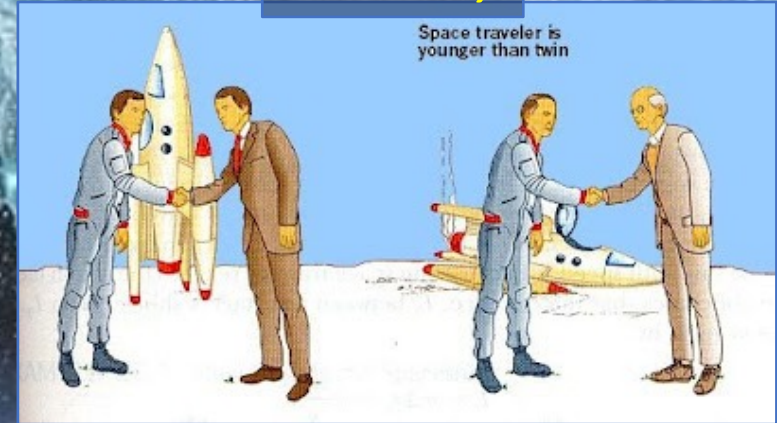


The Relativistic Quantum World

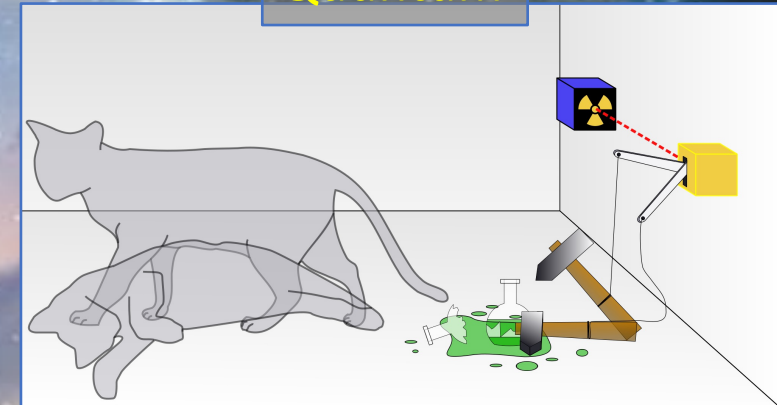
A lecture series on
Relativity Theory and Quantum Mechanics

Marcel Merk
Studium Generale Maastricht
Nov 1 – Nov 29, 2023

Relativity



Quantum



Relativity

Nov. 1:

Lecture 1: The Principle of Relativity and the Speed of Light
Lecture 2: Time Dilation and Lorentz Contraction

Nov. 8:

Lecture 3: The Lorentz Transformation and Paradoxes
Lecture 4: General Relativity and Gravitational Waves

Quantum Mechanics

Nov. 15:

Lecture 5: The Early Quantum Theory
Lecture 6: Feynman's Double Slit Experiment

Nov 22:

Lecture 7: Wheeler's Delayed Choice and Schrodinger's Cat
Lecture 8: Quantum Reality and the EPR Paradox

Standard Model

Nov. 29:

Lecture 9: The Standard Model and Antimatter
Lecture 10: Why is there something rather than nothing?

