Neutrino masses from Instantons in RCFT orientifolds



München 13-11-2007



ORIENTIFOLDS

ORIENTIFOLD PARTITION FUNCTIONS



ORIENTIFOLD PARTITION FUNCTIONS

$$\bigcirc \text{Closed} \qquad \frac{1}{2} \left[\sum_{ij} \chi_i(\tau) Z_{ij} \chi_i(\bar{\tau}) + \sum_i K_i \chi_i(2\tau) \right]$$

$$\bigcirc \text{Open} \qquad \frac{1}{2} \left[\sum_{i,a,n} N_a N_b A^i{}_{ab} \chi_i(\frac{\tau}{2}) + \sum_{i,a} N_a M^i{}_a \hat{\chi}_i(\frac{\tau}{2} + \frac{1}{2}) \right]$$

- i: Primary field label (finite range)
- a: Boundary label (finite range)
- χ_i : Character
- N_a : Chan-Paton (CP) Multiplicity

ALGEBRAIC CHOICES

- Basic CFT (N=2 tensor, free fermions...)
 (Type IIB closed string theory)
- Chiral algebra extension(*)
 May imply space-time symmetry (e.g. Susy: GSO projection).
 Reduces number of characters.
- Modular Invariant Partition Function (MIPF)(*) May imply bulk symmetry (e.g Susy), not respected by all boundaries. Defines the set of boundary states (Sagnotti-Pradisi-Stanev completeness condition)

(*) all these choices are simple current related

ACCESSIBLE RCFT'S

- Free fermions (4n real + (9-2n) complex)
- Sazama-Suzuki models (requires exact spectrum computation)
- Permutation orbifolds

....

(*) See also: Angelantonj et. al. Blumenbagen et. al. Aldazabal et. al. Brunner et. al.

MIPFs*

 $\begin{aligned} & \bigcirc \text{Choose a rational matrix } X_{\alpha\beta} \text{ obeying} \\ & 2X_{\alpha\beta} = Q_{J_{\alpha}}(J_{\beta}) \mod 1, \alpha \neq \beta \\ & X_{\alpha\alpha} = -h_{J_{\alpha}} \\ & N_{\alpha}X_{\alpha\beta} \in \mathbb{Z} \text{ for all } \alpha, \beta \\ & Q_{J}(a) = h(a) + h(J) - h(Ja) \end{aligned} \\ & \textcircled{O} \text{ This defines the torus partition function as} \\ & Z_{ij} \text{ is the number of currents } L \in \mathcal{H} \text{ such that} \\ & j = Li \end{aligned}$

 $Q_M(i) + X(M,L) = 0 \mod 1$ for all $M \in \mathcal{H}$.

*Gato-Rivera, Kreuzer, Schellekens (1991-1993)

ORIENTIFOLD CHOICES*

 $\beta_K(J)\beta_K(J') = \beta_K(JJ')e^{2\pi i X(J,J')} \quad , J, J' \in \mathcal{H}$

*Huiszoon, Sousa, Schellekens (1999-2000)

BOUNDARIES AND CROSSCAPS



$$R_{[a,\psi_a](m,J)} = \sqrt{\frac{|\mathcal{H}|}{|\mathcal{C}_a||\mathcal{S}_a|}} \psi_a^*(J) S_{am}^J$$

Generation Crosscap coefficients

$$U_{(m,J)} = \frac{1}{\sqrt{|\mathcal{H}|}} \sum_{L \in \mathcal{H}} e^{\pi i (h_K - h_{KL})} \beta_K(L) P_{LK,m} \delta_{J,0}$$

Cardy (1989) Sagnotti, Pradisi, Stanev (~1995) Huiszoon, Fuchs, Schellekens, Schweigert, Walcher (2000)

