

# INSTANTON INDUCED NEUTRINO MASSES IN 

RCFT ORIENTIFOLDS

## "OPEN DESCENDANT" PARTITION FUNCTIONS

等. Closed

$$
\frac{1}{2}\left[\sum_{i j} \chi_{i}(\tau) Z_{i j} \chi_{i}(\bar{\tau})+\sum_{i} K_{i} \chi_{i}(2 \tau)\right]
$$

Open

$$
\left.\frac{1}{2}\left[\sum_{[, a, n} N_{a} N_{b} A_{a b}^{i} \alpha_{i}\left(\frac{\tau}{2}\right)+\sum_{i, a} N_{a} M_{a}^{i} \hat{\chi}_{i} \frac{\tau}{2}+\frac{1}{2}\right]\right]
$$

$i$ : Primary field label (finite range)
$a$ : Boundary label (finite range)
$\chi_{i}$ : Character
$N_{a}$ : Chan-Paton (CP) Multiplicity

## ORIENTIFOLD PARTITION FUNCTIONS



## BOUNDARIES AND CROSSCAPS*

## 諩 Boundary coefficients

$$
R_{\left[a, \psi_{a}\right](m, J)}=\sqrt{\frac{|\mathcal{H}|}{\left|\mathcal{C}_{a}\right|\left|\mathcal{S}_{a}\right|}} \psi_{a}^{*}(J) S_{a m}^{J}
$$

粼 Crosscap coefficients

$$
U_{(m, J)}=\frac{1}{\sqrt{|\mathcal{H}|}} \sum_{L \in \mathcal{H}} e^{\pi i\left(h_{K}-h_{K L}\right)} \beta_{K}(L) P_{L K, m} \delta_{J, 0}
$$

*Huiszoon, Fuchs, Schellekens, Schweigert, Walcher (2000)

## COEFFICIENTS

## 蝪 Klein bottle

$$
K^{i}=\sum_{m, J, J^{\prime}} \frac{S^{i}{ }_{m} U_{(m, J)} g_{J, J^{\prime}}^{\Omega, m} U_{\left(m, J^{\prime}\right)}}{S_{0 m}}
$$

粼 Annulus

$$
A_{\left[a, \psi_{a}\right]\left[b, \psi_{b}\right]}^{i}=\sum_{m, J, J^{\prime}} \frac{S^{i}{ }_{m} R_{\left[a, \psi_{a}\right](m, J)} g_{J, J^{\prime}}^{\Omega, m} R_{\left[b, \psi_{b}\right]\left(m, J^{\prime}\right)}}{S_{0 m}}
$$

䗱 Moebius

$$
M_{\left[a, \psi_{a}\right]}^{i}=\sum_{m, J, J^{\prime}} \frac{P^{i}{ }_{m} R_{\left[a, \psi_{a}\right](m, J)} g_{J, J^{\prime}}^{\Omega, m} U_{\left(m, J^{\prime}\right)}}{S_{0 m}}
$$

$g_{J, J^{\prime}}^{\Omega, m}=\frac{S_{m 0}}{S_{m K}} \beta_{K}(J) \delta_{J^{\prime}, J^{c}}$

## TADPOLES \＆ANOMALIES

絆 Tadpole cancellation condition：

$$
\sum_{b} N_{b} R_{b(m, J)}=4 \eta_{m} U_{(m, J)}
$$

期 Cubic $\operatorname{Tr} F^{3}$ anomalies cancel

暽 Remaining anomalies by Green－Schwarz mechanism

溸 In rare cases，additional conditions for global anomaly cancellation＊

## FORMALISM CAN BE APPLIED TO：

＂Gepner Models＂
（minimal $N=2$ tensor products）
糔 Free fermions（4n real $+(9-2 \mathrm{n})$ complex）

紫 Kazama－Suzuki models
（requires exact spectrum computation）

鞠 Permutation orbifolds

## GEPNER ORIENTIFOLDS

C. Angelantonj, M. Bianchi, G. Pradisi, A. Sagnotti and Y. S. Stanev, Phys. Lett. B 387 (1996) 743 [arXiv:hep-th/9607229].
R. Blumenhagen and A. Wisskirchen, Phys. Lett. B 438, 52 (1998) [arXiv:hep-th/9806131].
G. Aldazabal, E. C. Andres, M. Leston and C. Nunez, JHEP 0309, 067 (2003) [arXiv:hep-th/0307183].
I. Brunner, K. Hori, K. Hosomichi and J. Walcher, arXiv:hep-th/0401137.
R. Blumenhagen and T. Weigand, JHEP 0402 (2004) 041 [arXiv:hep-th/0401148].
G. Aldazabal, E. C. Andres and J. E. Juknevich, JHEP 0405, 054 (2004) [arXiv:hep-th/0403262].