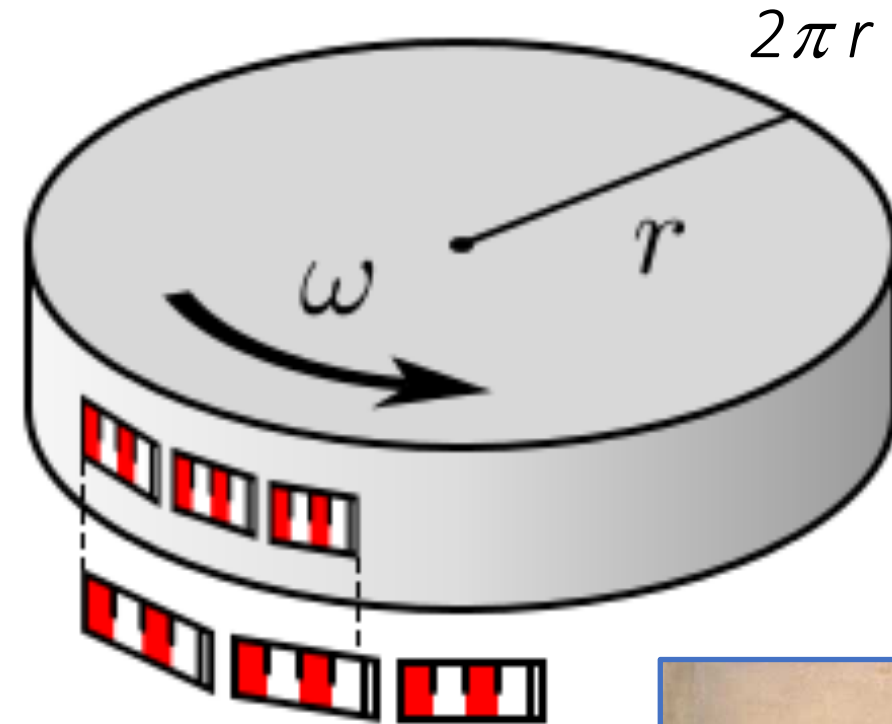
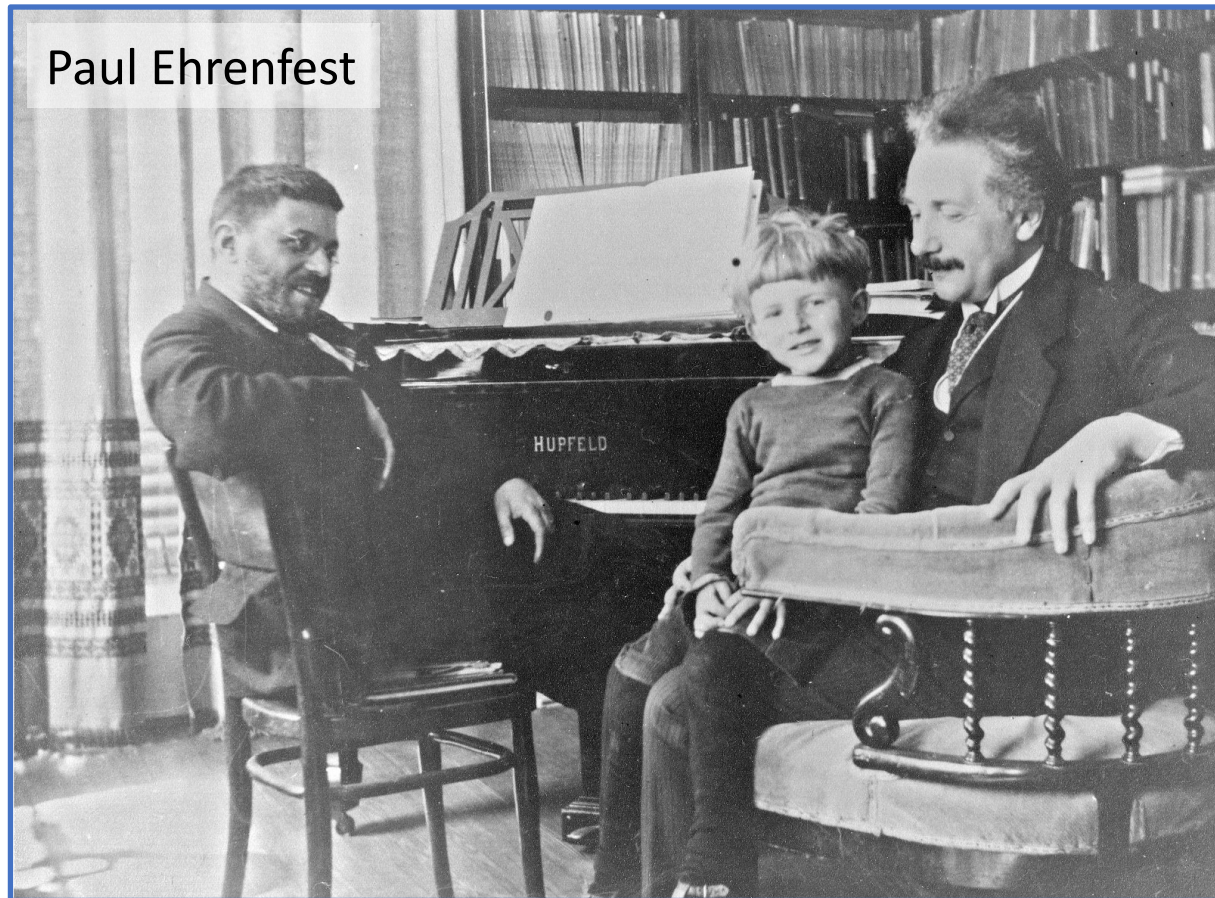
Lecture notes, written for this course, are available: www.nikhef.nl/~i93/Teaching/
Prerequisite for the course: High school level physics & mathematics.

Lecture 4

General Relativity and Gravitational Waves

*“Do not worry about your difficulties in mathematics.
I can assure you mine are still greater.”*
- Albert Einstein

Rotating disk with ruler on the edge:
Circumference: $C = 2 \pi r$



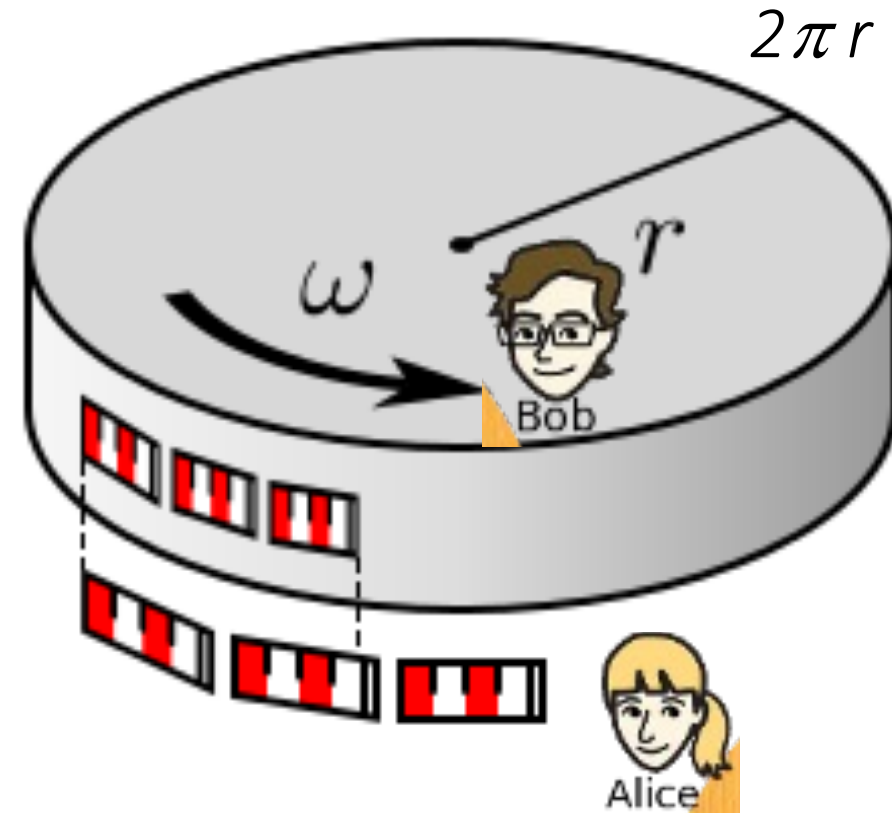
Rotating disk with ruler on the edge:
Circumference: $C = 2 \pi r$

