



# Grid-tune of QCDNUM and SPLINT

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## New SPLINT routines

- Added a few routines to manage spline objects in memory, see section 7 write-up

<code>nw = isp_SpSize(ia)</code>	Get object size
<code>call ssp_Erase(ia)</code>	Clear memory
<code>call ssp_SpDump(ia, 'filename')</code>	Dump spline to disk
<code>ia = isp_SpRead('filename')</code>	Read spline from disk
<code>call ssp_SpSetVal(ia, i, dd)</code>	Write info into a spline
<code>dd = dsp_SpGetVal(ia, i)</code>	Read info from a spline

- Can write splines to disk and read them back in another run to use as a reference
- Can store extra info in such splines like evolution parameters, particle code, *etc.*
- New version QCDNUM-17-01-82 at [www.nikhef.nl/user/h24/download](http://www.nikhef.nl/user/h24/download)

# Structure function splines

- Fill 2-dim 100×50 spline with NNLO  $F_2$
- With a  $\sqrt{s}$  cut of 300 GeV this takes 4.9 CPU secs (stf computed with `ZMSTFUN`)
- Added a new structure function grid-point routine to `ZMSTF`

```
F = ZMSTFIJ( istf, def, ix, iq, ichk )
```

- It now takes 1.4 CPU secs → grid routine is factor 3 faster
- Still quite slow: introduce fast structure function spline filling routine in `SPLINT` that makes use of the list processing capability of `ZMSTFUN`

```
call ssp_S2F123( ia, iset, def, istf, rs )
```

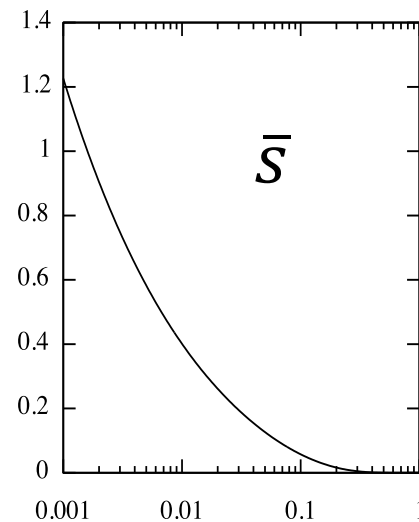
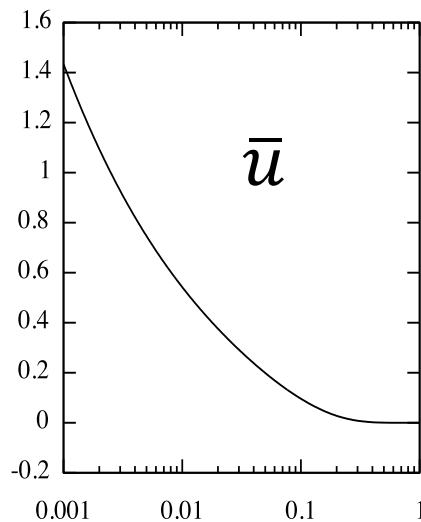
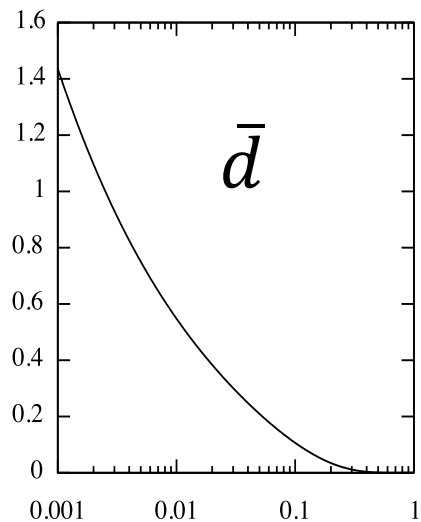
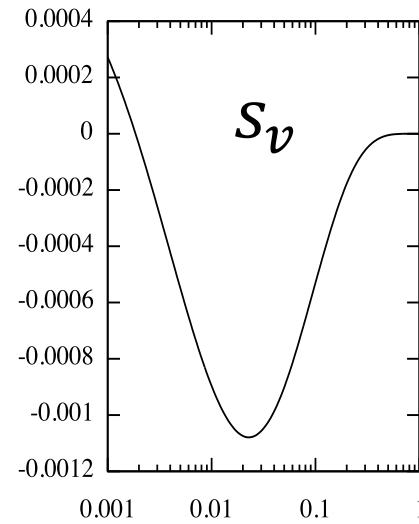
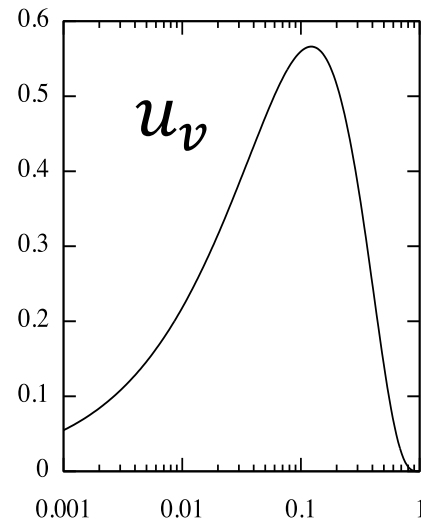
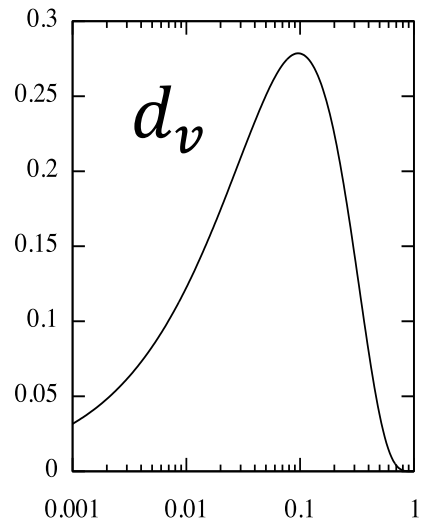
- Filling now takes 0.013 CPU secs: factor 100 faster!

# Timing study of `zmstf`: CPU per structure function

- Random set of grid points or  $x-Q^2$  points
- Compute  $F_2$  in loop or as a list
- `ZMSTFIJ` factor 3 faster than `ZMSTFUN`
- `ZMSTFUN` little difference in on/off-grid loop
- `ZMSTFUN` list is factor 10-100 faster than loop
- `SPLINE` faster than `ZMSTFIJ` for  $n > 10$
- Faster than `ZMSTFUN` list for  $n > 100$
- Factor 4 faster than `ZMSTFUN` for large  $n$

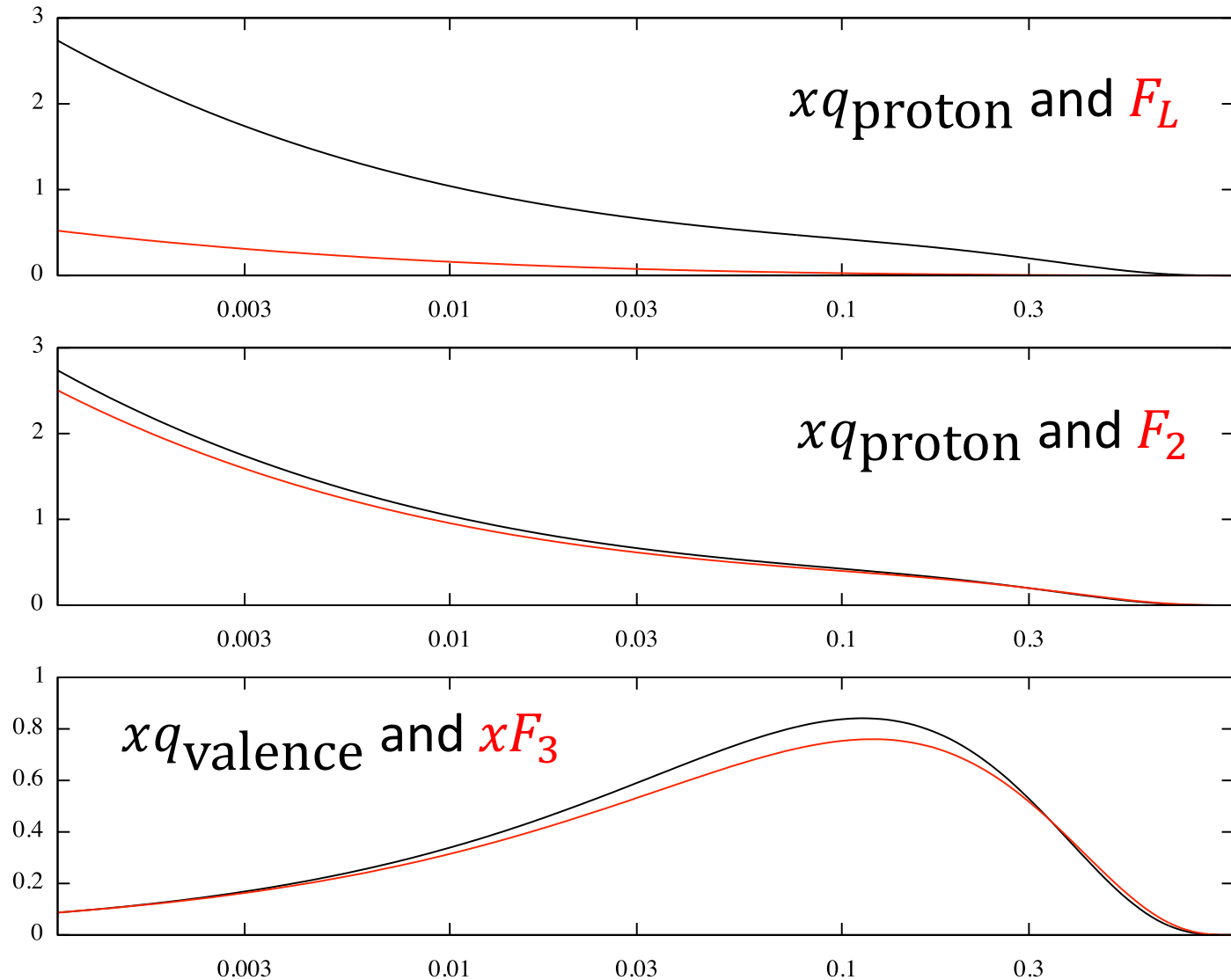
Routine	On/off-grid	Loop/List	$n$	$t/n$ [ms]
ZMSTFIJ	on-grid	loop	10	0.117
ZMSTFUN	on-grid	loop	10	0.417
ZMSTFUN	off-grid	loop	10	0.437
ZMSTFUN	on-grid	list	10	0.050
ZMSTFUN	on-grid	list	100	0.009
ZMSTFUN	on-grid	list	1000	0.003
ZMSTFUN	off-grid	list	10	0.065
ZMSTFUN	off-grid	list	100	0.014
ZMSTFUN	off-grid	list	1000	0.004
SPLINE	off-grid	loop	10	0.108
SPLINE	off-grid	loop	100	0.014
SPLINE	off-grid	loop	1000	0.001

