{ Hz}}{f} \right]^2 \text{ Hz}^{-1/2}, \Delta\alpha = 0.28^\circ$$

$$D = 2.2 \text{ m} : h(f) = 8.53 \cdot 10^{-25} \left[ \frac{10 \text{ Hz}}{f} \right]^2 \text{ Hz}^{-1/2}, \Delta\alpha = 0.19^\circ$$

$$D = 5 \text{ m} : h(f) = 1.75 \cdot 10^{-24} \left[ \frac{10 \text{ Hz}}{f} \right]^2 \text{ Hz}^{-1/2}, \Delta\alpha = 0.09^\circ$$

$$D = 3 \text{ km} : h(f) = 4.17 \cdot 10^{-24} \left[ \frac{10 \text{ Hz}}{f} \right]^2 \text{ Hz}^{-1/2}, \Delta\alpha = 0.04^\circ$$

Reflexion of scattered light by the inner edge of a baffle  
or by a reflecting element under grazing incidence