Alice stands next to the disk and sees
rulers on disk Lorentz contracted:
 $C = 2 \pi r / \gamma$

→ Circumference is smaller!

Bob moves on the disk and sees
rulers next to disk contracted:
 $C = 2 \pi r \cdot \gamma$

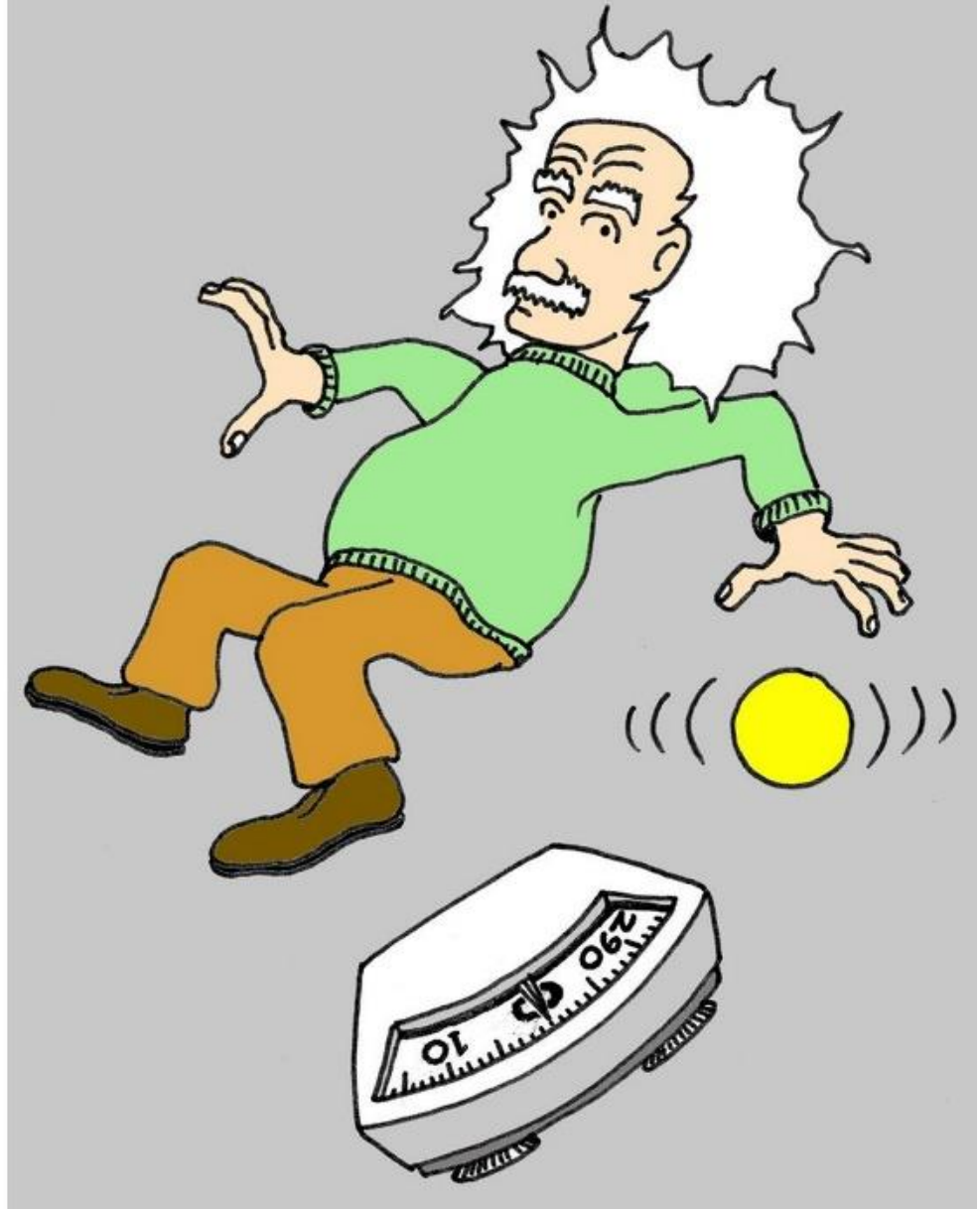
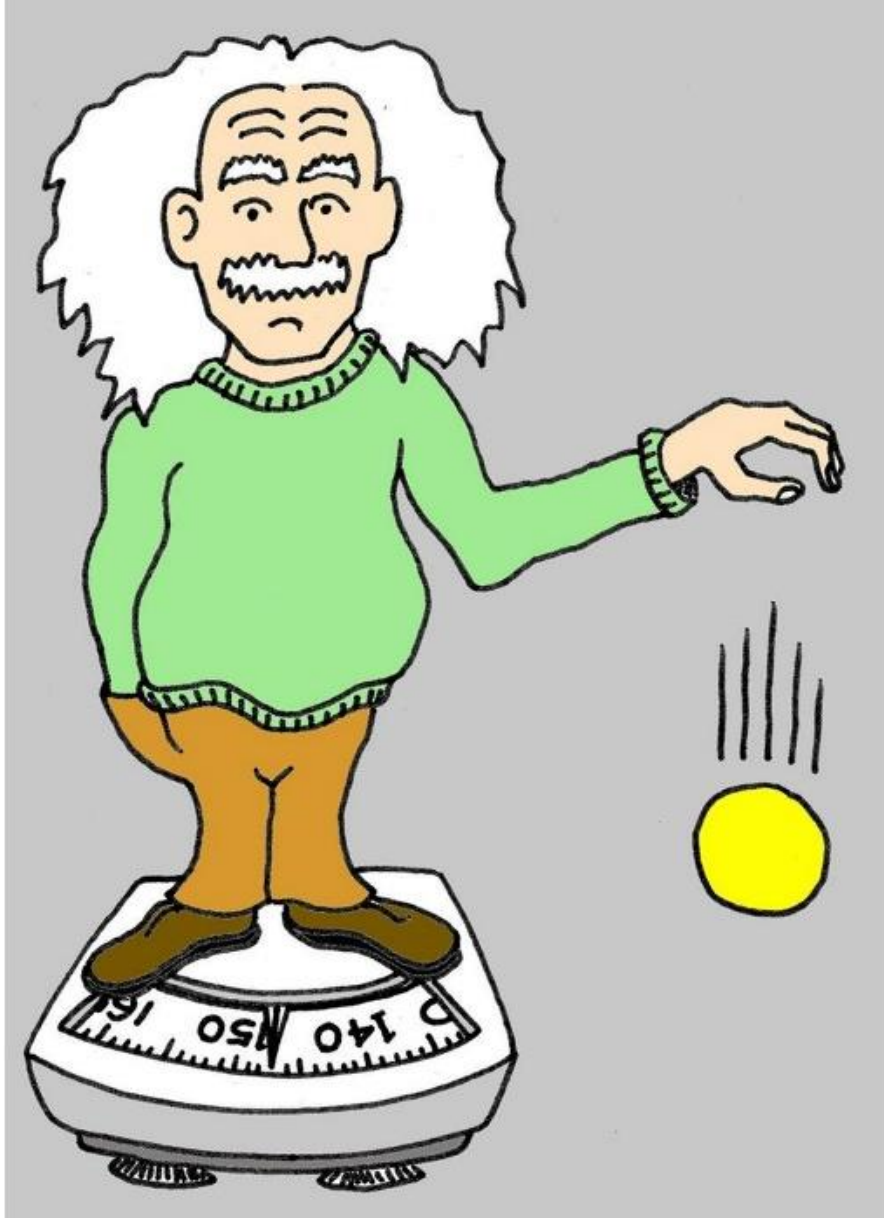
→ Circumference is larger!