## THE SM SPECTRUM

## Current experimental information:

## 3 chiral families + vector-like states

Possible vector-like states:

Higgs?
right-handed neutrinos?
squarks, sleptons?
gluinos?
who knows what else?
(Some constraints from unification, if you believe it)

## MODELS

3 families

+ anything vector-like



## DE－CONFUSION

蝶 Space－time susy imposed（not necessary）．
锰 No moduli stabilization．

溸 Boundary state $\approx$ brane
粼 Complete set of CFT boundary states （in the sense of Sagnotti，Pradisi，Stanev）

曗 But：not the complete set of geometric branes．

## DATA

|  | $2004-2005^{*}$ | $2005-2006^{\dagger}$ |
| :---: | :---: | :---: |
| Trigger | "Madrid" | All 3 family models |
| Chiral types | 19 | 19345 |
| Tadpole-free(per type) | 18 | 1900 |
| Total configs | $45 \times 10^{6}$ | $145 \times 10^{6}$ |
| Tadpole free, distinct | 210.000 | 1900 |
| Max. primaries | $\infty$ | 1750 |

[^0]
## SU(5)



Note: gauge group is just $\operatorname{SU}(5)$ !

## A curiosity



## Truly hidden

 hidden sector
## GOALS

Finding the standard model?
Only a small part of the orientifold landscape ( 880 out of 30000 hodge numbers....)
Only rational points in moduli space.
No chance unless SM is extremely abundant.

## CENTRAL QUESTION

"Is the standard model a plausible solution to landscape and anthropic constraints?"

Too hard, even for string theorists... but some simpler sub-questions may be within reach

Q Does it exist in the landscape?
Q Which BSM versions can be realized?
Q Generic features?
Q Correlations?

- Are some SM features extremely rare without a potential anthropic explanation?


## NEUTRINO MASSES＊

镤 In field theory：easy；several solutions．
数 In string theory：non－trivial． （String theory is much more falsifiable！）．

政 Potentially anthropic．
（＊）Ibañez，Schellekens，Uranga，arXiv：0704．1079，JHEP（to appear）

The following ingredients cannot be taken for granted in String Theory：

糍 Existence of a Weinberg operator．

$$
\mathcal{L}_{W}=\frac{\lambda}{M}(L \bar{H} L \bar{H})
$$

粈 Existence of right－handed neutrinos．
䗲 Existence of non－zero Dirac masses．
維 Absence of massless B－L vector bosons．
教 Existence of Majorana masses．

## RIGHT-HANDED NEUTRINOS



|  | Total number <br> in 2005/2006 database |
| :--- | :---: |
| 0 | 16766656 |
| 1 | 475928 |
| 2 | 502820 |
| 3 | 62149717 |
| 4 | 686961 |

## The Madrid Model



Chiral $\operatorname{SU}(3) \times \operatorname{SU}(2) \times \mathrm{U}(1)$ spectrum:

$$
3(u, d)_{L}+3 u_{L}^{c}+3 d_{L}^{c}+3\left(e^{-}, \nu\right)_{L}+3 e_{L}^{+}
$$

Y massless

$$
Y=\frac{1}{6} Q_{a}-\frac{1}{2} Q_{c}-\frac{1}{2} Q d
$$

## MADRID MODELS

All these models have three right-handed neutrinos (required for cubic anomaly cancellation)

In most of these models:
B-L survives as an exact gauge symmetry
Neutrino's can get Dirac masses, but not Majorana masses (both needed for see-saw mechanism).