# BATUNE00: Evolve test-job pdfs in VFNS NNLO from 1-100 GeV<sup>2</sup>



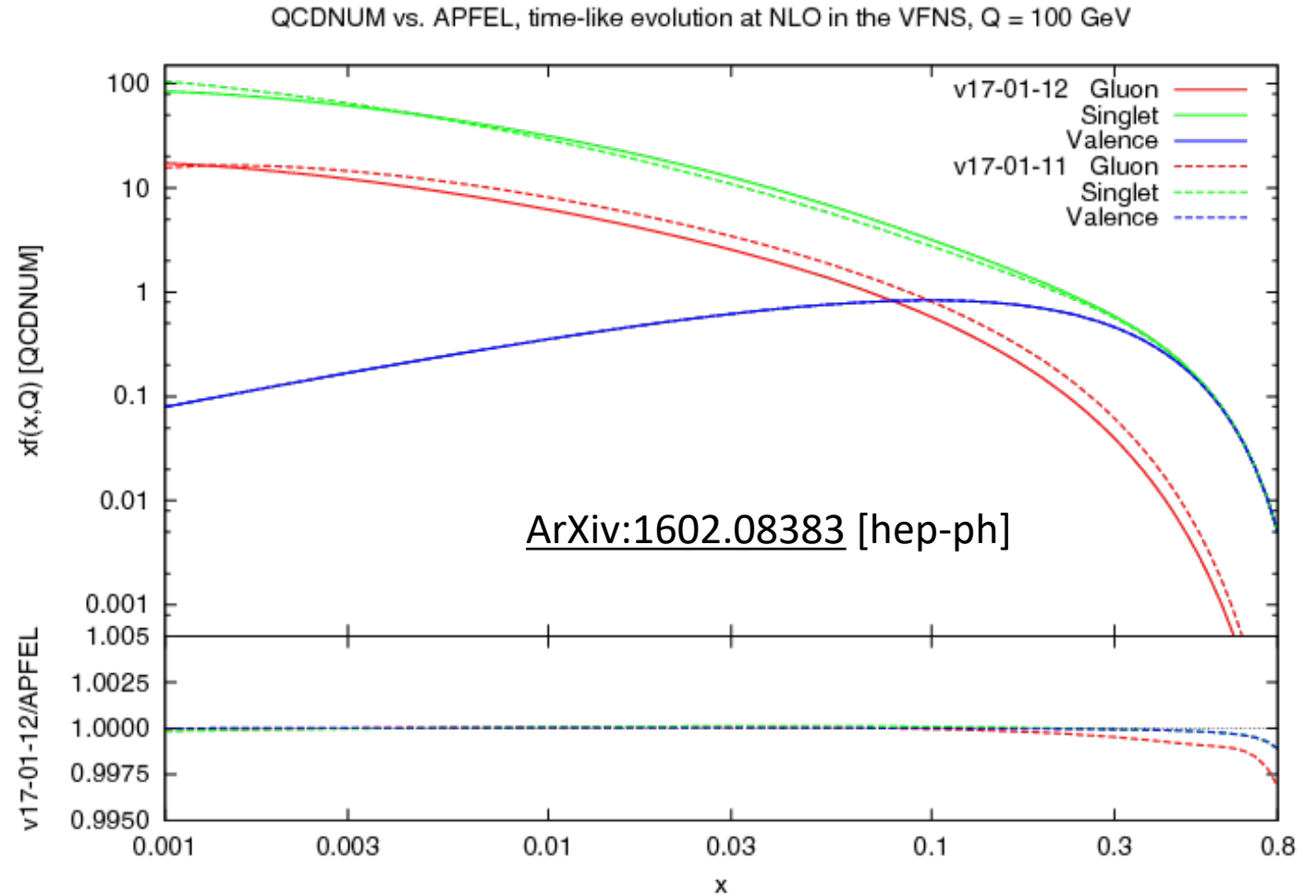
- $u, d, s$  evolved to 100 GeV<sup>2</sup>
- $c, b$  not shown
- $s, c, b$  NNLO evolution generates a small valence component with  $\bar{q} > q$  at intermediate  $x$
- Starting pdfs for BATUNE01

# Pdfs and structure functions versus $x$ at $\mu^2 = 100 \text{ GeV}^2$



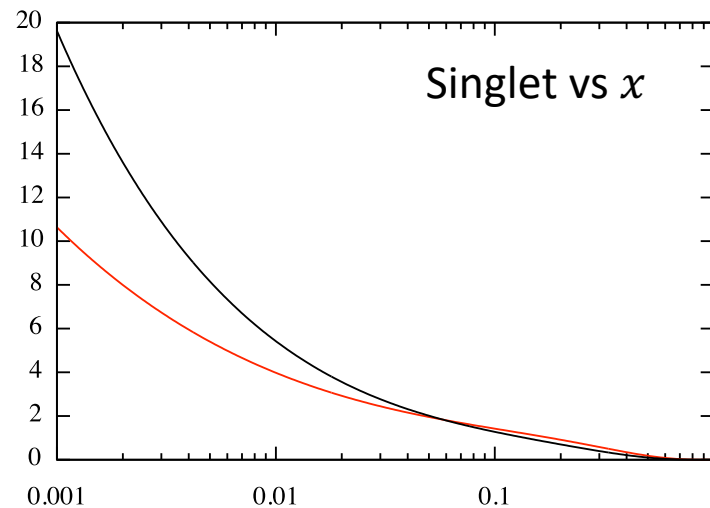
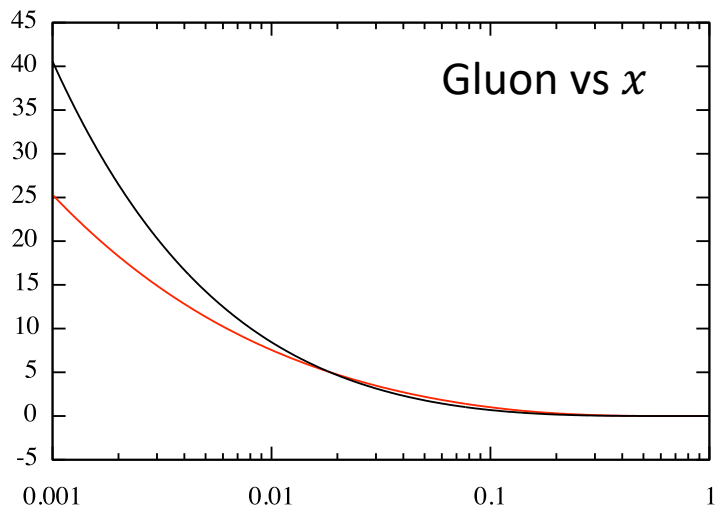
# BATUNE01: write high density reference splines

- Plot shows QCDNUM/APFEL comparison with pdfs evolved on a 200 point QCDNUM  $x$  grid
- We assume that pdfs evolved on a 300x150 grid are accurate enough for benchmarking
- Write high resolution (100x50 nodes) 2-dim reference splines of
  - Gluon
  - Singlet
  - $F_2$  proton
  - $F_L$  proton
  - $xF_3$  valence



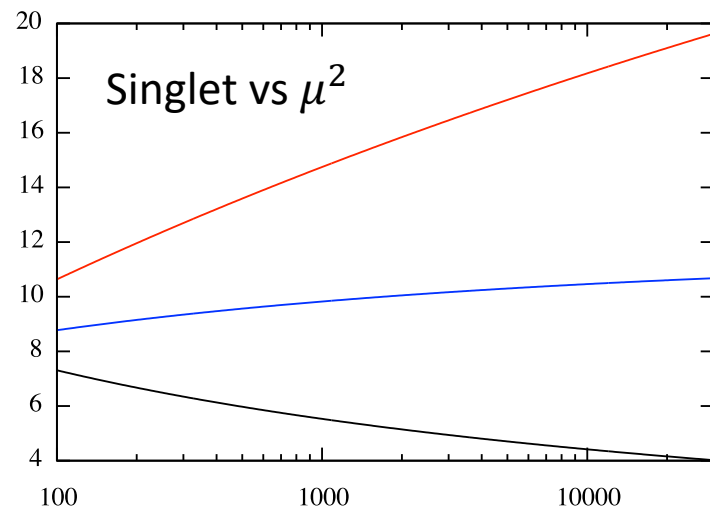
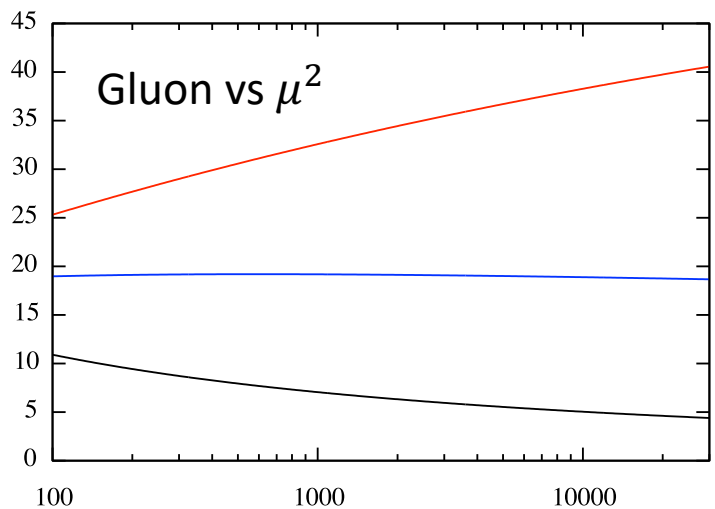
# BATUNE01: Evolve from 100-30000 GeV<sup>2</sup> at NNLO

$\mu^2 = 100 \text{ GeV}^2$   
 $\mu^2 = 3 \times 10^4 \text{ GeV}^2$



$\mu^2 = 100 \text{ GeV}^2$   
 $\mu^2 = 3 \times 10^4 \text{ GeV}^2$

$x = 10^{-3}$   
 $x = 0.02 (\times 4)$   
 $x = 0.5 (\times 1000)$



$x = 10^{-3}$   
 $x = 0.02 (\times 3)$   
 $x = 0.5 (\times 40)$

Almost pure logarithmic  $\mu^2$  dependence  $\rightarrow$  do not need many grid/node points