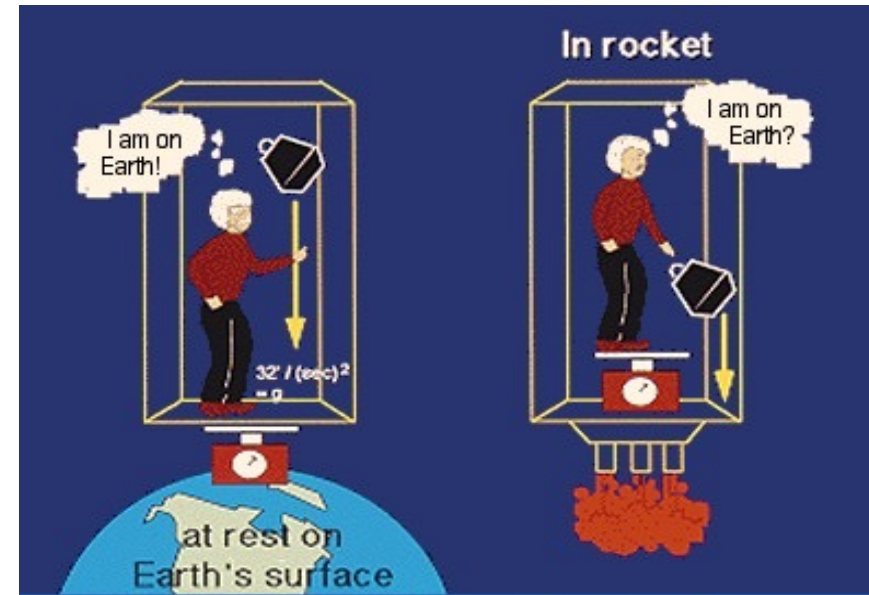
A rotating object is ***not an inertial*** frame:

- Postulate of relativity only worked for ***inertial frames***
- Need to adapt the postulates: special relativity → ***general relativity***

