# ACADEMIC LECTURES

BEYOND THE STANDARD MODEL



## SUPERSYMMETRY

#### SUPERSYMMETRY

# A symmetry between fermions and bosons

## **DE-MOTIVATIONS**

- \* No mass degeneracies among SM-particles with different spin: Not an exact symmetry.
- \*No SM-particles are each others partners: Doubling of the spectrum.
- Even that is not enough;Two Higgses are needed.
- \*Nucleon stability is not automatic.
- Huge number of parameters.

## MOTIVATIONS

#### **Nice**

Finiteness

String Theory

**W**Hierarchy problem

**Q**Dark Matter

Coupling constant convergence

## MOTIVATIONS

**Nice** 

**G**Finiteness

String Theory

Hierarchy problem

**Q**Dark Matter

Coupling constant convergence



Confidence

∼10%

**~**10%

~15%

∼20%

#### THE HIERARCHY PROBLEM

Loop correction to scalar masses



$$m_{\rm phys}^2 = m_{\rm bare}^2 + g\Lambda^2 << \Lambda^2$$

Fine tuning

# FERMIONS VS. SCALARS

+

+



#### Fermion

Scalar

 $\delta_m \propto g^2 m \log(\Lambda/m)$ 

2000

 $\delta_m = q\Lambda^2$ 

#### How does Susy solve this?

+

+







#### Fermion-Boson cancellation (if couplings match)

#### **TECHNICAL NATURALNESS**

Dirac Naturalness: Parameters should be of order 1 in natural units.

\* 't Hooft naturalness (Technical Naturalness): A parameter is naturally small if setting it to zero enhances the symmetry of the theory.

#### NATURALNESS

Some examples

Natural:  $\frac{m_{\rm top}}{M_Z}$ 

Unnatural, but technically natural:

 $i\bar{\psi}_L \gamma^\mu D_\mu \psi_L + i\bar{\psi}_R \gamma^\mu D_\mu \psi_R + m\bar{\psi}_L \psi_R + m\bar{\psi}_R \psi_L$ for m = 0 we can rotate  $\psi_L$  and  $\psi_R$ by *separate* phases (chiral symmetry)

 $m_e$ 

 $m_{\tau}$ 

Unnatural (by any definition) $m_Z$ (In SM + gravity) $m_{Planck}$ 

#### SUSY & THE HIERARCHY PROBLEM

- A priori, supersymmetry *only* solves the *technical* naturalness problem.
- It does not explain why M<sub>weak</sub> is much smaller than M<sub>planck</sub>.
  (cf. QCD and "dimensional transmutation")
- In fact, susy has a Higgs mass parameter that is unnatural (but technically natural): µ
- In supersymmetric theories additional mechanisms exist that do explain this ratio (require large M<sub>top</sub>).
- Cosmological constant hierarchy problem much worse, and not solved by Susy

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Does this justify (more than) doubling the particle spectrum?

## THE WESS-ZUMINO MODEL

 $\mathcal{L} = \mathcal{L}_{boson} + \mathcal{L}_{fermion}$ 

 $\mathcal{L}_{\rm boson} = \eta^{\mu\nu} \partial_{\mu} \phi^{\dagger} \partial_{\nu} \phi$ 

 $\mathcal{L}_{\text{fermion}} = i\psi\bar{\sigma}^{\mu}\partial_{\mu}\psi$ 

 $\sigma^{\mu} = (1, \vec{\tau})$  $\bar{\sigma}^{\mu} = (1, -\vec{\tau})$ 

 $\gamma^5 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$  $\gamma^{\mu} = \begin{pmatrix} 0 & -i\sigma^{\mu} \\ -i\bar{\sigma}^{\mu} & 0 \end{pmatrix}$ 

Free complex boson + left-handed (Weyl)fermion

Susy transformation on the scalar ( $\psi$  is a 2-component spinor)

$$\delta_{\varepsilon}\phi = \sqrt{2}\varepsilon\psi \equiv \sqrt{2}\varepsilon^{\alpha}\psi_{\alpha}$$

 $\epsilon$  is a constant spinor

$$\delta_{\varepsilon} \mathcal{L}_{\text{scalar}} = \sqrt{2} \left( \varepsilon \partial^{\mu} \psi \partial_{\mu} \phi^{\dagger} + \bar{\varepsilon} \partial^{\mu} \bar{\psi} \partial_{\mu} \phi \right)$$