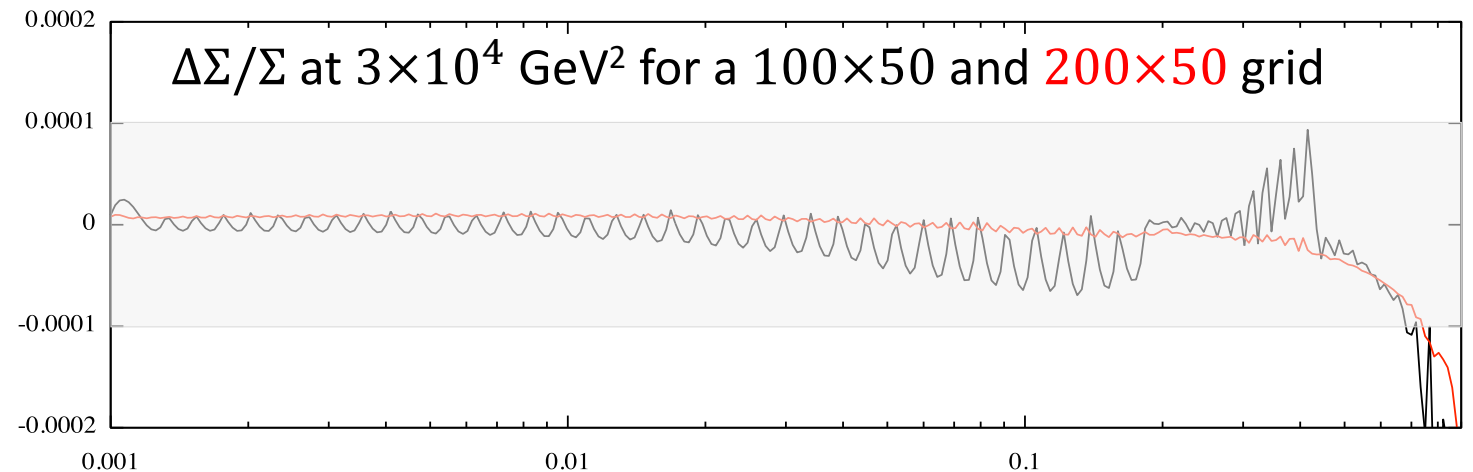
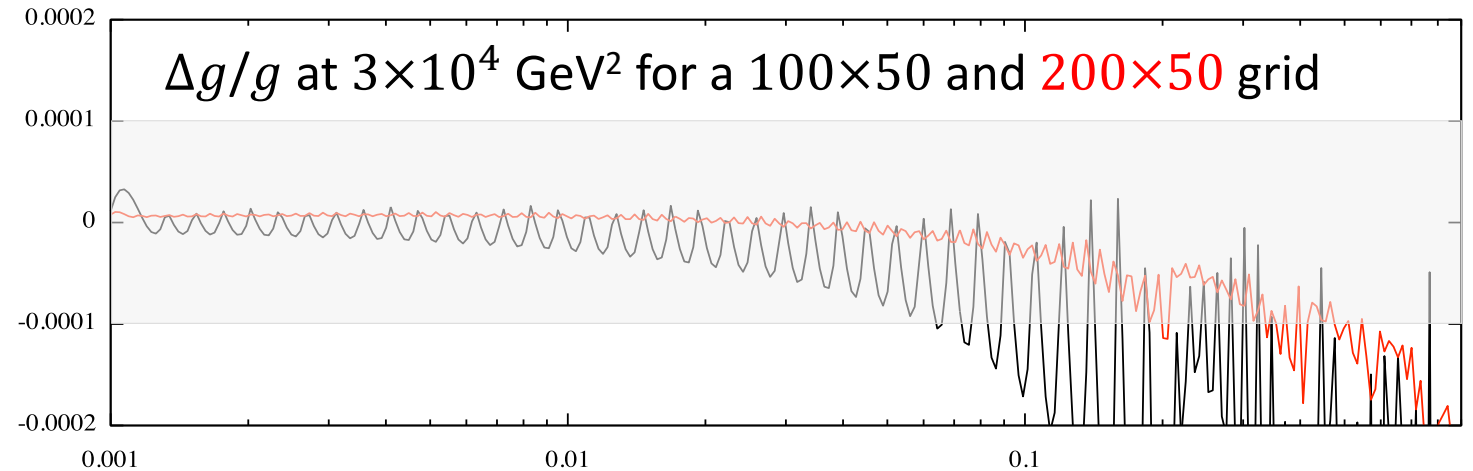


# BATUNE02: Tune $x$ - $\mu^2$ grid with respect to reference splines

## Subgrids in $x$

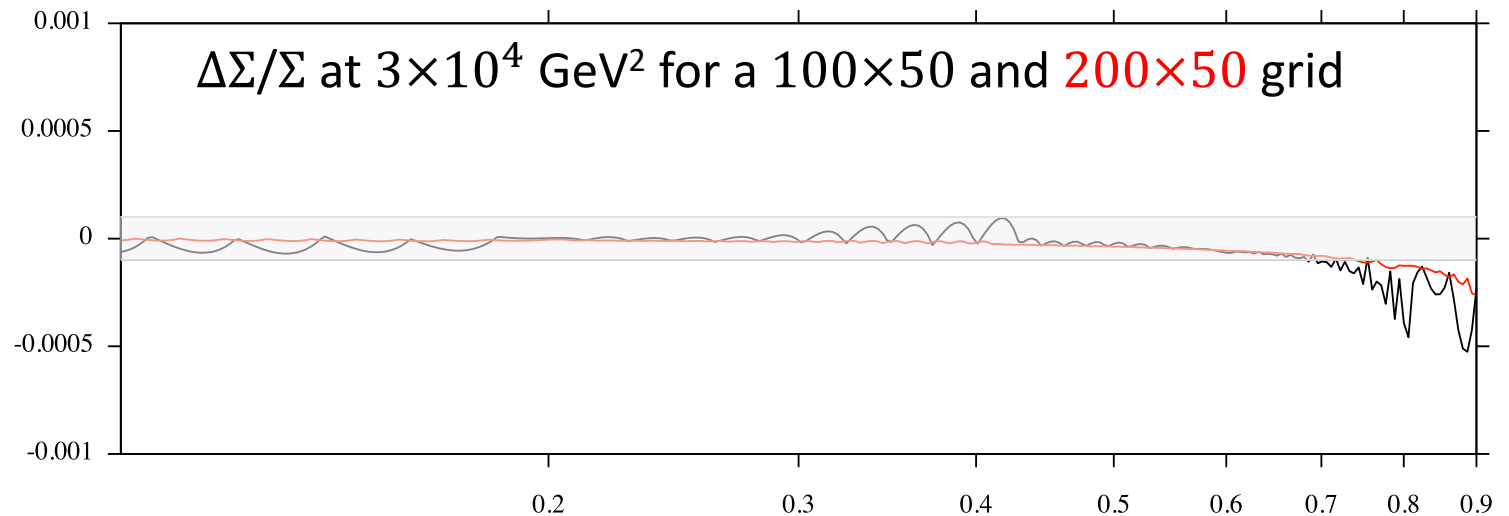
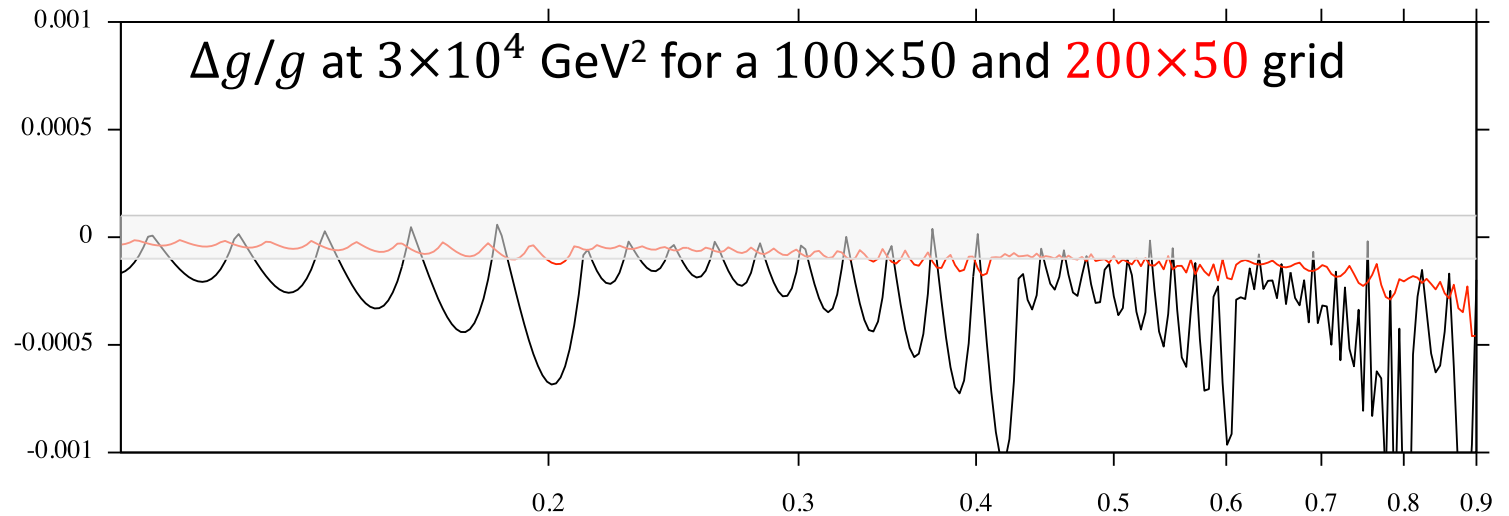
Lower Limit	Weight
$10^{-3}$	1
0.2	2
0.4	4
0.6	8
0.8	16

- 200x50 grid evolve in 12.8 ms
- 100x50 grid evolve in 5.1 ms
- $\Delta f/f < 10^{-4}$  for  $x < 0.1$
- Bit larger deviation at large  $x$  where pdfs are very small