In a very small** subset, B-L acquires a mass due to axion couplings.
(*) 391 out of 10000 models with $\mathrm{SU}(3) \times \operatorname{Sp}(2) \times \mathrm{U}(1) \times \mathrm{U}(1)$ (out of 211000 in total)

## Abelian Masses

Green-Schwarz mechanism


Axion-Vector boson vertex
-------MWW

Generates mass vector bosons of anomalous symmetries

$$
(e . g . B+L)
$$

But may also generate mass for non-anomalous ones

$$
(Y, B-L)
$$

## B-L VIOLATION

But even then, B-L still survives as a perturbative symmetry.
It may be broken to a discrete subgroup by instantons.
This possibility can be explored if the instanton is described by a RCFT brane M.
B-L violation manifests itself as:

$$
\begin{aligned}
& \quad I_{M \mathbf{a}}-I_{M \mathbf{a}^{\prime}}-I_{M \mathbf{d}}+I_{M \mathbf{d}^{\prime}} \neq 0 \\
& I_{M a}=\text { chiral }\left[\#\left(V, V^{* *}\right)-\#\left(V^{*}, V\right)\right] \text { between branes } M \text { and } a \\
& a^{\prime}=\text { boundary conjugate of a }
\end{aligned}
$$

## B-L ANOMALIES

$$
I_{M \mathbf{a}}-I_{M \mathbf{a}^{\prime}}-I_{M \mathbf{d}}+I_{M \mathbf{d}^{\prime}} \neq 0
$$

Implies a cubic $\mathrm{B}-\mathrm{L}$ anomaly if M is a "matter" brane (Chan-Paton multiplicity $\neq 0$ ).
$\Rightarrow$ M cannot be a matter brane:
non-gauge-theory instanton
(stringy instanton, exotic instanton)

Implies a (B-L) $\left(\mathrm{G}_{\mathrm{M}}\right)^{2}$ anomaly even if we cancel the cubic anomaly

$$
\Rightarrow B \text {-L must be massive }
$$

(The converse is not true: there are massive B-L models without such instanton branes)

## Instantons:

See talks by Lüst, Plauschinn, Lerda, Morales, Blumenhagen, Cvetic, Kiritsis

## REQUIRED ZERO-MODES

Neutrino mass generation by non-gauge theory instantons*

The desired neutrino mass term $v^{c} v^{c}$ violates c and d brane charge by two units. To compensate this, we must have

$$
I_{M \mathbf{c}}=2 ; I_{M \mathbf{d}}=-2 \quad \text { or } \quad I_{M \mathbf{d}^{\prime}}=2 ; I_{M \mathbf{c}^{\prime}}=-2
$$

and all other intersections 0 .
( $\mathrm{d}^{\prime}$ is the boundary conjugate of d )
(*) Blumenhagen, Cvetic, Weigand, hep-tb/0609191 Ibañez, Uranga, hep-th/0609213

## NEUTRINO-ZERO MODE COUPLING

The following world-sheet disk is allowed by all symmetries


$$
L_{\text {cubic }} \propto d_{a}^{i j}\left(\alpha_{i} \nu^{a} \gamma_{j}\right), a=1,2,3
$$

## ZERO-MODE INTEGRALS

$$
\int d^{2} \alpha d^{2} \gamma e^{-d_{a}^{i j}\left(a_{i} \nu^{a} \gamma_{j}\right)}=\nu_{a} \nu_{b}\left(\epsilon_{i j} \epsilon_{k l} d_{a}^{i k} d_{b}^{j^{l}}\right)
$$

Additional zero modes yield additional fermionic integrals and hence no contribution

Therefore $\mathrm{I}_{\mathrm{Ma}}=\mathrm{I}_{\mathrm{Mb}}=\mathrm{I}_{\mathrm{Mx}}=0$ ( $\mathrm{x}=$ Hidden sector), and there should be no vector-like zero modes.

There should also be no instanton-instanton zero-modes except 2 required by susy.

## INSTANTON TYPES

In orientifold models we can have complex and real branes

| Matter brane M | Instanton brane M |
| :---: | :---: |
| $\mathrm{U}(\mathrm{N})$ | $\mathrm{U}(\mathrm{k})$ |
| $\mathrm{O}(\mathrm{N})$ | $\mathrm{Sp}(2 \mathrm{k})$ |
| $\mathrm{Sp}(2 \mathrm{~N})$ | $\mathrm{O}(\mathrm{k})$ |

$$
I_{M \mathbf{c}}=2 ; I_{M \mathbf{d}}=-2 \text { or } \quad I_{M \mathbf{d}^{\prime}}=2 ; I_{M \mathbf{c}^{\prime}}=-2
$$