Now we look for a transformation of the fermion Lagrangian to cancel this; let us try

$$\delta_{\varepsilon}\psi_{\alpha} = i\lambda(\sigma^{\mu}\bar{\varepsilon})_{\alpha}\partial_{\mu}\phi = i\lambda\sigma^{\mu}_{\alpha\dot{\beta}}\bar{\varepsilon}^{\dot{\beta}}\partial_{\mu}\phi$$

$$\delta_{\varepsilon} \mathcal{L}_{\text{fermion}} = \lambda (\varepsilon \sigma^{\mu} \partial_{\mu} \phi^{\dagger} \bar{\sigma}^{\nu} \partial_{\nu} \psi - \bar{\psi} \bar{\sigma}^{\mu} \sigma^{\nu} \bar{\varepsilon} \partial_{\mu} \partial_{\nu} \phi) = \lambda (-\varepsilon \sigma^{\mu} \bar{\sigma}^{\nu} \psi \partial_{\mu} \partial_{\nu} \phi^{\dagger} - \bar{\psi} \bar{\sigma}^{\mu} \sigma^{\nu} \bar{\varepsilon} \partial_{\mu} \partial_{\nu} \phi)$$

Because of the symmetric appearance of the derivatives we may replace  $[\sigma^{\mu}\bar{\sigma}^{\nu}]^{\beta}_{\alpha}$  by

$$\frac{1}{2} [\sigma^{\mu} \bar{\sigma}^{\nu} + \sigma^{\nu} \bar{\sigma}^{\mu}]^{\beta}_{\alpha} = \eta^{\mu\nu} \delta^{\beta}_{\alpha}$$

Cancels bosonic variation if

$$\lambda = -\sqrt{2}$$

#### SUSY TRANSFORMATIONS

 $\delta_{\varepsilon}\phi = \sqrt{2}\varepsilon\psi \equiv \sqrt{2}\varepsilon^{\alpha}\psi_{\alpha}$  $\delta_{\varepsilon}\psi_{\alpha} = -i\sqrt{2}\sigma^{0}(\sigma^{\mu}\bar{\varepsilon})_{\alpha}\partial_{\mu}\phi$ 

Define an operator that generates this transformation on all fields

$$\left(\varepsilon Q + \bar{Q}\bar{\varepsilon}\right)X = \delta_{\varepsilon}X$$
$$X = \phi \text{ or } \psi$$

## SUSY COMMUTATOR

 $\left[\varepsilon_1 Q + \bar{Q}\bar{\varepsilon}_1, \varepsilon_2 Q + \bar{Q}\bar{\varepsilon}_2\right] = -2i(\varepsilon_2 \sigma^{\mu}\bar{\epsilon}_1 - \varepsilon_1 \sigma^{\mu}\bar{\epsilon}_2)\partial_{\mu}$ 

Or, equivalently

 $\{Q_{\alpha}, \bar{Q}_{\dot{\alpha}}\} = 2i\sigma^{\mu}_{\alpha\dot{\alpha}}\partial_{\mu}$ 

#### SUSY ÅLGEBRA

 $[Q_{\alpha}, P_{\mu}] = 0$  $\{Q_{\alpha}, Q_{\beta}\} = 0$  $\{\bar{Q}_{\dot{\alpha}}, \bar{Q}_{\dot{\beta}}\} = 0$  $\{Q_{\alpha}, \bar{Q}_{\dot{\beta}}\} = 2\sigma^{\mu}_{\alpha\dot{\beta}}P_{\mu}$ 

Non-trivial extension of the Poincaré Algebra.

## VACUUM ENERGY

 $\{Q_{\alpha}, \bar{Q}_{\dot{\beta}}\} = 2\sigma^{\mu}_{\alpha\dot{\beta}}P_{\mu}$  Implies:  $H = P^{0} = \frac{1}{4}(\bar{Q}_{1}Q_{1} + Q_{1}\bar{Q}_{1} + \bar{Q}_{2}Q_{2} + Q_{2}\bar{Q}_{2})$ Therefore

 $\langle \Psi | H | \Psi \rangle = \frac{1}{4} (|Q_1 | \Psi \rangle |^2 + |\bar{Q}_1 | \Psi \rangle |^2) + |Q_2 | \Psi \rangle |^2 + |\bar{Q}_2 | \Psi \rangle |^2)$ 

If the vacuum is supersymmetric

 $Q_{\alpha} |0\rangle = \bar{Q}_{\dot{\alpha}} |0\rangle = 0 \qquad \longrightarrow \qquad \langle 0| H |0\rangle = 0$ 

Vacuum energy is  $\pm^{\infty}$  in non-supersymmetric QFT

## SUPERMULTIPLETS

**Chiral Multiplet(\*)**complex scalar + left-handed fermion

**Vector Multiplet**Vector + Majorana fermion

**Graviton Multiplet** Graviton + Gravitino

(\*)CPT-Conjugate: complex scalar + right-handed fermion [not needed]

#### (EXTENDED SUPERSYMMETRY)

$$Q^i_{\alpha}; \quad i=1,\ldots N$$

#### Generates larger multiplets.

- ☑ N=1: previous slide.
- N=2: smallest multiplet has two complex scalars, plus a left and a right-handed spinor. Non-chiral!
- N=4: smallest multiplet contains a vector, six complex scalars plus four left- and right-handed spinors. Finite!
- N=8: smallest multiplet contains graviton.
- ☑ N>8: smallest multiplet contains spin 5/2