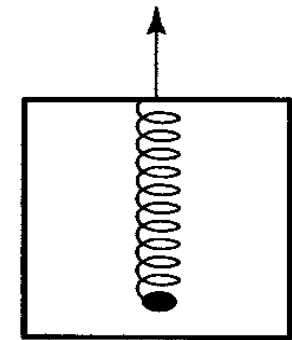
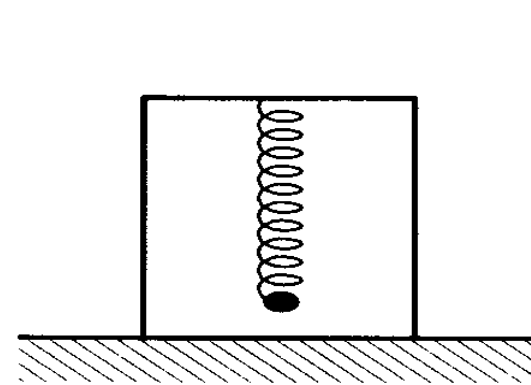
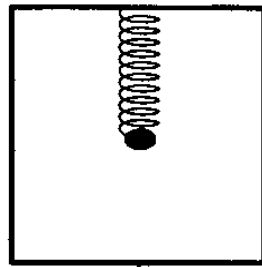
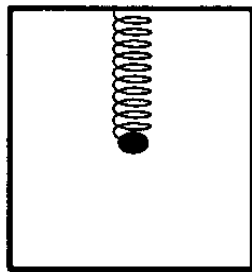
Einstein's "happiest thought"



The Equivalence Principle

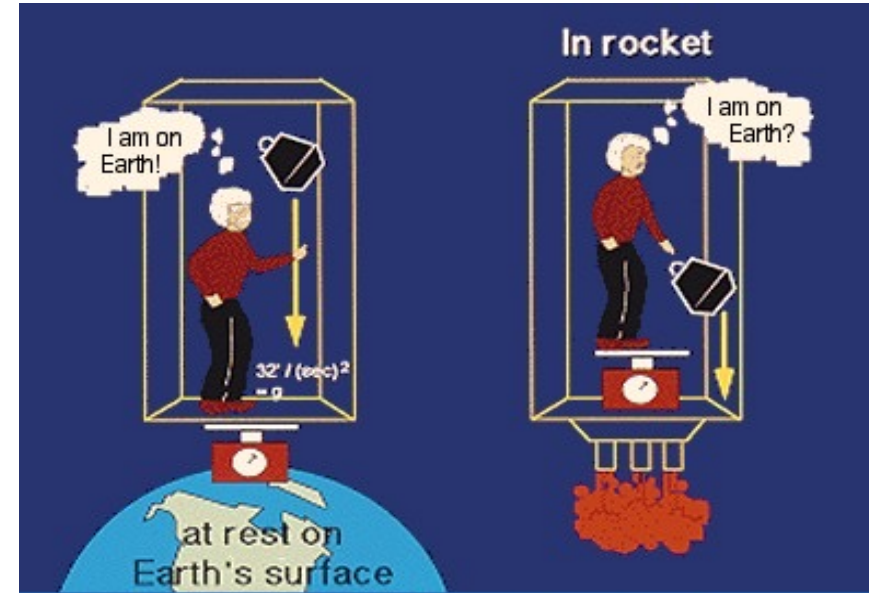


(Inertial Frame)

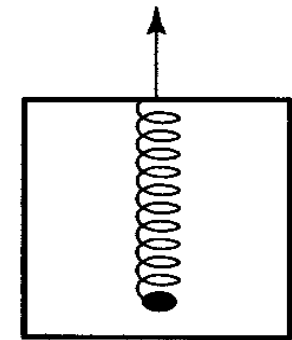
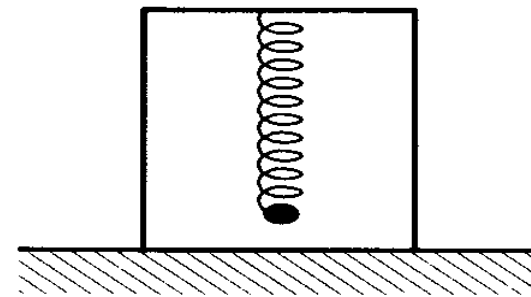
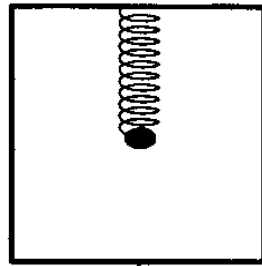
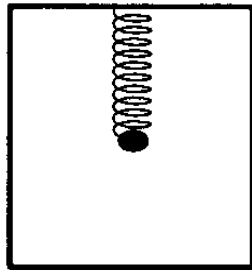


Einstein's "happiest thought": there is **no way** to determine whether you are standing on the earth or accelerating upwards in a rocket in space!

The Equivalence Principle



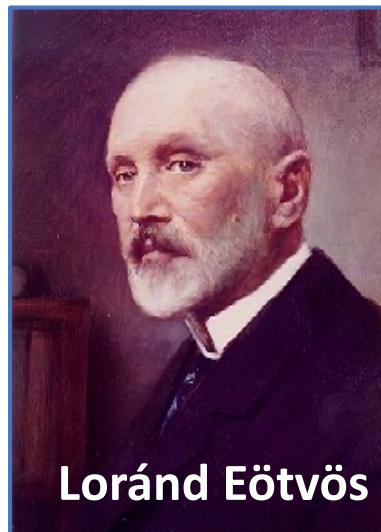
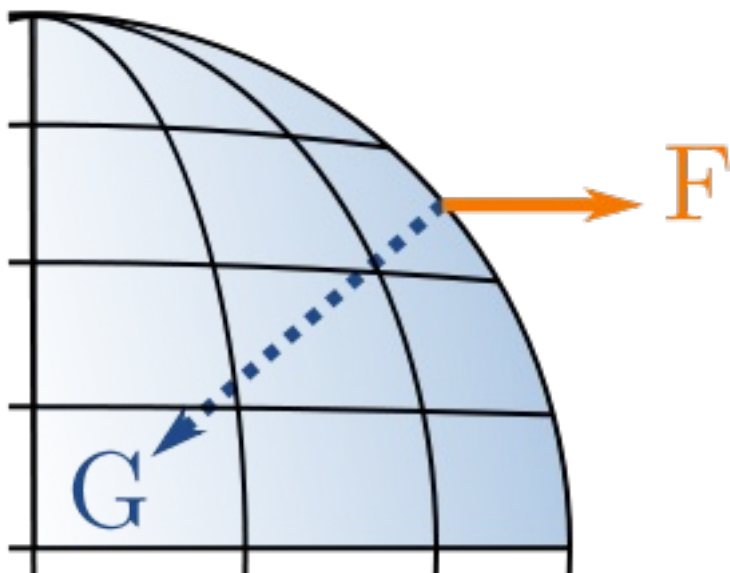
(Inertial Frame)