COEFFICIENTS



$$K^{i} = \sum_{m,J,J'} \frac{S^{i}{}_{m}U_{(m,J)}g^{\Omega,m}_{J,J'}U_{(m,J')}}{S_{0m}}$$





$$A^{i}_{[a,\psi_{a}][b,\psi_{b}]} = \sum_{m,J,J'} \frac{S^{i}_{\ m} R_{[a,\psi_{a}](m,J)} g^{\Omega,m}_{J,J'} R_{[b,\psi_{b}](m,J')}}{S_{0m}}$$

$$M_{[a,\psi_a]}^i = \sum_{m,J,J'} \frac{P_m^i R_{[a,\psi_a](m,J)} g_{J,J'}^{\Omega,m} U_{(m,J')}}{S_{0m}}$$

 $g_{J,J'}^{\Omega,m} = \frac{S_{m0}}{S_{mK}} \beta_K(J) \delta_{J',J^c}$

TADPOLES & ANOMALIES





TADPOLES & ANOMALIES

GTadpole cancellation condition: $\sum_{b} N_b R_{b(m,J)} = 4\eta_m U_{(m,J)}$ © Cubic TrF³ anomalies cancel

Semaining anomalies by Green-Schwarz mechanism

In rare cases, additional conditions for global anomaly cancellation* *Gato-Rivera,

*Gato-Rivera, Schellekens (2005)



MODEL BUILDING







Vector-like: mass allowed by $SU(3) \times SU(2) \times U(1)$ (Higgs, right-handed neutrino, gauginos, sparticles....)

THE MADRID MODEL*



(*) Ibanez, Marchesano, Rabadan

ABELIAN MASSES

Green-Schwarz mechanism



Axion-Vector boson vertex

·----

Generates mass vector bosons of anomalous symmetries (e.g. B + L) But may also generate mass for non-anomalous ones (Y, B-L)



SEARCHES

DATA

	2004-2005*	2005-2006†
Trigger	"Madrid"	All 3 family models
Chiral types	19	19345
Tadpole-free(per type)	18	1900
Total configs	$45 \ge 10^6$	145 x 10 ⁶
Tadpole free, distinct	210.000	1900
Max. primaries	∞	1750

(*) Huiszoon, Dijkstra, Schellekens

(†) Anastasopoulos, Dijkstra, Kiritsis, Schellekens

A "MADRID" MODEL

Gauge group: Exactly $SU(3) \times SU(2) \times U(1)!$ [U(3)×Sp(2)×U(1)×U(1), Massive B-L, No hidden sector]

3 x (V	V	0	(1) chirality 3	0
	,•	,0	(0) chirality (2)	
3 X (V	,0	,V	,0) chirality -3	\bigcup
3 x (V	,0	,V*	,0) chirality -3	D*
3 x (0	,V	,0	,V) chirality 3	L
5 x (0	,0	,V	,V) chirality -3	$E^{*}+(E+E^{*})$
3 x (0	,0	,V	,V*) chirality 3	N*
18 x (0	,V	,V	,0)	Higgs
2 x (V	,0	,0	,V)	00
2 x (Ad	,0	,0	,0)	
2 x (A	,0	,0	,0)	
6 x (S	,0	,0	,0)	
14 x (0	,A	,0	,0)	
6 x (0	,S	,0	,0)	
9 x (0	,0	,Ad	,0)	
6 x (0	,0	,А	,0)	
14 x (0	,0	,S	,0)	
3 x (0	,0	,0	,Ad)	
4 x (0	,0	,0	,A)	

6 x (0 ,0 ,0 ,S)

A "MADRID" MODEL

Gauge group: Exactly $SU(3) \times SU(2) \times U(1)!$ [U(3)×Sp(2)×U(1)×U(1), Massive B-L, No hidden sector]