## SUPERMULTIPLETS (N=1)

**Chiral Multiplet(\*)**complex scalar + left-handed fermion

**Vector Multiplet**Vector + Majorana fermion

Graviton Multiplet
 Graviton + Gravitino

(\*)CPT-Conjugate: complex scalar + right-handed fermion [not needed]

#### PHYSICAL STATE COUNTING

Туре	Example	# d.o.f
Real Scalar	$\pi^0$	1
Complex Scalar	$\pi^+,\pi^-$	2
Dirac fermion	$e_L^-, e_R^-, e_L^+, e_R^+$	4
Weyl fermion	$e_L^-, e_R^+ \text{ (mass = 0)}$	2
Majorana fermion	$\nu_R, \nu_L^c (\text{charge} = 0)$	2
Vector boson	photon	2
Gravitino		2
Graviton		2

# THE SSM

SM particle	SSM partner	Multiplet
$e_L^ (l_L^-)$	selectron <sub>1</sub> (slepton <sub>1</sub> )	Chiral
$e_L^+$ $(l_L^+)$	selectron <sub>2</sub> (slepton <sub>2</sub> )	Chiral
$q_L$	squark1	Chiral
$q_R$	squark <sub>2</sub>	Chiral
photon	photino (Majorana fermion)	Vector
gluon	gluino (Eight Majorana fermions)	Vector
$W^+, W^-, Z$	Wino <sup>±</sup> , Zino (Three Majorana fermions)	Vector
Higgs	???	Chiral

 $\phi: (1,2,\frac{1}{2}) \rightarrow \text{Weyl fermion}(1,2,\frac{1}{2})_L$   $\downarrow \text{CPT}$   $\phi^*: (1,2,-\frac{1}{2}) \rightarrow \text{Weyl fermion}(1,2,-\frac{1}{2})_L$ 

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 $\phi: (1, 2, \frac{1}{2}) \rightarrow \text{Weyl fermion}(1, 2, \frac{1}{2})_L$ CPT

 $\phi^*: (1,2,-\frac{1}{2}) \to \text{Weyl fermion}(1,2,-\frac{1}{2})_L$ 

#### Two distinct options for supermultiplet of SM Higgs

$$\phi: (1,2,\frac{1}{2}) \rightarrow \text{Weyl fermion}(1,2,\frac{1}{2})_L$$
  
CPT

 $\phi^*: (1,2,-\frac{1}{2}) \to \text{Weyl fermion}(1,2,-\frac{1}{2})_L$ 

or

 $H_2$ 

 $H_1$ 

Two distinct options for supermultiplet of SM Higgs

Both are needed to cancel anomalies

# THE MSSM (1)

MSSM spectrum:

Quarks + sQuarks
 Leptons+sLeptons
 Gauge bosons + gauginos
 H<sub>1</sub>,H<sub>2</sub> + Higgsinos
 **•** NOTHING

#### INTERACTIONS

SSM action:

 $\int d^4x (d^2\theta \mathcal{L}_F + c.c) + \int d^4x d^4\theta \mathcal{L}_D$ 

"F-terms"

"D-terms"

Origin of:

Most interactions; Gauge kinetic terms Scalar and fermion kinetic terms and their gauge couplings

Hard

$$-\frac{1}{4} \sum_{I=1}^{12} F^{I}_{\mu\nu} F^{\mu\nu,I}$$

$$i\sum_{\ell=1}^{15} \bar{\psi}_{\ell} \gamma^{\mu} D_{\mu} \psi_{\ell}$$

$$- (D_{\mu}\phi)^{\dagger}(D^{\mu}\phi) - \mu^{2}\phi^{\dagger}\phi - \frac{1}{4}\lambda(\phi^{\dagger}\phi)^{2}$$

 $g_{\mathcal{U}}^{\alpha\beta}\bar{\psi}_{L}^{\mathcal{Q},\alpha} \ [\mathbf{C}\phi^{*}]\psi_{R}^{\mathcal{U},\beta} + g_{\mathcal{D}}^{\alpha\beta}\bar{\psi}_{L}^{\mathcal{Q},\alpha}\phi\psi_{R}^{\mathcal{D},\beta} + g_{\mathcal{E}}^{\alpha\beta}\bar{\psi}_{L}^{\mathcal{L},\alpha}\phi\psi_{R}^{\mathcal{E},\beta} + \text{c.c.}$ 

(+ neutrino contributions)

 $(+F_{\mu\nu}\tilde{F}^{\mu\nu} \text{ terms})$ 

$$-\frac{1}{4} \sum_{I=1}^{12} F^{I}_{\mu\nu} F^{\mu\nu,I} + i \sum_{\ell=1}^{15} \bar{\psi}_{\ell} \gamma^{\mu} D_{\mu} \psi_{\ell}$$
  