There is no difference between acceleration force and gravitation.
“Gravitational mass” = “Inertial mass”

The Eötvös Experiment

Direction of gravity and centrifugal force on earth

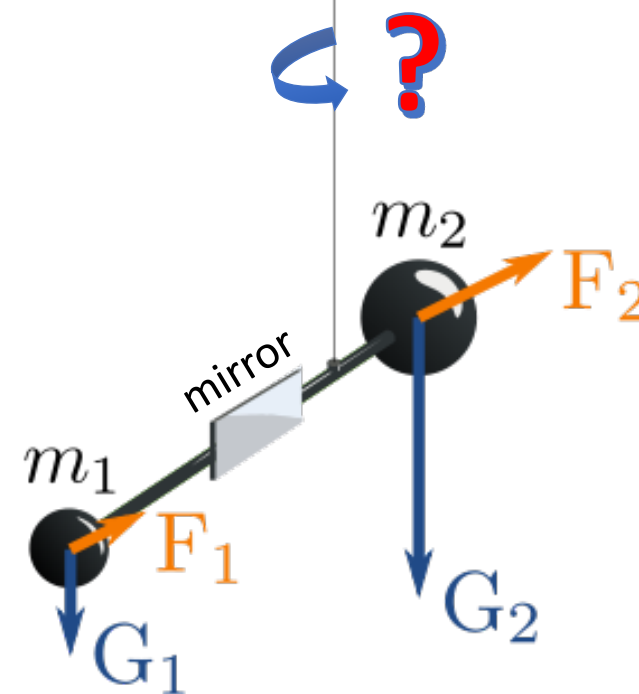


Loránd Eötvös

$$F = m_i a \text{ with } a = \frac{v^2}{R}$$
$$G = m_g g \text{ with } g = \frac{GM_{\oplus}}{R^2}$$

→ $m_i = m_g$

Small (m_1) and big (m_2) mass on a rod suspended by a thin fiber



Gravity force **G** depends on Newton's law of gravity: *gravitational mass*

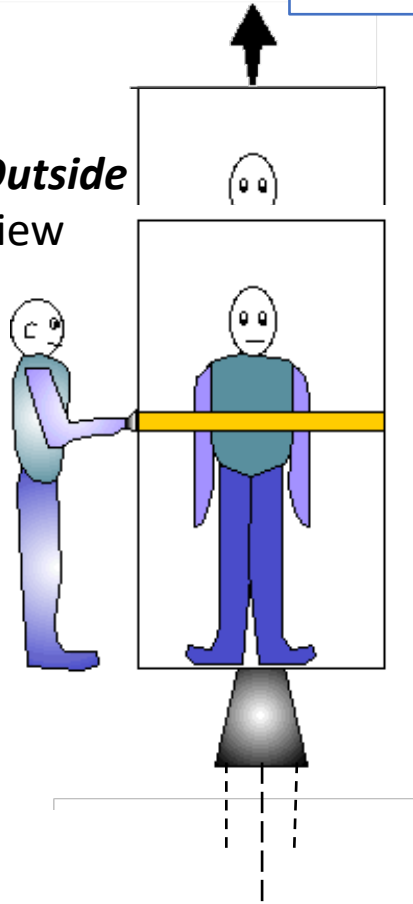
Centrifugal force **F** depends on Newton's law of motion inertial mass: *inertial mass*

The system did **not** rotate. → $F_1/F_2 = G_1/G_2$

→ Experimental proof that indeed gravitational mass is equivalent to inertial mass.

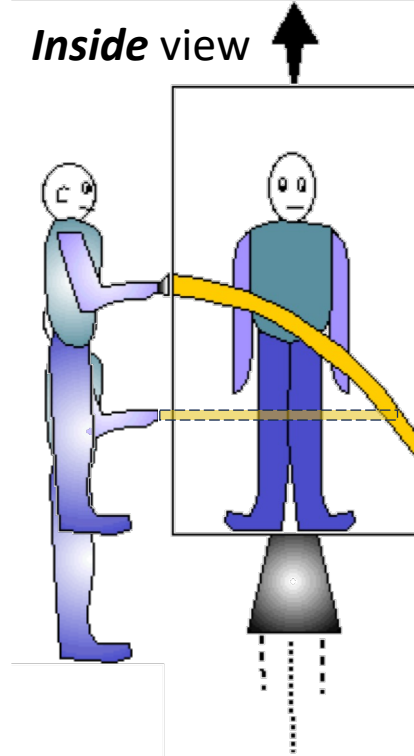
Consequence of acceleration and gravity being equivalent