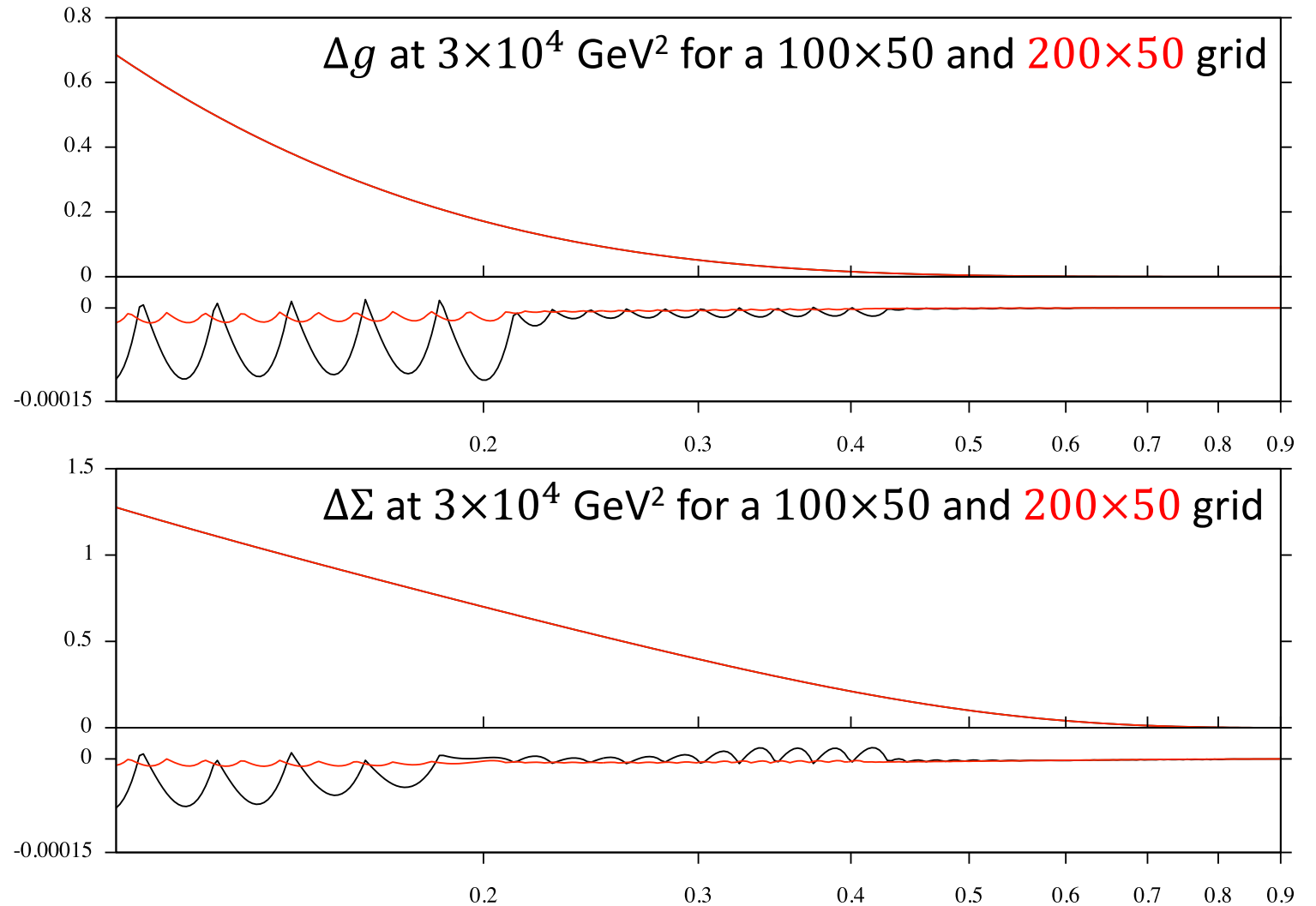
# BATUNE02: $\Delta g/g$ and $\Delta\Sigma/\Sigma$ at large $x$

- 200x50 evolution accurate to  $10^{-4}$  up to large  $x$
- Also true for 100x50 singlet
- 100x50 gluon less accurate but still within  $10^{-3}$
- No need for high accuracy since gluon vanishes quickly at large  $x$



# BATUNE02: Absolute deviation $\Delta g$ and $\Delta s$ at large $x$

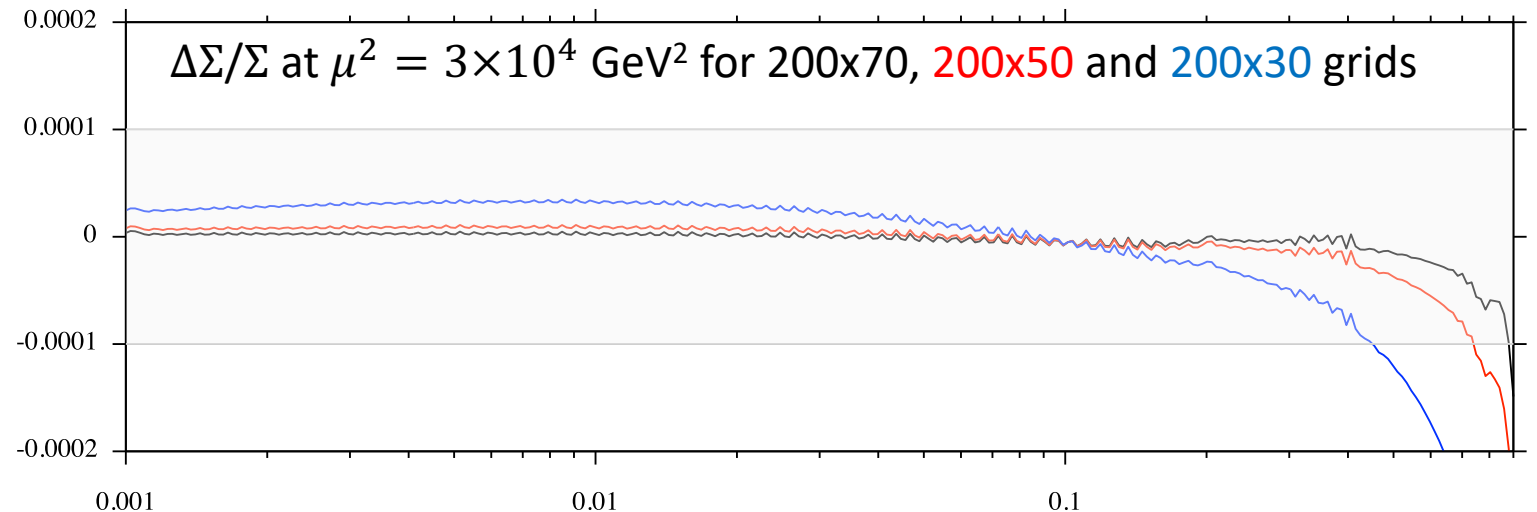
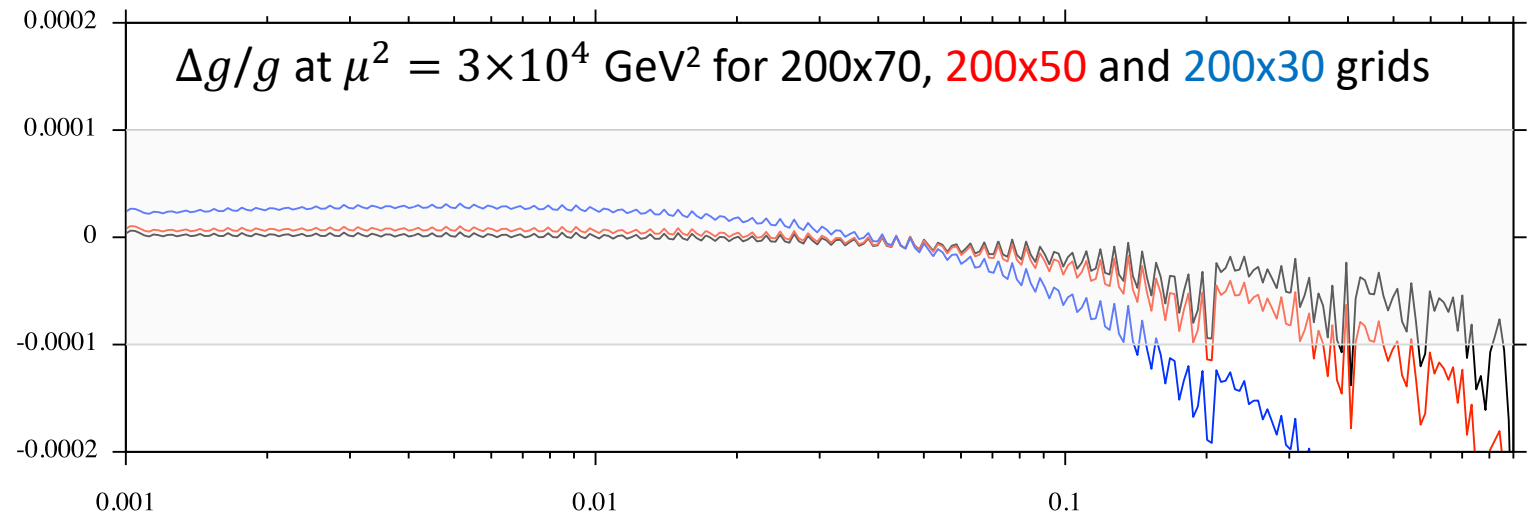
- Gluon vanishes faster than singlet at large  $x$
- Tuning for large relative gluon accuracy is not useful here
- 100x50 is my preferred grid