+  $(D_{\mu}\phi)^{\dagger} (D^{\mu}\phi) - \frac{\mu^{2} \phi^{\dagger} \phi}{4} - \frac{1}{4} \lambda (\phi^{\dagger}\phi)^{2} +$ 

$$g_{\mathcal{U}}^{\alpha\beta}\bar{\psi}_{L}^{\mathcal{Q},\alpha} \ [\mathbf{C}\phi^{*}]\psi_{R}^{\mathcal{U},\beta} + g_{\mathcal{D}}^{\alpha\beta}\bar{\psi}_{L}^{\mathcal{Q},\alpha}\phi\psi_{R}^{\mathcal{D},\beta} + g_{\mathcal{E}}^{\alpha\beta}\bar{\psi}_{L}^{\mathcal{L},\alpha}\phi\psi_{R}^{\mathcal{E},\beta} + \text{c.c.}$$

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(+ neutrino contributions)

$$(+F_{\mu\nu}\tilde{F}^{\mu\nu} \text{ terms})$$

#### SUPERSPACE

Extend space-time by four anti-commuting variables

 $\theta_{\alpha}, \alpha = 1, 2$  $\theta_{\alpha}, \alpha = 1, 2$ (Spinor index)  $\{\theta_{\alpha}, \theta_{\beta}\} = 0$  $\{\theta_{\alpha}, \theta_{\beta}\} = 0$  $\{\bar{\theta}_{\alpha},\bar{\theta}_{\beta}\}=0$  $\theta^2 \equiv \theta_1 \theta_2$  $\theta_1^2 = \theta_2^2 = \bar{\theta}_1^2 = \bar{\theta}_2^2 = 0$  $\bar{\theta}^2 \equiv \bar{\theta}_1 \bar{\theta}_2$ Superspace integrals  $\begin{cases} \int d^4x d^2\theta \\ \int d^4x d^2\theta d^2\bar{\theta} \equiv \int d^4x d^4\theta \end{cases}$
# CHIRAL SUPERFIELDS

For each supermultiplet  $(\phi, \psi_L)$  define a *Chiral Superfield* 

$$\phi_L(x,\theta) = \varphi(x) + \sqrt{2\theta}\psi_L(x) + \theta^2 F(x)$$
  

$$F(x): \text{ auxiliary field}$$

Conjugate:  $\phi_L^{\dagger}(x,\bar{\theta}) = \varphi^*(x) + \sqrt{2}\bar{\psi}_L(x)\bar{\theta} + \bar{\theta}^2 F^*(x)$ 

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# VECTOR SUPERFIELDS

To describe vector bosons we need an additional kind of superfield

Vector Superfield

 $V(x,\theta,\bar{\theta}) = -\theta \rho_{\mu} \bar{\theta} V^{\mu} + i\theta^2 \bar{\theta} \bar{\lambda} - i\bar{\theta}^2 \theta \lambda + \frac{1}{2} \theta^2 \bar{\theta}^2 D$  $V^{\mu} \text{ is a vector boson}$  $\lambda \text{ is a Majorana fermion}$ D is the auxiliary field

Satisfies  $V = V^{\dagger}$ 

# INTERACTIONS

SSM action:

 $\int d^4x (d^2\theta \mathcal{L}_F + c.c) + \int d^4x d^4\theta \mathcal{L}_D$ 

Supersymmetric if and only if  $\mathcal{L}_F$  is only a function of  $\theta$  (and NOT of  $\overline{\theta}$ )  $\rightarrow \mathcal{L}_F$  is a chiral superfield  $\mathcal{L}_D = (\mathcal{L}_D)^{\dagger}$   $\rightarrow \mathcal{L}_D$  is a vector superfield Built with Fundamental Superfields

# THE SUPERPOTENTIAL

F-terms 
$$\int d^4x (d^2\theta \mathcal{L}_F + c.c)$$

 $d^2\theta \equiv$  "Expand in  $\theta$  and keep only the quadratic terms"  $\mathcal{L}_F =$  gauge kinetic terms +  $W(\phi)$ (Superpotential)

Superpotential:  $W(\phi) = \text{Any polynomial in the superfields}$ (but NOT their conjugates) Contains most of the information about couplings



Superpotential contains all allowed terms Renormalizability: at most order 3 in superfield

# **D-TERMS**

 $\int d^4x d^4\theta \mathcal{L}_D$ 

 $\mathcal{L}_D = (\mathcal{L}_D)^{\dagger}$ 

 $d^4\theta \equiv$  "expand to order  $\theta^2 \bar{\theta}^2$  and take its coefficient" Example

Yields

$$\mathcal{L}_D = \phi^{\dagger} e^{2gV} \phi$$

$$-|D_{\mu}\varphi|^{2} - i\psi\sigma^{\mu}D_{\mu}\bar{\psi} + 2ig[\varphi^{*}\lambda\psi - \varphi\bar{\lambda}\bar{\psi}] + FF^{*} + g\varphi^{*}D\varphi$$



A single chiral superfield  $\phi$ , with superpotential

$$W(\phi) = \frac{1}{2}m\phi^2 + \frac{1}{3}\lambda\phi^3$$
$$\phi(x,\theta) = \varphi(x) + \sqrt{2}\theta\psi(x) + \theta^2 F(x)$$