Outside view

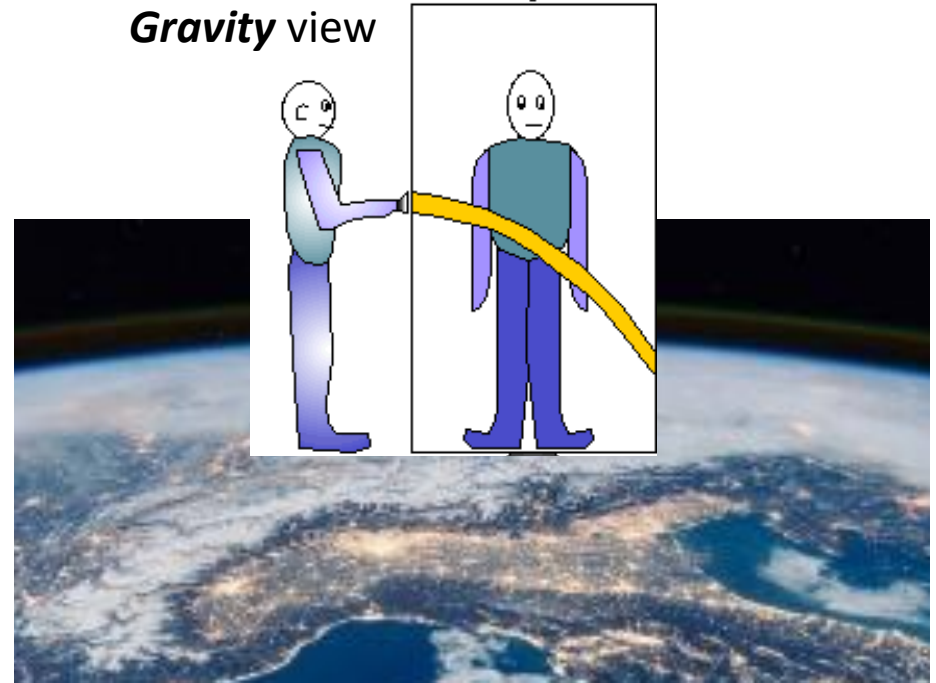


A: Lightbeam in accelerating rocket

Inside view



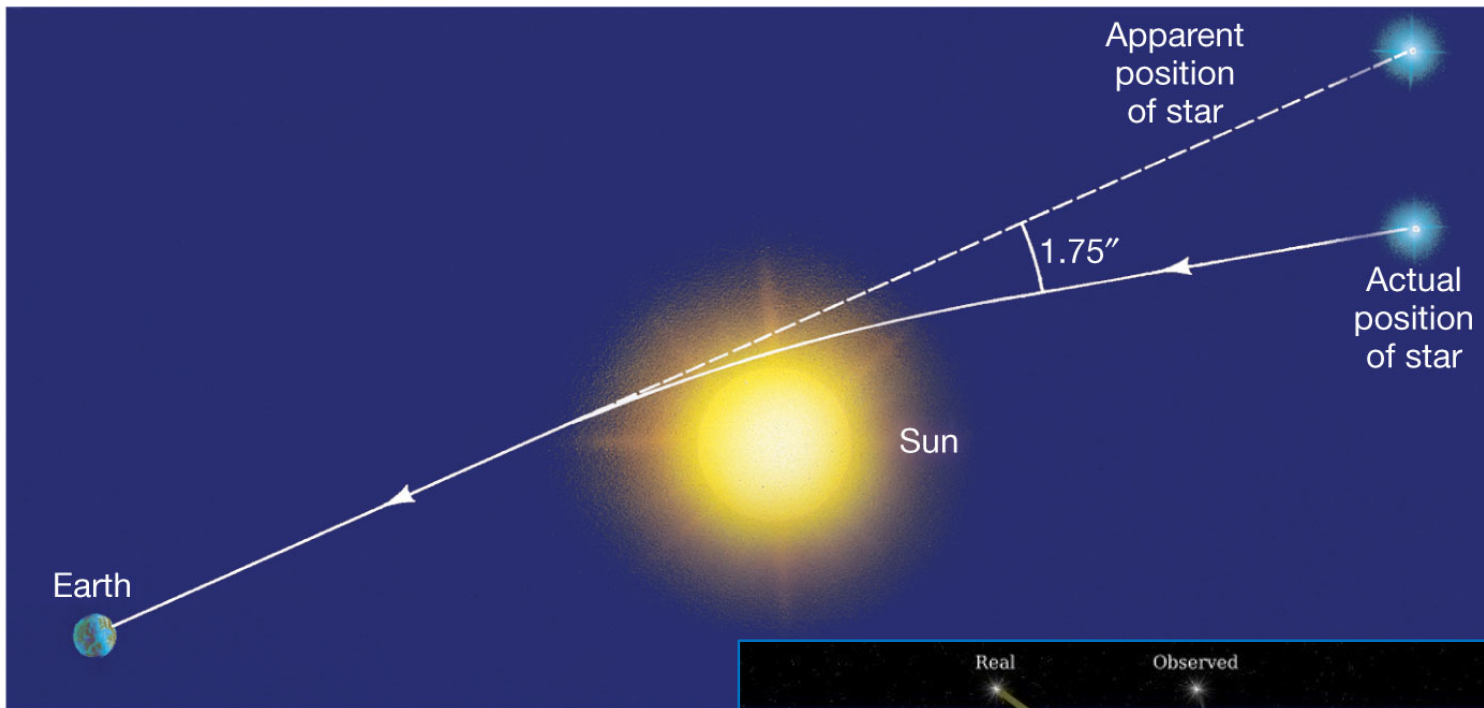
Gravity view



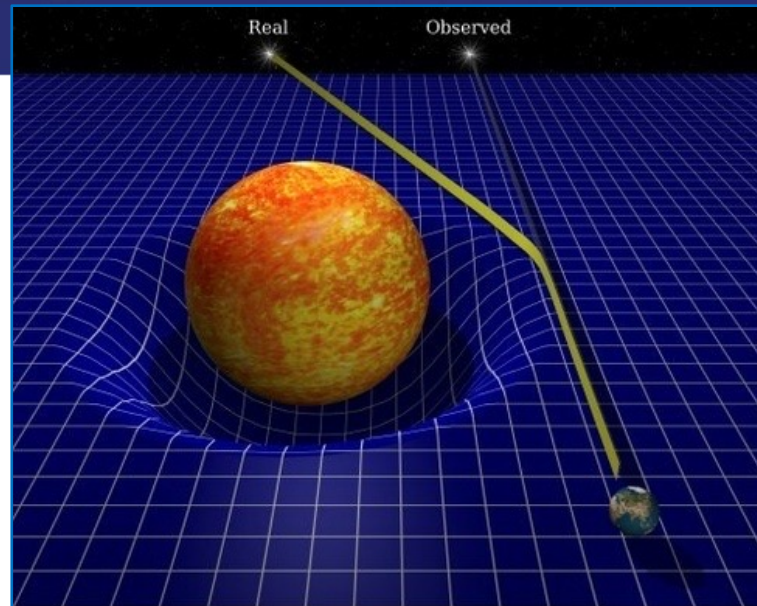