# BATUNE02: Vary the number of $\mu^2$ grid points

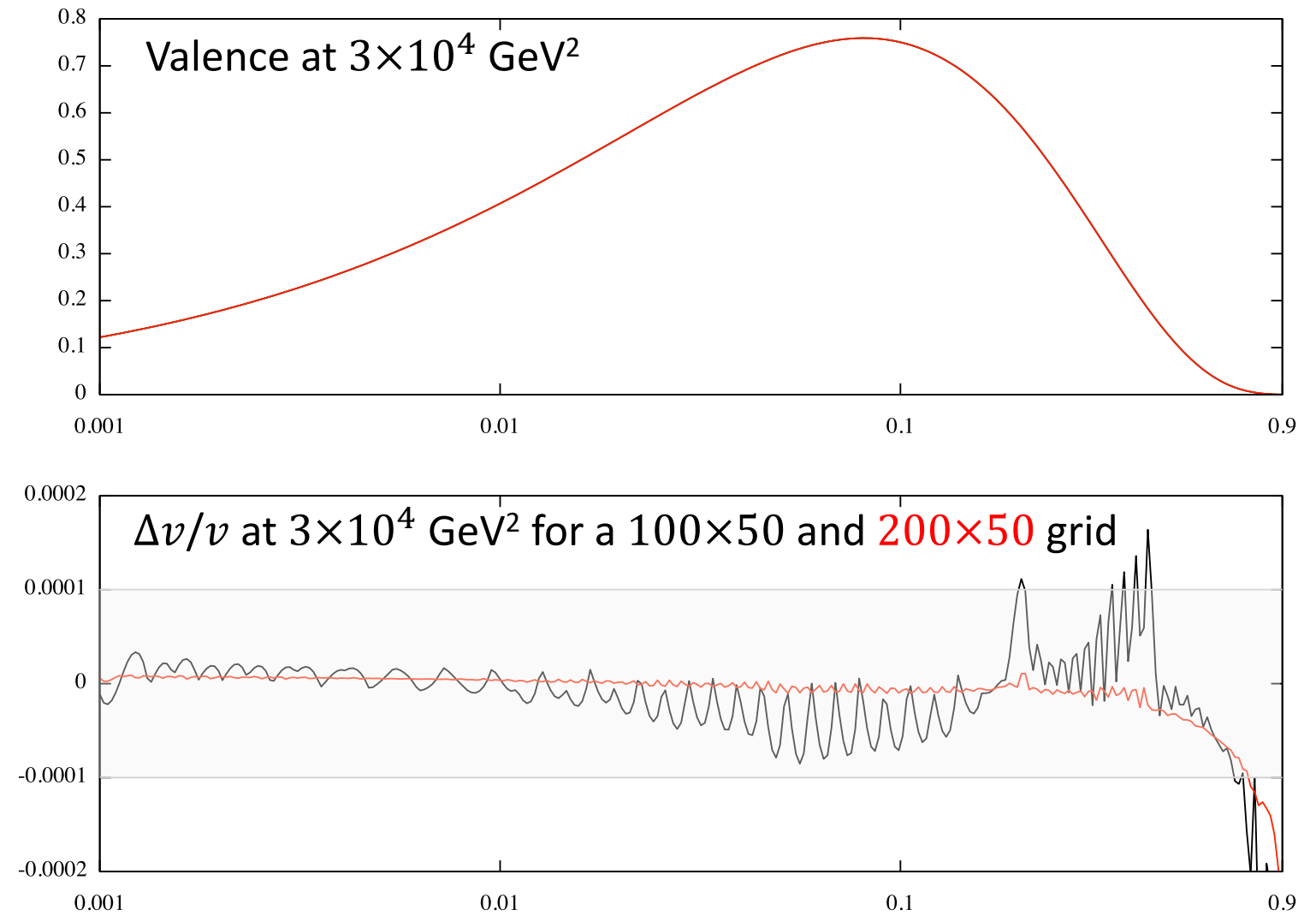
- Take  $n_x = 200$  to have clear plots of  $\Delta f/f$
- Small bias develops for  $n_q < 50 \rightarrow$  I prefer  $n_q = 50$
- Here are the CPU times of evolution with various grid settings

$n_x$	$n_q$	$t$ [ms]	factor
300	150	56.8	11.1
200	90	22.3	4.4
200	70	17.6	3.5
200	50	12.8	2.5
200	30	8.1	1.6
100	50	5.1	1.0



# BATUNE02: Valence pdf and $\Delta v/v$

- Tuning also OK for the valence pdf



# QCDNUM grid tuning conclusion

## Sub-grids in $x$

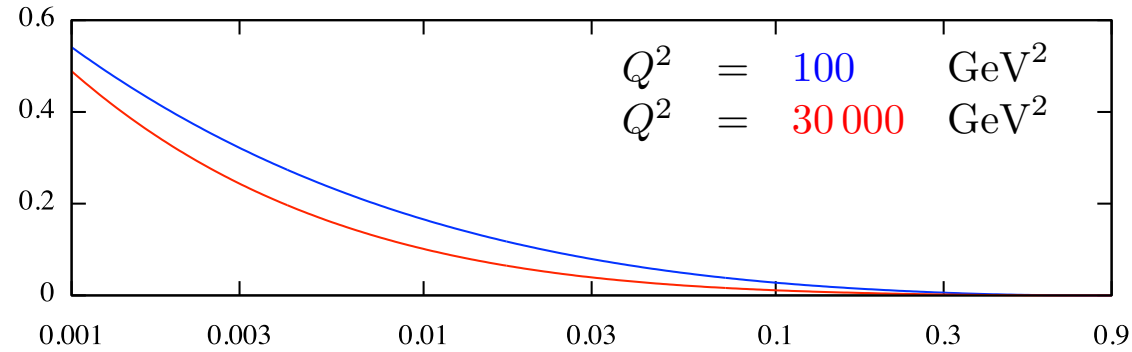
Lower Limit	Weight
$10^{-3}$	1
0.2	2
0.4	4
0.6	8
0.8	16

- My preferred grid is 100 points in  $x$  and 50 points in  $\mu^2$  with 5 sub-grids as given in the table
- NNLO evolution on this grid costs about 5 msec in CPU
- Numerical error  $\Delta f / f < 10^{-4}$  for  $x < 0.1$
- Bit larger relative deviation at large  $x$  where pdfs are very small

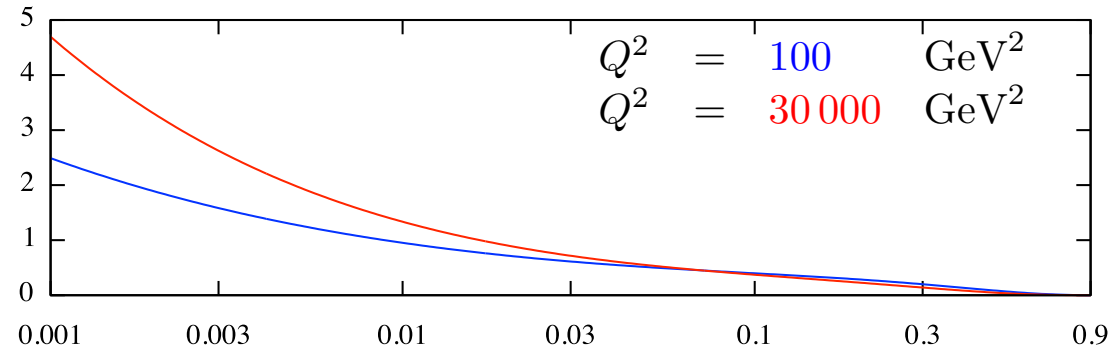
Increasing to  $n_x = 200$  gives  $\sim 10$  times better accuracy with double the cost in CPU  
Increasing  $n_q$  does not give much gain in precision