Coefficient of  $\theta^2$ :

$$m(\varphi F - \frac{1}{2}\psi^2) + \lambda(F\varphi^2 - \varphi\psi^2)$$

Kinetic terms (from D-terms)

$$-\partial_{\mu}\varphi\partial^{\mu}\varphi + i\psi\sigma^{\mu}\partial_{\mu}\bar{\psi} + FF^{*}$$

#### **ELIMINATION OF AUXILIARY FIELDS**

#### Complete action:

 $-\partial_{\mu}\varphi\partial^{\mu}\varphi + i\psi\rho^{\mu}\partial_{\mu}\bar{\psi} + FF^* + \left[m(\varphi F - \frac{1}{2}\psi^2) + \lambda(F\varphi^2 - \varphi\psi^2) + c.c\right]$ 

#### Equation of motion for F:

$$F = -m\varphi^* - \lambda^*(\varphi^*)^2$$

Substitute back into action:

 $\mathcal{L} = -\partial_{\mu}\varphi\partial^{\mu}\varphi + i\psi\rho^{\mu}\partial_{\mu}\bar{\psi} - \frac{1}{2}m(\psi^{2} + \bar{\psi}^{2}) - \lambda\varphi\psi^{2} - \lambda^{*}\varphi^{*}\bar{\psi}^{2} - |m\varphi + \lambda\varphi^{2}|^{2}$ 

#### **ELIMINATION OF AUXILIARY FIELDS**

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Note: Zero electric charge (Majorana mass allowed)

# REMARKS

 $\mathcal{L} = -\partial_{\mu}\varphi\partial^{\mu}\varphi + i\psi\rho^{\mu}\partial_{\mu}\bar{\psi} - \frac{1}{2}m(\psi^{2} + \bar{\psi}^{2}) - \lambda\varphi\psi^{2} - \lambda^{*}\varphi^{*}\bar{\psi}^{2} - |m\varphi + \lambda\varphi^{2}|^{2}$ 

Scalar and fermion have equal mass

Quartic terms derived from cubic and quadratic terms in the superpotential.
 (no additional parameters).

# THE MSSM INTERACTIONS: THE GOOD

#### Yukawa's

# $\mathcal{L}_{Y} = g_{\mathcal{U}}^{\alpha\beta} \bar{\psi}_{L}^{\mathcal{Q},\alpha} [\mathbf{C}\phi^{*}] \psi_{R}^{\mathcal{U},\beta} + g_{\mathcal{N}}^{\alpha\beta} \bar{\psi}_{L}^{\mathcal{L},\alpha} [\mathbf{C}\phi^{*}] \psi_{R}^{\mathcal{N},\beta} + g_{\mathcal{D}}^{\alpha\beta} \bar{\psi}_{L}^{\mathcal{Q},\alpha} \phi \psi_{R}^{\mathcal{D},\beta} + g_{\mathcal{E}}^{\alpha\beta} \bar{\psi}_{L}^{\mathcal{L},\alpha} \phi \psi_{R}^{\mathcal{E},\beta} + \text{c.c.}$

# THE MSSM INTERACTIONS: THE GOOD

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$$\begin{array}{cccc} (3,2,\frac{1}{6}) & \begin{pmatrix} u_L \\ d_L \end{pmatrix} & \mathcal{Q} \\ d_L \end{pmatrix} & \mathcal{U} \\ 3^*,1,-\frac{2}{3}) & u_L^c & \mathcal{U} \\ (3^*,1,\frac{1}{3}) & d_L^c & \mathcal{D} \\ (1,2,-\frac{1}{2}) & \begin{pmatrix} \nu_L \\ e_L^- \end{pmatrix} & \mathcal{L} \\ (1,1,1) & e_L^+ & \mathcal{E} \\ (1,1,0) & \nu_L^c & \mathcal{N} \end{array}$$

## MSSM INTERACTIONS: THE GOOD

#### Yukawa's

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 $\begin{array}{ll} \alpha,\beta: & \text{family labels} \\ \psi^{\mathcal{Q}}: & \text{fermion field} \\ \mathcal{Q}: & \text{corresponding superfield} \end{array}$ 

## MSSM INTERACTIONS: THE GOOD

#### Yukawa's

$$\mathcal{L}_{Y} = g_{\mathcal{U}}^{\alpha\beta} \bar{\psi}_{L}^{\mathcal{Q},\alpha} [\mathbf{C}\phi^{*}] \psi_{R}^{\mathcal{U},\beta} + g_{\mathcal{N}}^{\alpha\beta} \bar{\psi}_{L}^{\mathcal{L},\alpha} [\mathbf{C}\phi^{*}] \psi_{R}^{\mathcal{N},\beta} + g_{\mathcal{D}}^{\alpha\beta} \bar{\psi}_{L}^{\mathcal{Q},\alpha} \phi \psi_{R}^{\mathcal{D},\beta} + g_{\mathcal{E}}^{\alpha\beta} \bar{\psi}_{L}^{\mathcal{L},\alpha} \phi \psi_{R}^{\mathcal{E},\beta} + \text{c.c.}$$