3 x (V	,V	,0	,0) c	hirality 3	Q
3 x (V	,0	,V	,0) c	hirality -3	U*
3 x (V	,0	,V*	,0) c	hirality -3	D*
3 x (0	,V	,0	,V) c	hirality 3	L
5 x (0	,0	,V	,V) c	hirality -3	$E^{*}+(E+E^{*})$
3 x (0	,0	,V	,V*) c	hirality 3	N*
18 x (0	,V	,V	,0)		Higgs
2 x (V	,0	,0	,V)		
2 x (A	d ,0	,0	,0)		
2 x (A	,0	,0	,0)		
6 x (S	,0	,0	,0)	Vector-	like matter
14 x (0	,A	,0	,0)	V=vecto	or
6 x (0	,S	,0	,0)	A=Anti-	symm. tensor
9 x (0	,0	,Ad	,0)	S=Symr	netric tensor
6 x (0	,0	,A	,0)	Ad=Adj	oint
14 x (0	,0	,S	,0)		
- / -	-	-			

 $\begin{array}{c} 14 \times (0 & ,0 & ,S & ,0) \\ 3 \times (0 & ,0 & ,0 & ,Ad) \\ 4 \times (0 & ,0 & ,0 & ,A) \\ 6 \times (0 & ,0 & ,0 & ,S) \end{array}$

AN SU(5) MODEL

Gauge group is just SU(5)!



Top quark Yukawa's?



NEUTRINO MASSES

NEUTRINO MASSES*

 In field theory: easy; several solutions.
 Most popular: add three right-handed neutrinos add "natural" Dirac & Majorana masses (see-saw)

$$m_{\nu} = \frac{(M_D)^2}{M_M}; \quad M_D \approx 100 \text{ MeV}, \qquad M_M \approx 10^{11} \dots 10^{13} \text{ GeV}$$

In string theory: non-trivial.
 (String theory is much more falsifiable!).

Questions:

Can we find vacua with small neutrino masses? Are they generically small? If not, why do we observe small masses?

(*) Ibañez, Schellekens, Uranga, arXiv:0704.1079, JHEP (to appear) Blumenhagen, Cvetic, Weigand, hep-th/0609191 Ibañez, Uranga, hep-th/0609213

Other ideas: see e.g. Conlon, Cremades; Giedt, Kane, Langacker, Nelson; Buchmuller, Hamaguchi, Lebedev, Ratz,



NEUTRINOS IN MADRID MODELS

All these models have three right-handed neutrinos (required for cubic anomaly cancellation)

In most of these models: B-L survives as an exact gauge symmetry

Neutrino's can get Dirac masses, but not Majorana masses (both needed for see-saw mechanism).

In a very small* subset, B-L acquires a mass due to axion couplings.

(*) 391 out of 213000, all with $SU(3) \times Sp(2) \times U(1) \times U(1)$

B-L VIOLATION

But even then, B-L still survives as a perturbative symmetry. It may be broken to a discrete subgroup by instantons.

This possibility can be explored if the instanton is described by a RCFT brane M. B-L violation manifests itself as:

$$I_{M\mathbf{a}} - I_{M\mathbf{a}'} - I_{M\mathbf{d}} + I_{M\mathbf{d}'} \neq 0$$

 I_{Ma} = chiral [# (V,V*) - # (V*,V)] between branes M and a

a' = boundary conjugate of a







B-L ANOMALIES

$$I_{M\mathbf{a}} - I_{M\mathbf{a}'} - I_{M\mathbf{d}} + I_{M\mathbf{d}'} \neq 0$$

Implies a cubic B-L anomaly if M is a "matter" brane (Chan-Paton multiplicity $\neq 0$).

⇒ M cannot be a matter brane: non-gauge-theory instanton

Implies a $(B-L)(G_M)^2$ anomaly even if we cancel the cubic anomaly

 \Rightarrow B-L must be massive

(The converse is not true: there are massive B-L models without such instanton branes)

REQUIRED ZERO-MODES

Neutrino mass generation by non-gauge theory instantons*

The desired neutrino mass term v^cv^c violates c and d brane charge by two units. To compensate this, we must have

$$I_{M\mathbf{c}} = 2 ; I_{M\mathbf{d}} = -2 \text{ or } I_{M\mathbf{d}'} = 2 ; I_{M\mathbf{c}'} = -2$$

and all other intersections 0. (d' is the boundary conjugate of d)