Possible for:
(2 $\mathrm{U}, \mathrm{k}=1$ or 2

- $\mathrm{Sp}, \mathrm{k}=1$
- $\mathrm{O}, \mathrm{k}=1,2$


## UNIVERSAL INSTANTONINSTANTON ZERO-MODES

Q $\mathrm{U}(\mathrm{k}) \mathrm{i} 4 \mathrm{Adj}$<br>- $\operatorname{Sp}(2 \mathrm{k}): 2 \mathrm{~A}+2 \mathrm{~S}$<br>O $\mathrm{O}(\mathrm{k}): 2 \mathrm{~S}+2 \mathrm{~A}$

Only $O$ (1) has the required 2 zero modes
(See also: Argurio, Bertolini, Ferretti, Lerda,Peterson, arXiv:0704.0262)

## INSTANTON SCAN

Can we find such branes $M$ in the 391 models with massive B-L?

Q About 30.000 "instanton branes" (violations of the sum rule, i.e. $I_{M \mathbf{a}}-I_{M \mathbf{a}^{\prime}}-I_{M \mathrm{~d}}+I_{M \mathrm{~d}^{\prime}} \neq 0$ )
Q Violations between -8 and +8
Q Quantized in units of 1,2 or 4
(1 may give R-parity violation, 4 means no Majorana mass)
Q Some models have no RCFT instantons
Q 1315 instantons have the right number of zero modes, counted chirally.
Q None of these models has R-parity violating instantons.

| Tensor | MIPF | Orientifold | Instanton | Solution |
| :--- | :--- | :--- | :--- | :--- |
| $(2,4,18,28)$ | 17 | 0 |  |  |
| $(2,4,22,22)$ | 13 | 3 | $S 2^{+}!, S 2^{-}!$ | Yes! |
| $(2,4,22,22)$ | 13 | 2 | $S 2^{+}!, S 2^{-}$! | Yes |
| $(2,4,22,22)$ | 13 | 1 | $S 2^{+}, S 2^{-}$ | No |
| $(2,4,22,22)$ | 13 | 0 | $S 2^{+}, S 2^{-}$ | Yes |
| $(2,4,22,22)$ | 31 | 1 | $U 1^{+}, U 1^{-}$ | No |
| $(2,4,22,22)$ | 20 | 0 |  |  |
| $(2,4,22,22)$ | 46 | 0 |  |  |
| $(2,4,22,22)$ | 49 | 1 | $O 2^{+}, O 2^{-}, O 1^{+}, O 1^{-}$ | Yes |
| $(2,6,14,14)$ | 1 | 1 | $U 1^{+}$ | No |
| $(2,6,14,14)$ | 22 | 2 |  |  |
| $(2,6,14,14)$ | 60 | 2 |  |  |
| $(2,6,14,14)$ | 64 | 0 |  |  |
| $(2,6,14,14)$ | 65 | 0 |  | Yes |
| $(2,6,10,22)$ | 22 | 2 |  | No |
| $(2,6,8,38)$ | 16 | 0 |  |  |
| $(2,8,8,18)$ | 14 | 2 | $S 2^{+}!, S 2^{-}!$ | $S 2^{+}!, S 2^{-!}$ |
| $(2,8,8,18)$ | 14 | 0 | $U 1^{+}, U 1^{-}$ |  |
| $(2,10,10,10)$ | 52 | 0 |  |  |
| $(4,6,6,10)$ | 41 | 0 |  |  |
| $(4,4,6,22)$ | 43 | 0 |  |  |
| $(6,6,6,6)$ | 18 | 0 |  |  |

## A MODEL WITH S2 INSTANTONS

```
5 x (V ,V ,0,0 ) chirality 3
3x (V ,0,V ,0 ) chirality -3
3x (V ,0,V*,0 ) chirality -3
3 x (0,V ,0,V ) chirality 3
5 x (0,0,v ,V ) chirality -3
3x (0,0,V ,V*) chirality 3
6x (V,0,0,V )
18 x (0,V ,V ,0 )
    2 x (Ad,0,0,0 )
    2 x (A ,0,0,0 )
    2 x (S ,0,0,0 )
14 x (0,A ,0,0 )
    6x (0,S ,0,0 )
    9 x (0,0,Ad,0 )
    6 x (0,0,A ,0 )
14 x (0,0,S,0 )
    3x (0,0,0,Ad)
    4 x (0,0,0,A )
6x (0,0,0,S )
```

Gauge group: $\mathrm{SU}(3) \times \mathrm{SU}(2) \times \mathrm{U}(1) \times$ Nothing. Exactly the correct number of instanton zero modes
(except for 2 universal symmetric tensors)