Note: the Higgs field  $\phi$  is needed with and without conjugate

Hence both H1 and H2 are needed to get all required Yukawa's

 $g_{\mathcal{D}}\mathcal{Q}H_1\bar{\mathcal{D}} + g_{\mathcal{E}}\mathcal{L}H_1\bar{\mathcal{E}}$ 

 $g_{\mathcal{U}}\mathcal{Q}H_{2}\bar{\mathcal{U}}+g_{\mathcal{N}}\mathcal{L}H_{2}\bar{\mathcal{N}}$ 

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# MSSM INTERACTIONS: THE BAD

Some undesirable terms are also allowed:

 $QL\bar{D}; LL\bar{E}; \bar{U}\bar{U}\bar{D}; LH_2$ 

Violate Baryon number and/or Lepton number Not allowed in SM because of odd number of fermions



Disastrous unless very small, or sparticles very heavy

# MSSM INTERACTIONS: THE BAD

How to prevent this? Note that

## $QL\bar{D}; LL\bar{E}; \bar{U}\bar{U}\bar{D}; LH_2$

Violate B-L.

Hence we may postulate B-L as an exact symmetry of nature (not possible for B and L separately!)

But this would forbid Majorana neutrino masses!

A less restrictive constraint is R-parity

 $R_p = (-1)^{3(B-L)+2S}$ 

# **R-PARITY**

$$R_p = (-1)^{3(B-L)+2S}$$

S: Spin



#### SM particles: R-parity + Superpartners: R-parity -

Two important consequences



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# MSSM INTERACTIONS: THE UGLY

Another undesirable term

$$\mu H_1 H_2$$

This gives an equal mass to Higgses and Higgsinos  $\mu^2(|h_1|^2+|h_2|^2)$ 

Problems:

 Natural size: M<sub>Planck</sub> or M<sub>GUT</sub> (but technically natural)
 Positive definite: No "Mexican hat" (but susy still unbroken)

"The µ problem"

At low energy, susy is broken. At high energy it can be:

At low energy, susy is broken. At high energy it can be:

A fundamental symmetry of nature

Not a fundamental symmetry of nature.

At low energy, susy is broken. At high energy it can be:

A fundamental symmetry of nature
 Then it must be symmetry of gravity as well:
 Supergravity.
 This symmetry is not a symmetry of the vacuum:
 Spontaneous breaking.

Not a fundamental symmetry of nature.

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 Then it must be symmetry of gravity as well:
 Supergravity.
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 Spontaneous breaking.

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Accidental low-energy symmetry: Explicit breaking

## SONTANEOUSLY BROKEN SUPERGRAVITY

Vacuum not invariant:

$$Q_{\alpha}|0> \neq 0$$

This would lead to a massless Goldstone particle in the spectrum; Because susy is fermionic this particle is fermion: The Goldstino.

Supergravity implies that supersymmetry is a local symmetry. The gauge boson is a spin-3/2 particle: The Gravitino.

Symmetry breaking now leads to a Higgs-like mechanism: The Gravitino eats the Goldstino and become massive

# SOFT SUSY BREAKING

Parametrization of broken supersymmetry (independent of how it is broken).

*Soft supersymmetry breaking term:* term in the action that breaks susy, but not its good properties at high energies: "non-renormalization theorems".

In particular, these terms respect the absence of quadratic divergencies for scalar masses: The hierarchy problem is solved in the technical sense.

#### **ALLOWED SOFT BREAKING TERMS**

Allowed:

 $m_{ij}\varphi_i\varphi_j^*$ ;  $\alpha_{ij}\varphi_i\varphi_j + c.c$ ;  $\beta_{ijk}\varphi_i\varphi_j\varphi_k + c.c$ ;  $\mu(\lambda\lambda + \bar{\lambda}\bar{\lambda})$ 

 $\lambda$  can be any gaugino in the theory  $\varphi_i$  can be any scalar in the theory

All superpartners plus the Higgs can get a mass after susybreaking, but before  $SU(3) \times SU(2) \times U(1)$  breaking.

Not allowed: Fourth order scalar terms.

### MSSM (2) SOFT BREAKING PARAMETERS

 $\mathcal{L}_{soft} =$ 

$$\begin{split} &-\sum_{i} m_{i}^{2} |\varphi_{i}|^{2} \\ &-\frac{1}{2} \sum_{a} M_{a} \bar{\lambda}_{a} \lambda_{a} \\ &+ [m_{12}^{2} h_{1} h_{2} + c.c] \\ &+ [g_{\mathcal{U}} A_{\mathcal{U}} \varphi_{\mathcal{Q}} \varphi_{\bar{\mathcal{U}}} h_{2} + g_{\mathcal{N}} A_{\mathcal{N}} \varphi_{\mathcal{L}} \varphi_{\bar{\mathcal{N}}} h_{2} + c.c] \\ &+ [g_{\mathcal{D}} A_{\mathcal{D}} \varphi_{\mathcal{Q}} \varphi_{\bar{\mathcal{D}}} h_{1} + g_{\mathcal{L}} A_{\mathcal{L}} \varphi_{\mathcal{L}} \varphi_{\bar{\ell}} h_{1} + c.c] \end{split}$$