(*)Blumenhagen, Cvetic, Weigand, hep-th/0609191 Ibañez, Uranga, hep-th/0609213

NEUTRINO-ZERO MODE COUPLING

The following world-sheet disk is allowed by all symmetries



 $L_{cubic} \propto d_a^{ij} (\alpha_i \nu^a \gamma_j) , a = 1, 2, 3$

ZERO-MODE INTEGRALS

$$\int d^2 \alpha \, d^2 \gamma \, e^{-d_a^{ij} \, (\alpha_i \nu^a \gamma_j)} = \nu_a \nu_b \left(\epsilon_{ij} \epsilon_{kl} d_a^{ik} d_b^{jl} \right)$$

Additional zero modes yield additional fermionic integrals and hence no contribution

Therefore $I_{Ma}=I_{Mb}=I_{Mx}=0$ (x = Hidden sector), and there should be no vector-like zero modes.

There should also be no instanton-instanton zero-modes except 2 required by susy.

INSTANTON TYPES

In orientifold models we can have complex and real branes

Matter brane M	Instanton brane M
U(N)	U(k)
O(N)	Sp(2k)
Sp(2N)	O(k)

 $I_{M\mathbf{c}} = 2$; $I_{M\mathbf{d}} = -2$ or $I_{M\mathbf{d}'} = 2$; $I_{M\mathbf{c}'} = -2$

Possible for:

UNIVERSAL INSTANTON-INSTANTON ZERO-MODES

 $\bigcirc U(k): 4 Adj$ $\bigcirc Sp(2k): 2 A + 2 S$ $\bigcirc O(k): 2 A + 2 S$

Only O(1) has the required 2 zero modes

INSTANTON SCAN

Can we find such branes M in the 391 models with massive B-L?

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(1 may give R-parity violation, 4 means no Majorana mass)

Some models have no RCFT instantons

- 9 1315 instantons with correct *chiral* intersections
- Some of these models has R-parity violating instantons.
- Most instantons are symplectic in this sample.
- Solution of the spurious extra susy zero-modes (Sp(2) instantons).

INSTANTON SCAN

Can we find such branes M in the 391 models with massive B-L?

(1 may give R-parity violation, 4 means no Majorana mass)

Some models have no RCFT instantons

- 9 1315 instantons with correct *chiral* intersections
- Some of these models has R-parity violating instantons.
- Most instantons are symplectic in this sample.
- There are examples with exactly the right number, *non-chirally*, except for the spurious extra susy zero-modes (Sp(2) instantons).

...almost

Tensor	MIPF	Orientifold	Instanton	Solution
(1, 16, 16, 16)	12	0	$S2^+, S2^-$	Yes
(2,4,12,82)	19	0	S2 ⁻ !	?
(2,4,12,82)	19	0	$U2^{+}!, U2^{-}!$	No
(2,4,12,82)	19	0	$U1^{+}, U1^{-}$	No
(2,4,14,46)	10	0		
(2,4,14,46)	16	0		
(2,4,16,34)	15	0		
(2,4,16,34)	15	1		
(2,4,16,34)	16	0	$S2^+, S2^-$	Yes
(2,4,16,34)	16	1		
(2,4,16,34)	18	0	$S2^{-}$	Yes
(2,4,16,34)	18	0	$U1^+, U1^-, U2^+, U2^-$	No
(2,4,16,34)	49	0	$U2^+, S2^-!, U1^+$	Yes
			Continued on :	next page