## A MODEL WITH S2 INSTANTONS

```
5 x (V ,V ,0,0 ) chirality 3
3x (V ,0,V ,0 ) chirality -3
3x (V ,0,V*,0 ) chirality -3
3 x (0,V ,0,V ) chirality 3
5 x (0,0,v ,V ) chirality -3
3x (0,0,V ,V*) chirality 3
6x (V ,0,0,V )
18 x (0,V ,V ,0 )
    2 x (Ad,0,0,0 )
    2 x (A ,0,0,0 )
    2 x (S ,0,0,0 )
14 x (0,A ,0,0 )
    6 x (0,S,0,0 )
    9 x (0,0,Ad,0 )
    6 x (0,0,A ,0 )
14 x (0,0,S,0 )
    3x (0,0,0,Ad)
    4 x (0,0,0,A )
6x (0,0,0,S )
```

Gauge group: $\mathrm{SU}(3) \times \mathrm{SU}(2) \times \mathrm{U}(1) \times$ Nothing. Exactly the correct number of instanton zero modes
(except for 2 universal symmetric tensors)

$$
\begin{aligned}
\sin ^{2}\left(\theta_{w}\right) & =.5271853 \\
\frac{\alpha_{3}}{\alpha_{2}} & =3.2320501
\end{aligned}
$$

## THE O1 INSTANTON

Type:
Dimension

| U | S | U | U | U | O | O | U | O | 0 | O | U | S | S | O |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 3 | 2 | 1 | 1 | 1 | 2 | 2 | 3 | 1 | 2 | 3 | 1 | 2 | 2 | 2 |