### MSSM (2) SOFT BREAKING PARAMETERS

 $\mathcal{L}_{\mathrm{soft}} =$ 

 $-\sum_{i} m_{i}^{2} |\varphi_{i}|^{2} \qquad (5 \times 9) + 2$   $-\frac{1}{2} \sum_{a} M_{a} \bar{\lambda}_{a} \lambda_{a} \qquad 3$   $+[m_{12}^{2}h_{1}h_{2} + c.c] \qquad 1$   $+[g_{\mathcal{U}}A_{\mathcal{U}}\varphi_{\mathcal{Q}}\varphi_{\bar{\mathcal{U}}}h_{2} + g_{\mathcal{N}}A_{\mathcal{N}}\varphi_{\mathcal{L}}\varphi_{\bar{\mathcal{N}}}h_{2} + c.c]$   $+[g_{\mathcal{D}}A_{\mathcal{D}}\varphi_{\mathcal{Q}}\varphi_{\bar{\mathcal{D}}}h_{1} + g_{\mathcal{L}}A_{\mathcal{L}}\varphi_{\mathcal{L}}\varphi_{\bar{\mathcal{I}}}h_{1} + c.c] \qquad 54$ 

Lots of additional parameters(\*)

105 (+19 SM)

(\*) Ignoring neutrino masses

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### MSSM (2) SOFT BREAKING PARAMETERS

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Lots of additional parameters(\*)

But: just a parametrization of an unknown breaking mechanism

(\*) Ignoring neutrino masses

105 (+19 SM)

# CONSTRAINTS



$$\frac{1}{M^2} \left( \frac{\Delta \tilde{m}_{\mathcal{U}}^2}{\tilde{m}_{\mathcal{U}}^2} \right) < 10^{-7} \text{ GeV}^{-2}$$

## **ADDITIONAL ASSUMPTIONS**

To reduce the parameter space often some modelinspired assumptions are made for the parameter values at some high (GUT?) scale.

> $(m_i)^2 = (m_0)^2 \mathbf{1}$  [Universal scalar mass]  $M_a = m_{\frac{1}{2}}$  [Universal gaugino mass]  $A_x = Am_0 \mathbf{1}$  [Universal three-point coupling]

Then there are just 5 additional parameters:

$$u, m_{1/2}, m_0^2, m_{12}^2$$
 and A

The complete Higgs potential is

$$V(h_1, h_2) = \mu_1^2 |h_1|^2 + \mu_2^2 |h_2|^2 - (m_{12}^2 h_1 h_2 + \text{c.c})$$
  
+  $\frac{1}{8} (g_1^2 + g_2^2) (|h_1|^2 - |h_2|^2)^2 + \frac{1}{2} g_2^2 |h_1^{\dagger} h_2|^2$ 

 $h_i$ : Scalar in Higgs superfield  $H_i$  $\mu_i^2 = |\mu|^2 + m_{h_i}^2$ 

 $g_i$ : Gauge coupling  $m_{h_i}^2$  can be negative

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$$D\text{-terms}$$

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Soft breaking

 $g_i$ : Gauge coupling  $m_{h_i}^2$  can be negative

# HIGGS ALIGNMENT

$$H_{1}:(1,2,-\frac{1}{2}) \qquad \langle h_{1}\rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} v_{1} \\ 0 \end{pmatrix}$$
$$H_{2}:(1,2,\frac{1}{2}) \qquad \langle h_{2}\rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v_{2} \end{pmatrix}$$

$$T_3 + Y = 0$$

h1 direction is irrelevant

h<sub>2</sub> direction with respect
to h<sub>1</sub> *is* relevant:
If not exactly aligned,
the photon is massive!



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## HIGGS ALIGNMENT

Parametrize the h2 direction

$$\langle h_2 \rangle = \frac{1}{\sqrt{2}} \left( \begin{array}{c} v_2 \ e^{i\alpha} \sin \gamma \\ v_2 \ e^{i\eta} \cos \gamma \end{array} \right)$$

Then:

$$V(v_1, v_2, \alpha, \eta, \gamma) = \mu_1^2 v_1^2 + \mu_2^2 v_2^2 - 2m_{12}^2 v_1 v_2 \cos \eta \cos \gamma + \frac{1}{8} (g_1^2 + g_2^2) (v_1^2 - v_2^2)^2 + \frac{1}{2} g_2^2 g_2^2 v_1^2 v_2^2 \sin^2 \gamma$$

Minimum:  $\sin \gamma = 0$ 

(not true for general two-Higgs potential)