# BATUNE03: Read structure function reference splines

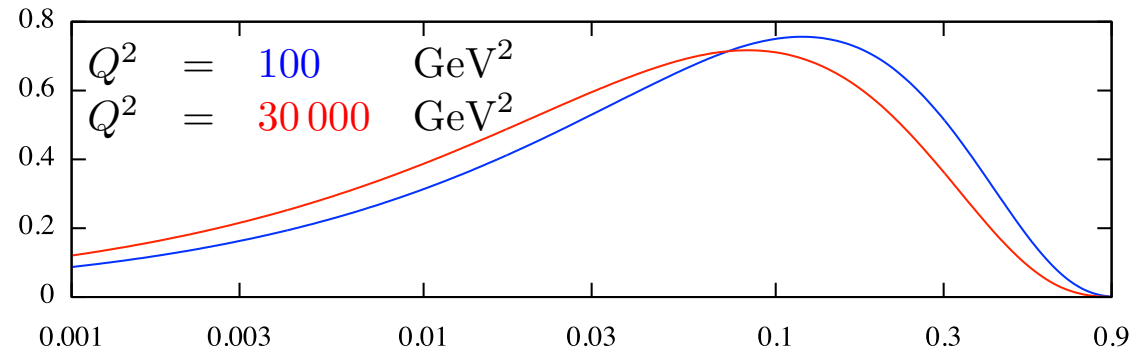
$F_L$  proton



$F_2$  proton

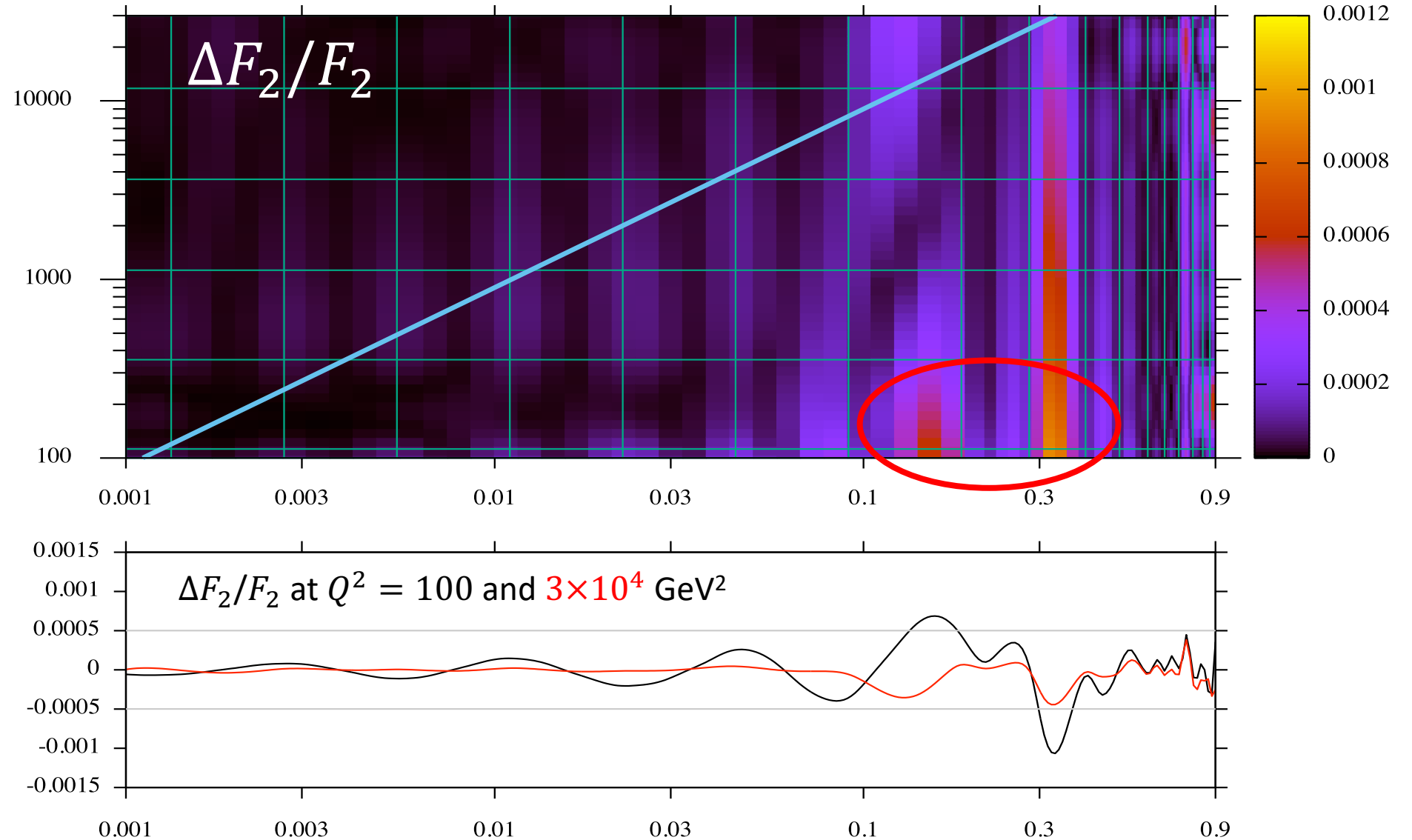


$\chi F_3$  valence



# BATUNE03: tune proton $F_2$ spline

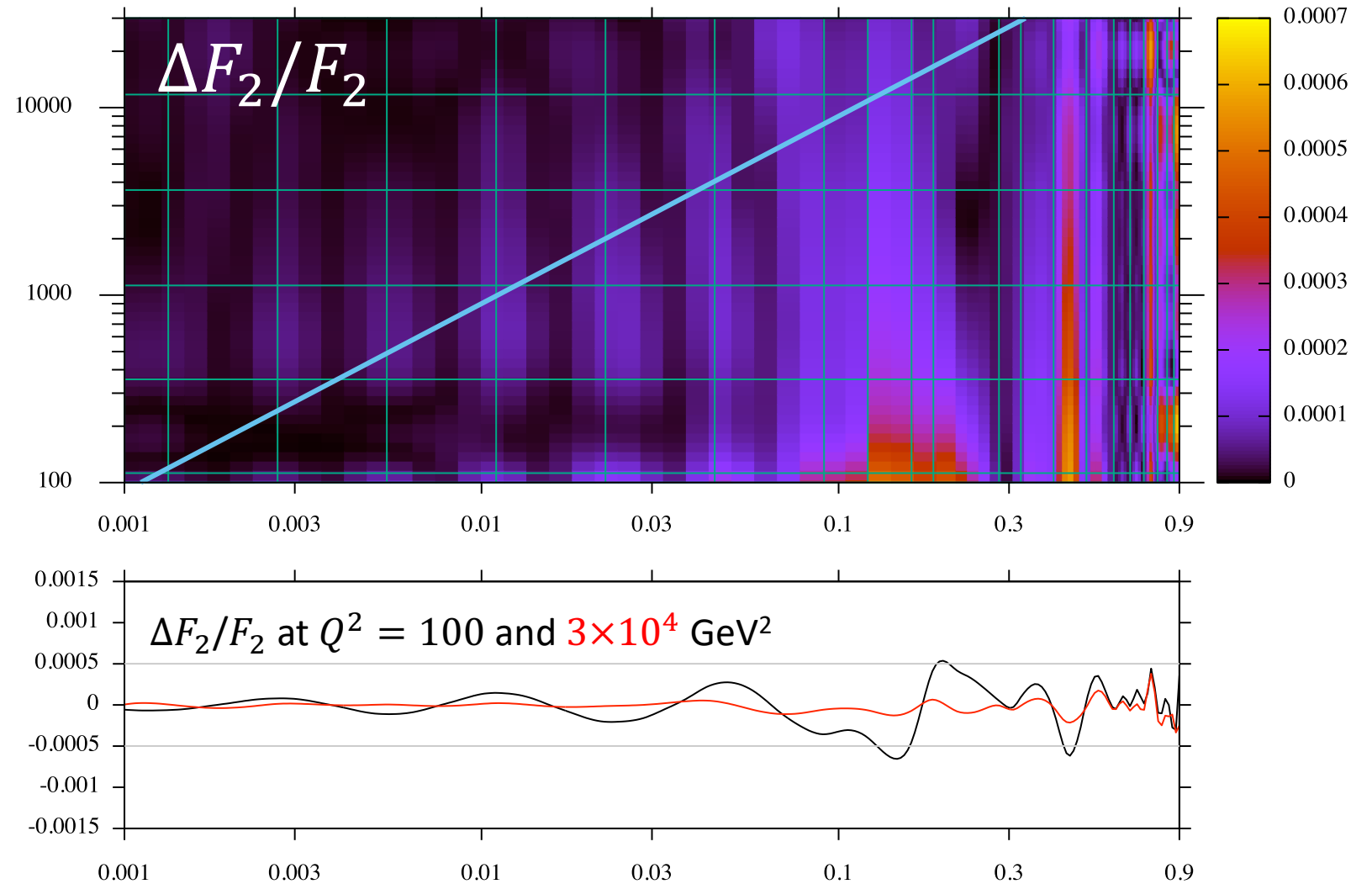
- Step-x = 5 and step-q = 10 gives 22×7 nodes
- Add 3 nodes to suppress hot spots at  $x \sim 0.2$



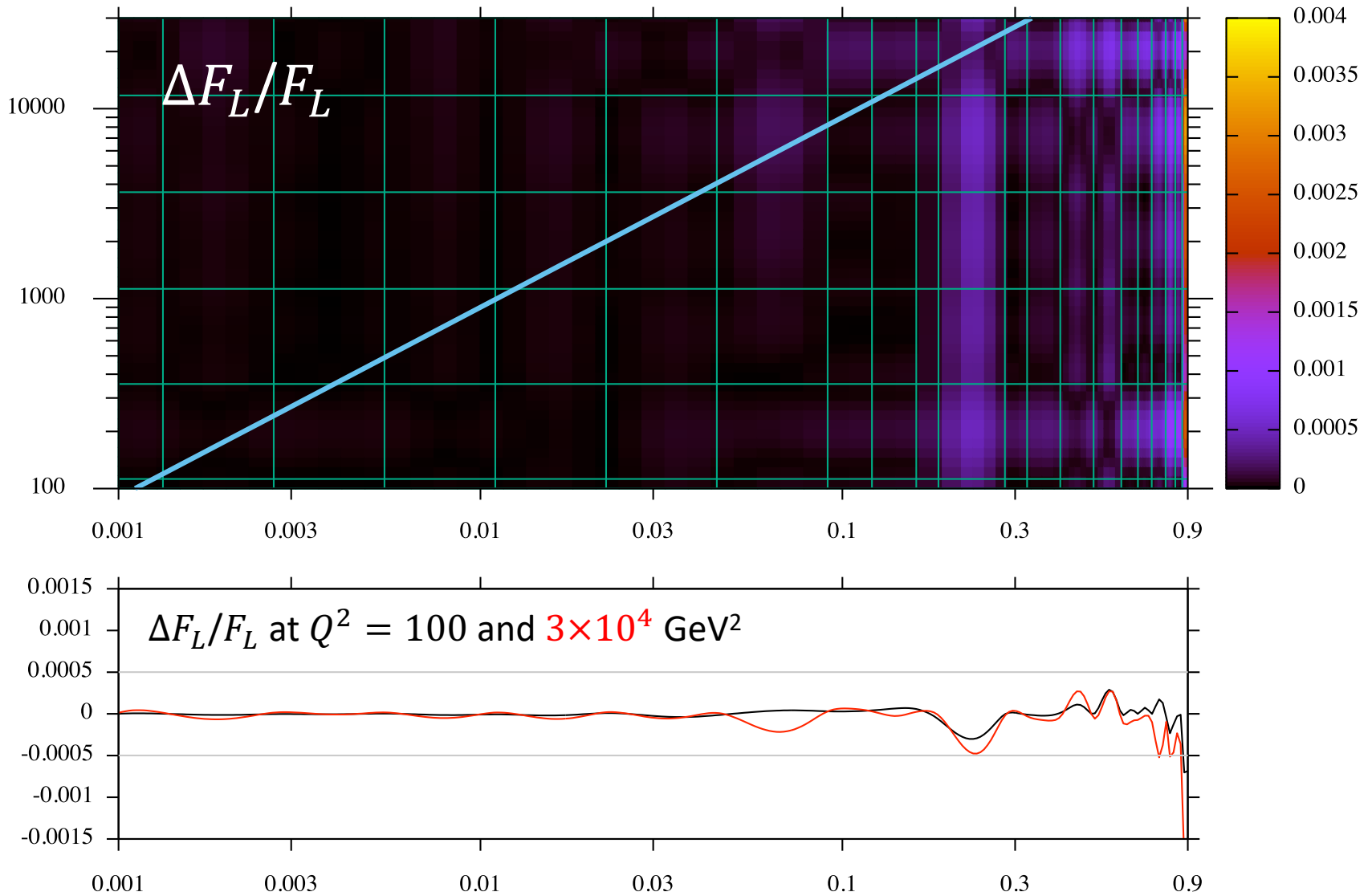


# BATUNE04: proton $\Delta F_2/F_2$ on a $25 \times 7$ node grid

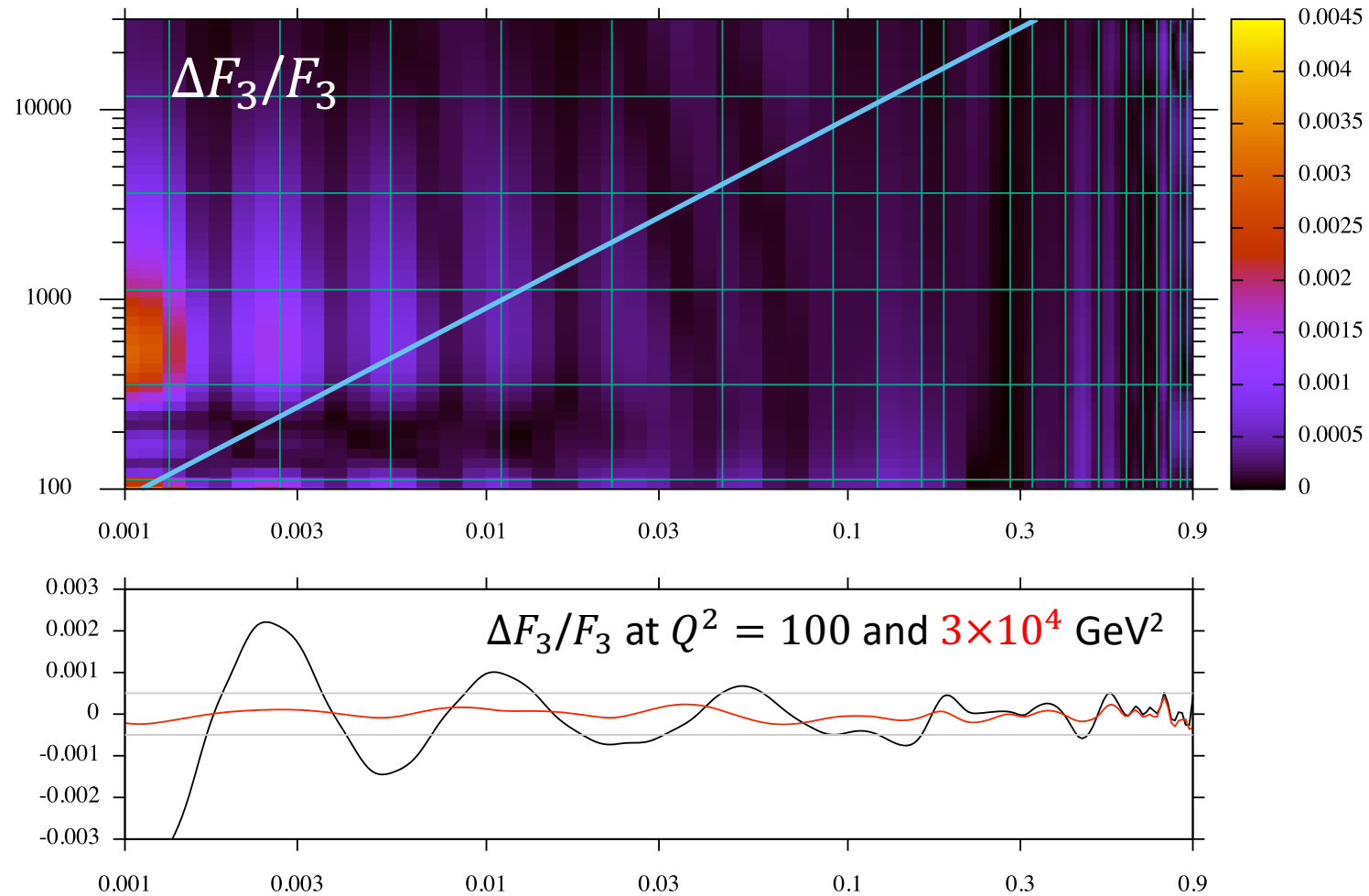
- Add nodes at  $x = 0.13, 0.15, 0.33$  to suppress some minor hot-spots
- Slight improvement
- $\Delta F_2/F_2 < 5 \times 10^{-4}$  almost everywhere