$2 \mathrm{x}(\mathrm{0}, 0,0, \mathrm{~V}, 0,0,0,0,0,0,0,0,0,0,0, \mathrm{~V})$ chirality 2
$5 \mathrm{x}(\mathrm{V}, 0, \mathrm{~V}, 0,0,0,0,0,0,0,0,0,0,0,0,0$ ) chirality -3
$5 \mathrm{x}(0,0, \mathrm{~V}, \mathrm{~V} *, 0,0,0,0,0,0,0,0,0,0,0,0)$ chirality 3
$12 \mathrm{x}(0,0, \mathrm{~V}, 0,0,0,0,0,0,0,0,0,0,0,0, \mathrm{~V})$ chirality -2
$3 \mathrm{x}(\mathrm{V}, 0, \mathrm{~V} *, 0,0,0,0,0,0,0,0,0,0,0,0,0)$ chirality -3
$3 \mathrm{x}(0,0, V, V, 0,0,0,0,0,0,0,0,0,0,0,0)$ chirality -3
$3 \mathrm{x}(\mathrm{V}, \mathrm{V}, 0,0,0,0,0,0,0,0,0,0,0,0,0,0)$ chirality 3
$3 \mathrm{x}(0, \mathrm{~V}, 0, \mathrm{~V}, 0,0,0,0,0,0,0,0,0,0,0,0)$ chirality 3
$25 \mathrm{x}(0,0, A d, 0,0,0,0,0,0,0,0,0,0,0,0,0)$ chirality 0
$2 \mathrm{x}(\mathrm{A}, 0,0,0,0,0,0,0,0,0,0,0,0,0,0,0)$ chirality 0
$4 \mathrm{x}(\mathrm{V}, 0,0, \mathrm{~V}, 0,0,0,0,0,0,0,0,0,0,0,0)$ chirality 0
$2 \mathrm{x}(0,0,0, \mathrm{~A}, 0,0,0,0,0,0,0,0,0,0,0,0$ ) chirality 0
$34 \mathrm{x}(0,0, A, 0,0,0,0,0,0,0,0,0,0,0,0,0)$ chirality 0
$14 \mathrm{x}(0,0, S, 0,0,0,0,0,0,0,0,0,0,0,0,0)$ chirality 0
$2 \mathrm{x}(\mathrm{V}, 0,0,0,0,0,0,0,0,0,0,0,0,0,0, \mathrm{~V})$ chirality 0
$2 \mathrm{x}(0,0,0, \mathrm{~V}, 0,0,0,0,0,0,0,0,0, \mathrm{v}, 0,0)$ chirality 0
$1 \mathrm{x}(\mathrm{Ad}, 0,0,0,0,0,0,0,0,0,0,0,0,0,0,0)$ chirality 0
$2 \mathrm{x}(0, \mathrm{~S}, 0,0,0,0,0,0,0,0,0,0,0,0,0,0)$ chirality 0
$1 \mathrm{x}(0,0,0, A d, 0,0,0,0,0,0,0,0,0,0,0,0)$ chirality 0
$6 \mathrm{x}(0,0, \mathrm{~V}, 0,0,0,0,0,0,0,0,0,0, \mathrm{~V}, 0,0$ ) chirality 0
$2 \mathrm{x}(\mathrm{S}, 0,0,0,0,0,0,0,0,0,0,0,0,0,0,0)$ chirality 0
$2 \mathrm{x}(0,0,0, S, 0,0,0,0,0,0,0,0,0,0,0,0)$ chirality 0
$2 \mathrm{x}(0, \mathrm{~V}, 0,0,0,0,0,0,0,0,0, \mathrm{~V}, 0,0,0,0$ ) chirality 0
$1 \mathrm{x}(0, \mathrm{~V}, 0,0,0,0,0,0, V, 0,0,0,0,0,0,0)$ chirality 0
$1 \mathrm{x}(0, \mathrm{~V}, 0,0,0,0,0,0,0, \mathrm{~V}, 0,0,0,0,0,0)$ chirality 0
$1 \mathrm{x}(0, \mathrm{~V}, 0,0,0,0,0,0,0,0, V, 0,0,0,0,0)$ chirality 0
$2 \mathrm{x}(\mathrm{V}, 0,0, \mathrm{~V} *, 0,0,0,0,0,0,0,0,0,0,0,0)$ chirality 0
$2 \mathrm{x}(\mathrm{V}, 0,0,0,0,0,0, \mathrm{~V}, 0,0,0,0,0,0,0,0)$ chirality 0
$2 \mathrm{x}(0,0,0, \mathrm{~V}, 0,0,0, \mathrm{~V}, 0,0,0,0,0,0,0,0)$ chirality 0
$2 \mathrm{x}(0,0, \mathrm{~V}, 0,0,0,0,0,0,0,0,0,0,0, \mathrm{~V}, 0$ ) chirality 0
$6 \mathrm{x}(0, \mathrm{~V}, \mathrm{~V}, 0,0,0,0,0,0,0,0,0,0,0,0,0$ ) chirality 0
$6 \mathrm{x}(0,0, V, 0,0,0,0,0,0,0,0,0, V, 0,0,0)$ chirality 0
$2 \mathrm{x}(\mathrm{V}, 0,0,0, \mathrm{~V}, 0,0,0,0,0,0,0,0,0,0,0)$ chirality 0
$2 \mathrm{x}(0,0,0, \mathrm{~V}, \mathrm{~V}, 0,0,0,0,0,0,0,0,0,0,0)$ chirality 0
$3 \mathrm{x}(0,0,0,0, S, 0,0,0,0,0,0,0,0,0,0,0)$ chirality -1
$3 \mathrm{x}(0,0,0,0,0, \mathrm{~V}, 0,0,0,0,0, \mathrm{~V}, 0,0,0,0)$ chirality 1
$1 \mathrm{x}(0,0,0,0, \mathrm{~A}, 0,0,0,0,0,0,0,0,0,0,0)$ chirality -1
$2 \mathrm{x}(0,0,0,0, \mathrm{~V}, 0, \mathrm{~V}, 0,0,0,0,0,0,0,0,0)$ chirality 2

## CONCLUSIONS

Q Many desirable SM features can be realized in the RCFT orientifold landscape...

- Chiral SM spectrum

9 No mirrors
Q No adjoints, rank-2 tensors
9 No hidden sector
Q No hidden-observable massless matter
9 Matter free hidden sector
9 Exact $\operatorname{SU}(3) \times \operatorname{SU}(2) \times U(1)$
9 O1 instantons
but not all at the same time.
Q Neutrino masses:
"an incomplete success."
With sufficient statistics, O1 instantons
without superfluous zero-modes will be found.
Q Boundary state statistics:
12 million Unitary
3 million Orthogonal
2 million Symplectic $\Rightarrow 270000$ O1
(0) But what is the real reason why neutrino masses are small?


[^0]:    (*) Huiszoon, Dijkstra, Schellekens
    Phys.Lett.B609:408-417,2005, Nucl.Phys.B710:3-57,2005
    $(\dagger)$ Anastasopoulos, Dijkstra, Kiritsis, Schellekens