Tensor	MIPF	Orientifold	Instanton	Solution
(2,4,18,28)	17	0		
(2,4,22,22)	13	3	$S2^+!, S2^-!$	Yes!
(2,4,22,22)	13	2	$S2^+!, S2^-!$	Yes
(2,4,22,22)	13	1	$S2^+, S2^-$	No
(2,4,22,22)	13	0	$S2^+, S2^-$	Yes
(2,4,22,22)	31	1	$U1^+, U1^-$	No
(2,4,22,22)	20	0		
(2,4,22,22)	46	0		
(2,4,22,22)	49	1	$O2^+, O2^-, O1^+, O1^-$	Yes
(2,6,14,14)	1	1	$U1^+$	No
(2,6,14,14)	22	2		
(2,6,14,14)	60	2		
(2,6,14,14)	64	0		
(2,6,14,14)	65	0		
(2,6,10,22)	22	2		
(2,6,8,38)	16	0		
(2,8,8,18)	14	2	$S2^+!, S2^-!$	Yes
(2,8,8,18)	14	0	$S2^{+}!, S2^{-}!$	No
(2,10,10,10)	52	0	$U1^+, U1^-$	No
(4, 6, 6, 10)	41	0		
(4,4,6,22)	43	0		
(6, 6, 6, 6)	18	0		

AN SP(2) INSTANTON MODEL

U3 S2 U1 U1 O

3	Х	(V	,V	,0	,0	,0)	chirality	3
3	х	(V	,0	,V	,0	,0)	chirality	-3
3	х	(V	,0	,V*	,0	,0)	chirality	-3
3	х	(0	,V	,0	,V	,0)	chirality	3
5	х	(0	,0	,V	,V	,0)	chirality	-3
3	х	(0	,0	,V	,V*	,0)	chirality	3
1	х	(0	,0	,V	,0	,V)	chirality	-1
1	х	(0	,0	,0	,V	,V)	chirality	1
18	Х	(0	,V	,V	,0	,0)		
2	х	(V	,0	,0	,V	,0)		
2	Х	(Ad,	, 0	,0	,0	,0)		
2	х	(А	,0	,0	,0	,0)		
6	х	(S	,0	,0	,0	,0)		
14	Х	(0	,А	,0	,0	,0)		
6	х	(0	,S	,0	,0	,0)		
9	х	(0	,0	,Ad,	0	,0)		
6	Х	(0	,0	,А	,0	,0)		
14	Х	(0	,0	,S	,0	,0)		
3	Х	(0	,0	,0	,Ad,	0)		
4	Х	(0	,0	,0	,А	,0)		
6	х	(0	,0	,0	,S	,0)		

AN SP(2) INSTANTON MODEL

U3 S2 U1 U1 O

	3	х	(۷	,V	,0	,0	,0)	chirality	3
	3	Х	(V	,0	,V	,0	,0)	chirality	-3
	3	Х	(۷	,0	,V*	,0	,0)	chirality	-3
	3	Х	(0	,V	,0	,V	,0)	chirality	3
	5	Х	(0	,0	,V	,V	,0)	chirality	-3
	3	Х	(0	,0	,V	,V*	,0)	chirality	3
	1	Х	(0	,0	,V	,0	,V)	chirality	-1
	1	Х	(0	,0	,0	,V	,V)	chirality	1
1	8	Х	(0	,V	,V	,0	,0)		
	2	Х	(V	,0	,0	,V	,0)		
	2	Х	(Ad,	, 0	,0	,0	,0)		
	2	Х	(А	,0	,0	,0	,0)		
	6	Х	(S	,0	,0	,0	,0)		
1	14	Х	(0	,А	,0	,0	,0)		
	6	Х	(0	,S	,0	,0	,0)		
	9	Х	(0	,0	,Ad,	0	,0)		
	6	Х	(0	,0	,А	,0	,0)		
1	4	Х	(0	,0	,S	,0	,0)		
	3	Х	(0	,0	,0	,Ad,	0)		
	4	Х	(0	,0	,0	,А	,0)		
	6	Х	(0	,0	,0	,S	,0)		

THE O1 INSTANTON

Type:

Гуре	:			U	S	U	U	U	0	0	U	0	0	0	U	S	S	0	S	Ľ.	
Dime	nsid	on	8	3	2	1	1	1	2	2	3	1	2	3	1	2	2	2		ł.	
	5	х	(V	,0	,v	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0)	chirality -3
	5	х	(0	,0	,v	,V*	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0)	chirality 3
	3	x	(V	,0	,V*	*,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0)	chirality -3
	3	х	(0	,0	,v	,v	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0)	chirality -3
	3	х	(v	,v	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0)	chirality 3
	3	x	(0	,v	,0	,v	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0)	chirality 3
	2	х	(0	,0	,0	,V	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,v)	chirality 2
	12	х	(0	,0	,V	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,V)	chirality -2
	1	х	(0	,0	,0	,0	,0	,0	,0	,0	,0	, V	,0	,0	,0	,0	,0	,V)	
	2	х	(0	,0	,0	,0	, V	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,V)	
	1	х	(0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,v	,V)	
	2	х	(0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,v	,0	,0	,V)	
	1	х	(0	,0	,0	,0	,0	, V	,0	,0	,0	,0	,0	,0	,0	,0	,0	,V)	
	3	х	(0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,s)	
	4	х	(0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,v	,0	,V)	
	2	х	(0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	, A)	
	2	х	(V	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,V)	
	3	х	(0	,0	,0	,0	,s	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0)	chirality -1
	3	х	(0	,0	,0	,0	,0	,v	,0	,0	,0	,0	,0	,v	,0	,0	,0	,0)	chirality 1
	1	x	(0	,0	,0	,0	,A	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0)	chirality -1
	2	x	(0	,0	,0	,0	,v	,0	,v	,0	,0	,0	,0	,0	,0	,0	,0	,0)	chirality 2
	1	х	(0	,0	,0	,0	,0	,0	,0	,v	,0	,0	,0	,0	,0	,0	,v	,0)	chirality -1
	1	x	(0	,0	,0	,0	, V	,0	,0	,0	,0	,v	,0	,0	,0	,0	,0	,0)	chirality -1
	1	х	(0	,0	,0	,0	,0	,0	,0	,0	,v	,0	,0	,v	,0	,0	,0	,0)	chirality 1
	1	х	(0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,v	,v	,0	,0	,0	,0)	chirality -1
	1	х	(0	,0	,0	,0	,0	,0	,v	,0	,0	,0	,0	,v	,0	,0	,0	,0)	chirality -1
	1	х	(0	,0	,0	,0	,V	,v	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0)	chirality -1
	1	х	(0	,0	,0	,0	,v	,0	,0	,0	,0	,0	,0	,v	,0	,0	,0	,0)	chirality 1
	1	х	(0	,0	,0	,0	,V	,0	,0	,0	,0	,0	,0	,V*	*,0	,0	,0	,0)	chirality -1
	3	х	(0	,0	,0	,0	,v	,0	,0	,0	,0	,0	,0	,0	,0	,0	,v	,0)	chirality 1
	1	х	(0	,0	,0	,0	,0	,0	,0	,v	,0	,v	,0	,0	,0	,0	,0	,0)	chirality 1
	2	х	(0	,0	,0	,v	,0	,0	,0	,0	,0	,0	,0	,0	,0	,v	,0	,0)	
	1	х	(Ac	1,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0)	
	2	x	(0	,s	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0)	
	1	х	(0	,0	,0	,Ad	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0)	
	6	х	(0	,0	,v	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,V	,0	,0)	
	1	х	(0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	,0	, A	,0)	
	1	x	(0	.0	.0	.0	. Ac	0.1	.0	.0	.0	.0	.0	.0	.0	.0	.0	. 0)	

OTHER OPERATORS

Operator	U	S	Ο
$ u_R \nu_R$	3	627	3
LHLH	5	550	3
LH	6	0	4
QDL	12	0	4
UDD	0	0	4
LLE	12	0	4
QQQL	4	0	3892
UUDE	4	0	3880

CONCLUSIONS

Seutrino masses: "an incomplete success." With sufficient statistics, O1 instantons without superfluous zero-modes will be found. Boundary state statistics: 12 million Unitary 3 million Orthogonal 2 million Symplectic => 270000 O1 But what is the real reason why neutrino masses are small?