# Node-tuning also fine for proton $F_L$ ...



... and acceptable for valence  $x F_3$



NB: there is only a small contribution from  $x F_3$  to the cross-section at high  $Q^2$

## BATUNE04: Structure function splines for xsec (25×7 nodes)

- $F_L$  for uptype and downtype  $\Sigma(q + \bar{q})$
- $F_2$  for uptype and downtype  $\Sigma(q + \bar{q})$
- $x F_3$  for uptype and downtype  $\Sigma(q - \bar{q})$



3 msec for 6  
25×7 splines

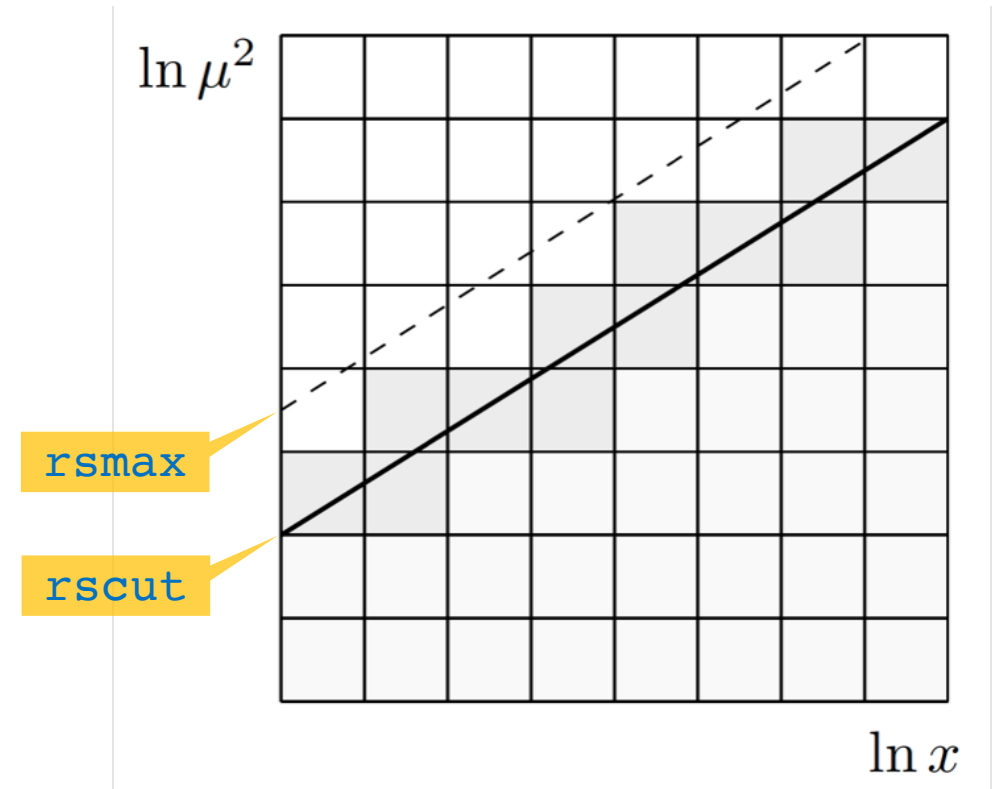
### Simplified cross-section

$$\begin{aligned} F_{2,L} &= \frac{1}{9} F_{\text{dtype}} + \frac{4}{9} F_{\text{utype}} & y &= Q^2/xs \\ x F_3 &= 0 & Y_{\pm} &= 1 \pm (1 - y)^2 \end{aligned}$$

$$\frac{d^2\sigma}{dx dQ^2} = \frac{2\pi\alpha^2}{xQ^4} [Y_+ F_2 - y^2 F_L + Y_- x F_3]$$

# $\sqrt{s}$ cut revisited

- In the transition region between `rscut` and `rsmax` the spline is not available everywhere
- This region should be avoided when reading a spline to fill another spline
- Cure: set the  $\sqrt{s}$  cut of the source spline above the `rsmax` of the target spline



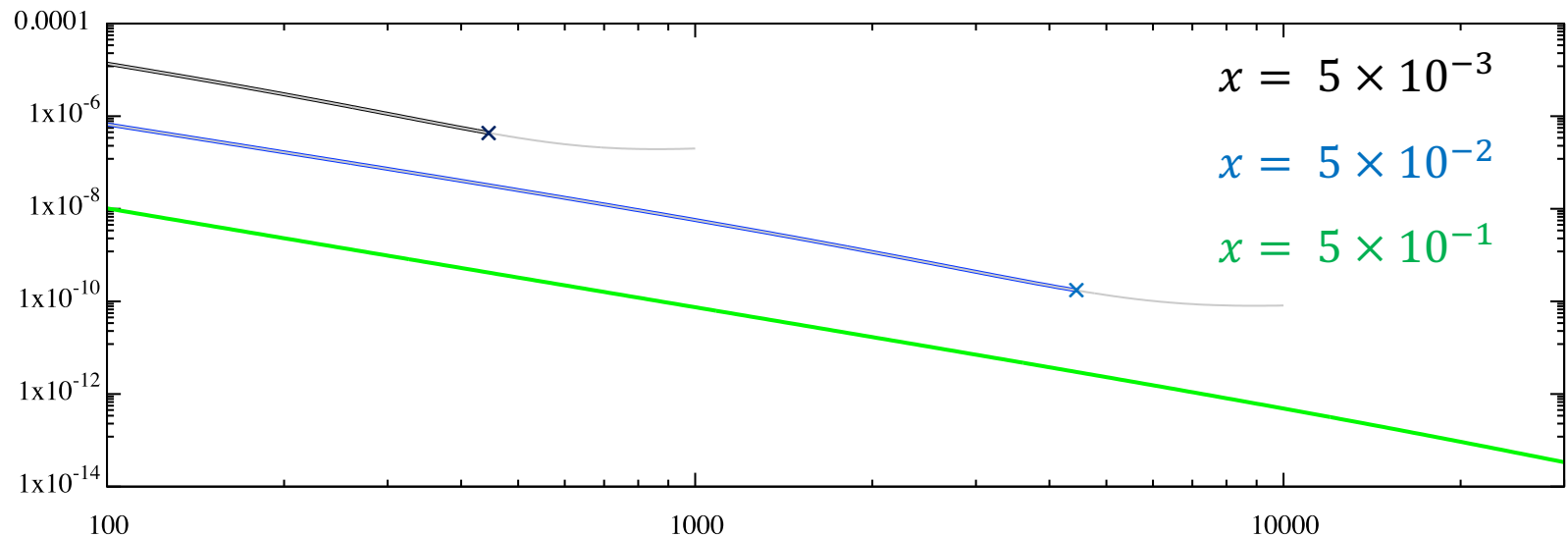
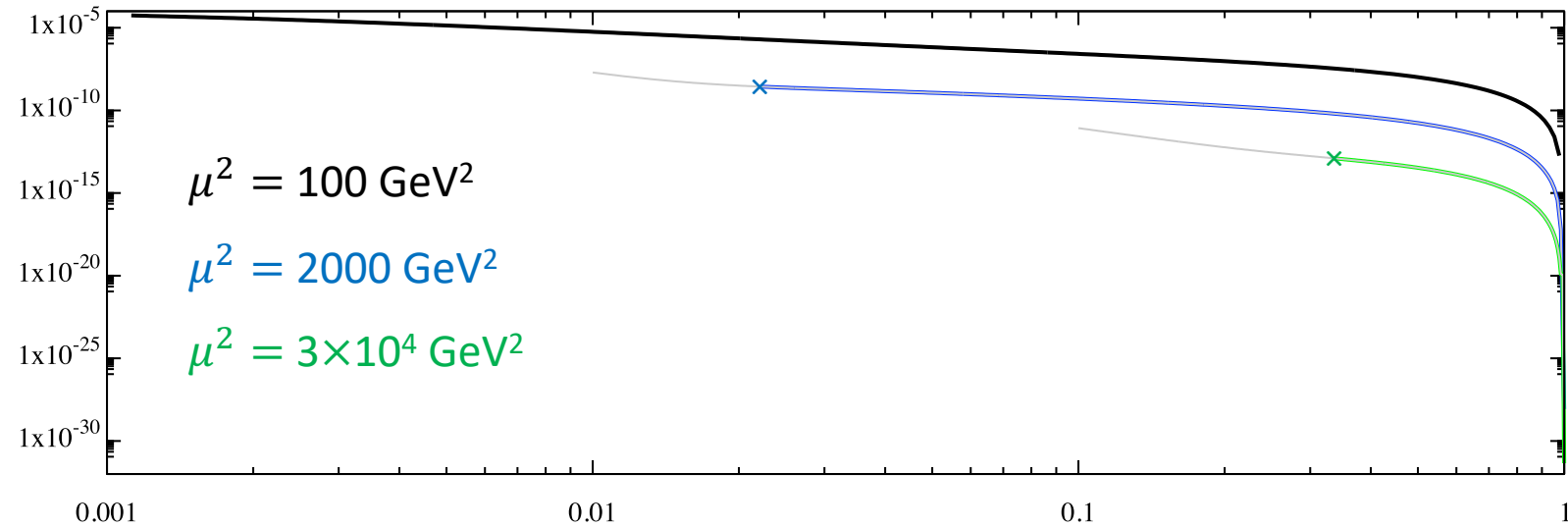
Add two new `SPLINT` routines

```
rsc = dsp_RsCut(ia)
```

```
rsm = dsp_RsMax(ia)
```