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**RADIATIVE BREAKING**  $V(h_1, h_2) = \mu_1^2 |h_1|^2 + \mu_2^2 |h_2|^2 - (m_{12}^2 h_1 h_2 + \text{c.c})$  $+ \frac{1}{8} (g_1^2 + g_2^2) (|h_1|^2 - |h_2|^2)^2 + \frac{1}{2} g_2^2 |h_1^{\dagger} h_2|^2$ 

Flat direction in quartic potential:  $h_2 = e^{i\alpha} C h_1^{\dagger}$  $(hCh \equiv h_i \epsilon_{ij} h_j = 0)$ 

Quadratic terms: $(\mu_1^2 + \mu_2^2 - 2m_{12}^2 \cos \alpha)|h_1|^2$ Positivity condition: $\mu_1^2 + \mu_2^2 \ge 2|m_{12}^2|$ Negative determinant: $|m_{12}^2|^2 > \mu_1^2\mu_2^2$ Incompatible if $\mu_1 = \mu_2$ (universal scalar masses)

## **RADIATIVE BREAKING**

 $h_2$ 

At some large scale:

 $m_{h_1} = m_{h_2} = m_0 \rightarrow \mu_1 = \mu_2$ Then both masses start running separately



 $\overline{t}$ 

 $h_2$ 

h<sub>1</sub>,h<sub>2</sub>: Eight real d.o.f. Three are "eaten" by W, Z Hence five massive scalars left (vs. just one in SM).

Electric charges per Higgs:  $2 \times 0$ , +1, -1

Charged Higges: two are eaten, two survive: H<sup>+</sup>, H<sup>-</sup>

Neutral Higges: Common phase is eaten; Scales and relative phase survive: h<sub>0</sub>, H<sub>0</sub>, A<sub>0</sub>

 $V(h_1, h_2) = \mu_1^2 |h_1|^2 + \mu_2^2 |h_2|^2 - (m_{12}^2 h_1 h_2 + \text{c.c})$ +  $\frac{1}{8} (g_1^2 + g_2^2) (|h_1|^2 - |h_2|^2)^2 + \frac{1}{2} g_2^2 |h_1^{\dagger} h_2|^2$ 

Invariant under  $h_i \rightarrow h_i^{\dagger}$ Provided  $m_{12}^2$  is real (can be chosen real w.l.o.g.) This symmetry can be extended to an approximate CP symmetry of the full Lagrangian.

Neutral mass eigenstates are approximate CP eigenstates

$$V(h_1, h_2) = \mu_1^2 |h_1|^2 + \mu_2^2 |h_2|^2 - (m_{12}^2 h_1 h_2 + \text{c.c})$$
  
+  $\frac{1}{8} (g_1^2 + g_2^2) (|h_1|^2 - |h_2|^2)^2 + \frac{1}{2} g_2^2 |h_1^{\dagger} h_2|^2$ 

Relative phase: Odd under CP, massless if  $m_{12}^2 = 0$ 

$$m_{A^0}^2 = \frac{m_{12}^2}{\cos\beta\sin\beta}$$
$$\tan\beta \equiv \frac{v_2}{v_1}$$

Charged Higgs masses

$$M_{\rm w}^2 + m_{A^0}^2$$

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Charged Higgs masses

$$M_{\rm w}^2 + m_{A^0}^2$$

Neutral, CP even

$$m_{H^0,h^0}^2 = \frac{1}{2} \left( m_{A^0}^2 + M_z^2 \pm \sqrt{(m_{A^0}^2 + M_z^2)^2 - 4m_{A^0}^2 M_z^2 \cos^2 2\beta} \right)$$

Lightest one has mass below M<sub>Z</sub>

Loop corrections (due to top quark loops)

$$\Delta M^2 = \frac{3}{8\pi^2} \frac{g_2^2 m_t^4}{M_{\rm w}^2 \sin^2 \beta} \log(1 + \frac{m_0^2}{m_t^2})$$

 $M_{h_0} < 135 \text{ GeV}$ 

### **COUPLING CONSTANT UNIFICATION**



#### **COUPLING CONSTANT UNIFICATION**





**Q** Hierarchy Problem

**Q** Dark Matter

**Q** Coupling Constant Convergence

Sunday, 2 May 2010

**Q** Hierarchy Problem

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Gauginos, sFermions

Gauginos

Gauginos

Sunday, 2 May 2010



**Q** Dark Matter

Coupling Constant Convergence

Gauginos

Gauginos

Gauginos, sr-

JUIIS

(Arkani-Hamed, Dimopoulos, Giudice, Romanino 2004)

# SUSY AT HADRON COLLIDERS



March 25, 1983

#### Experimental Consequences of Supersymmetry

10:00 C. Zachos - Introduction to Supersymmetry P.1

11:00 T. Taylor - Proton Decay. P.12

1:30 S. Dawson - Limits on Superparticles P.24

2:30 R. Huerta - Electron-Positron Collisions P. 53

3:15 Coffee Break

4:00 A. Schellekens - Hadron-Hadron' Colliders P.63



# CONCLUSION