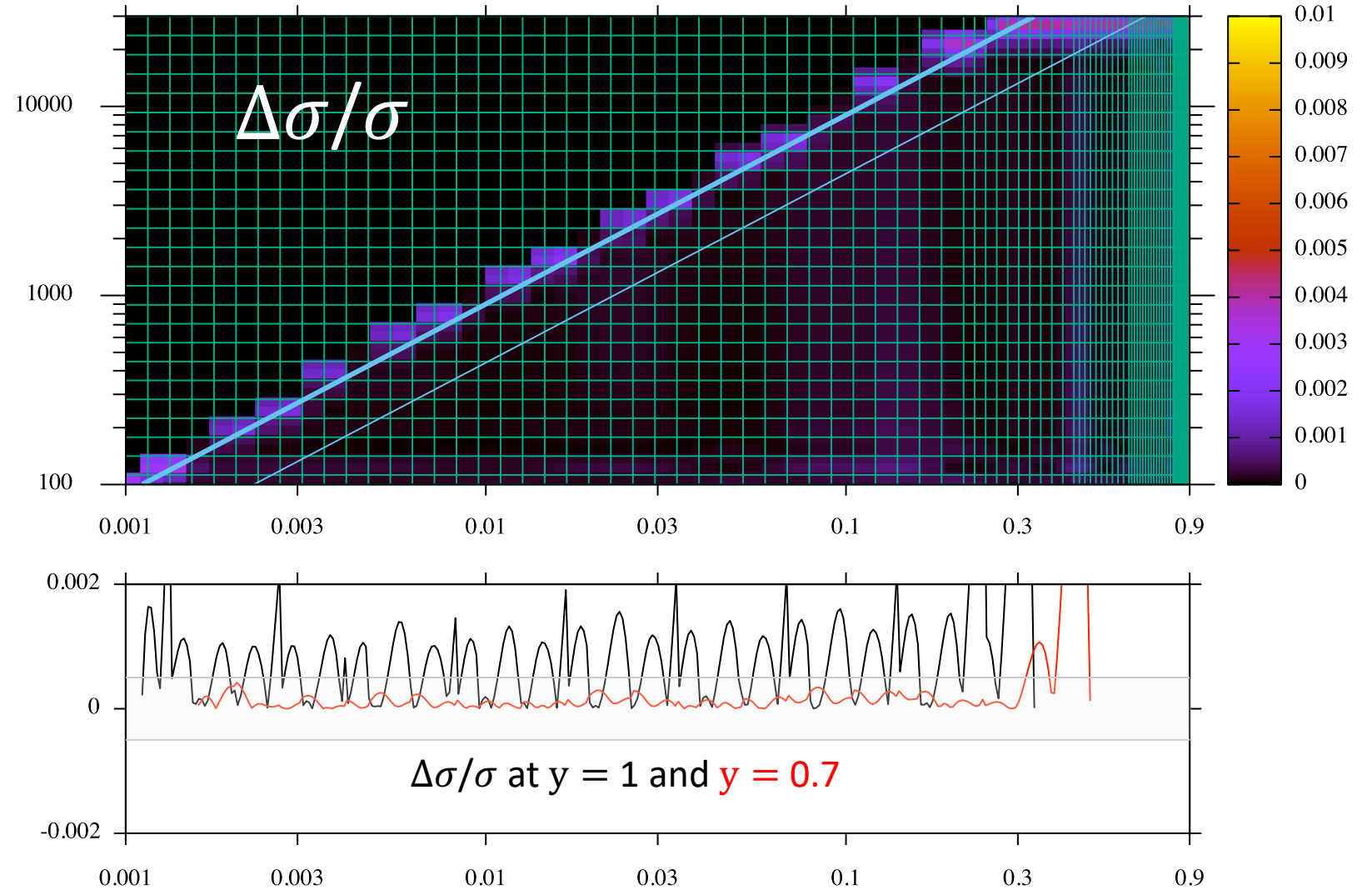
# BATUNE04: compute (simplified) cross-section

- Very steep fall-off over several orders of magnitude
- Such steep dependence can not be splined accurately on a coarse node-grid
- The grey curves show that the cross-section extrapolates smoothly beyond the kinematic cut  $\rightarrow$  no further action required
- Next: find optimal nodes for cross-section spline



# Cross-section spline with 100×25 nodes

- No  $\sqrt{s}$  cut at input structure function splines (not worth it)
- Set step-x = 1 and step-q = 2
- $\sqrt{s}$  cut at 300 GeV
- $\Delta\sigma/\sigma < 5 \times 10^{-4}$  except along the kinematic cut ( $y = 1$ )
- Since cross-section smoothly extrapolates we can raise the  $\sqrt{s}$  cut to improve at  $y = 1$

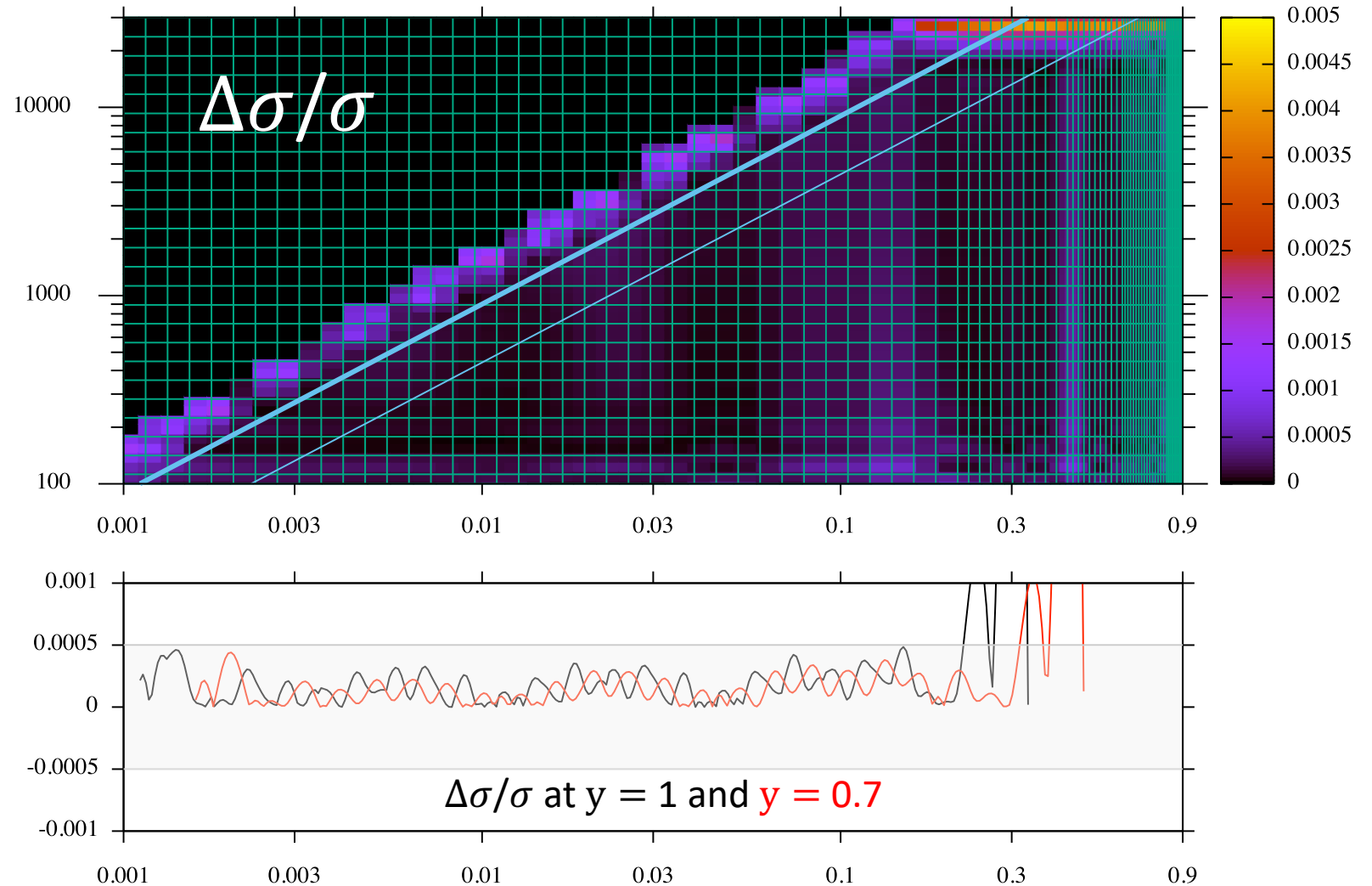


# Cross-section spline with 100×25 nodes

- Set step-x = 1 and step-q = 2
- $\sqrt{s}$  cut at 370 GeV
- $\Delta\sigma/\sigma \lesssim 5 \times 10^{-4}$  along the kinematic cut ( $y = 1$ )

## Timing (optimised code)

	$n_x$	$n_q$	$t$ [ms]
Evolution	100	50	3.6
6 Stf splines	22	7	2.9
	100	50	4.5
Xsec spline	100	25	2.2
	50	25	1.2





## Changes in the QCDNUM distribution

- Stand-alone write-up for `ZMSTF` (and `HQSTF` in the make)
- All write-ups are now collected in the `/doc` directory
- C++ test jobs are now in `/testjobs` (no more `/testjobsCxx`)
- Running jobs from the `/run` directory:
  - Must specify extension `.f` or `.cc`
  - Can specify source code directory (default `../testjobs`)

```
bash> ./runtest example.f
bash> ./runtest example.cc
bash> ./runtest mydir/myjob.cc
```

# BATUNE goodies in the QCDNUM-17-01-82 distribution

- **BAT** directory
  - `batune00 – batune04.f` tuning jobs
  - `batxxpyy.plt` selection of gnuplot plotting macros
  - `btxxyyyyy.dat` selection of gnuplot ASCII input files
- Can do straight-forward check of QCDNUM (without having to plot)

```
bash> cd qcdnum-17-01-82/run
bash> ./runtest ../BAT/batune00.f
bash> diff ../BAT/batune00.dat ../plots/batune00.dat
```

- The unix diff should give no differences
- My suggestion is to convert `batune00.f` to JULIA and do the same test to check for oscillations

## Tuning OK so what's next



- Timing of bin-integrations
- For this I need an ASCII file with bin limits  $x_1, x_2, \mu_1^2, \mu_2^2$
- Some streamlining of **SPLINT** code
- Bin integration at kinematic limit