B: Lightbeam in gravitational field

Prediction of Einstein: light beam bends under gravity!

Bending of light in gravitation field of the Sun

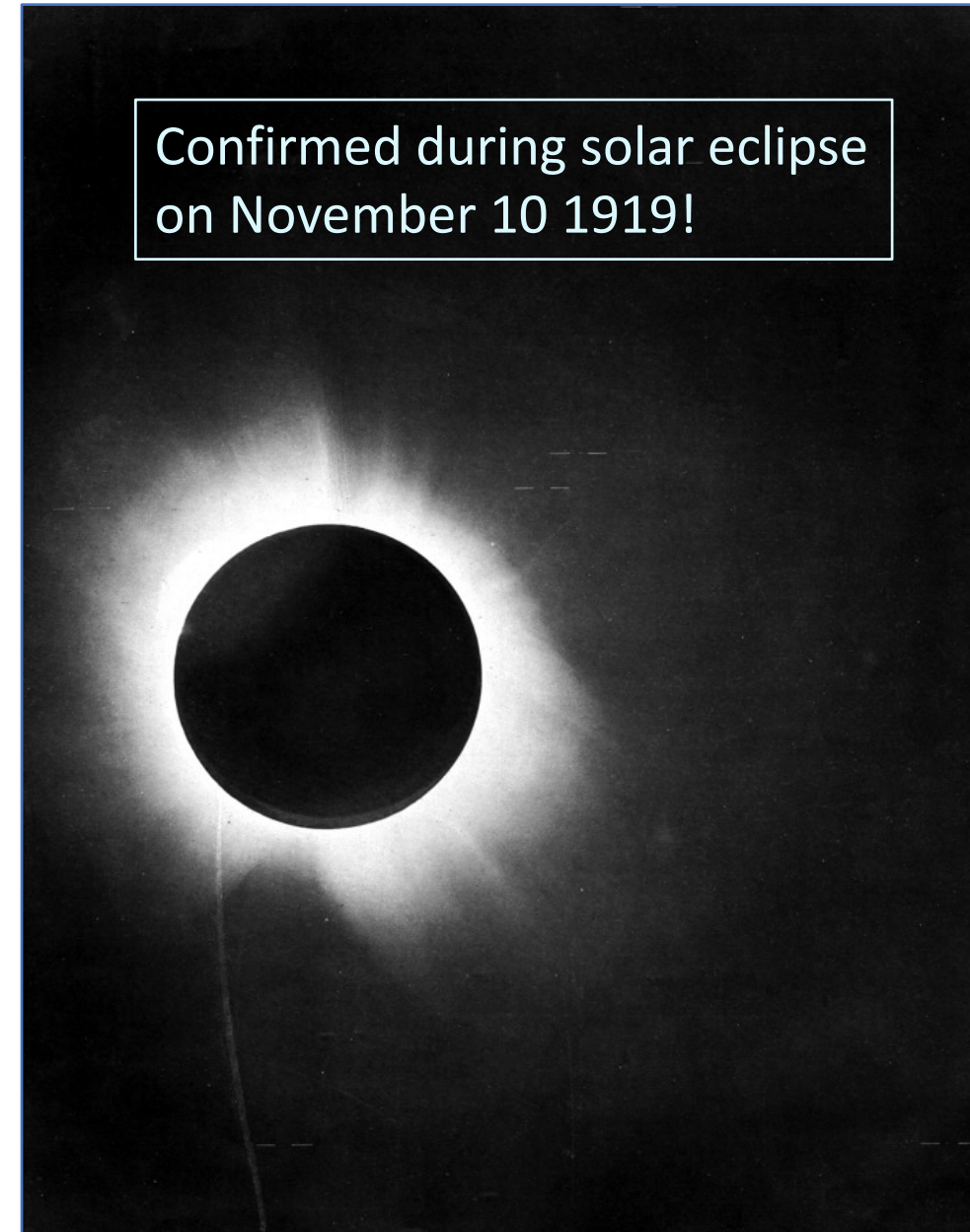


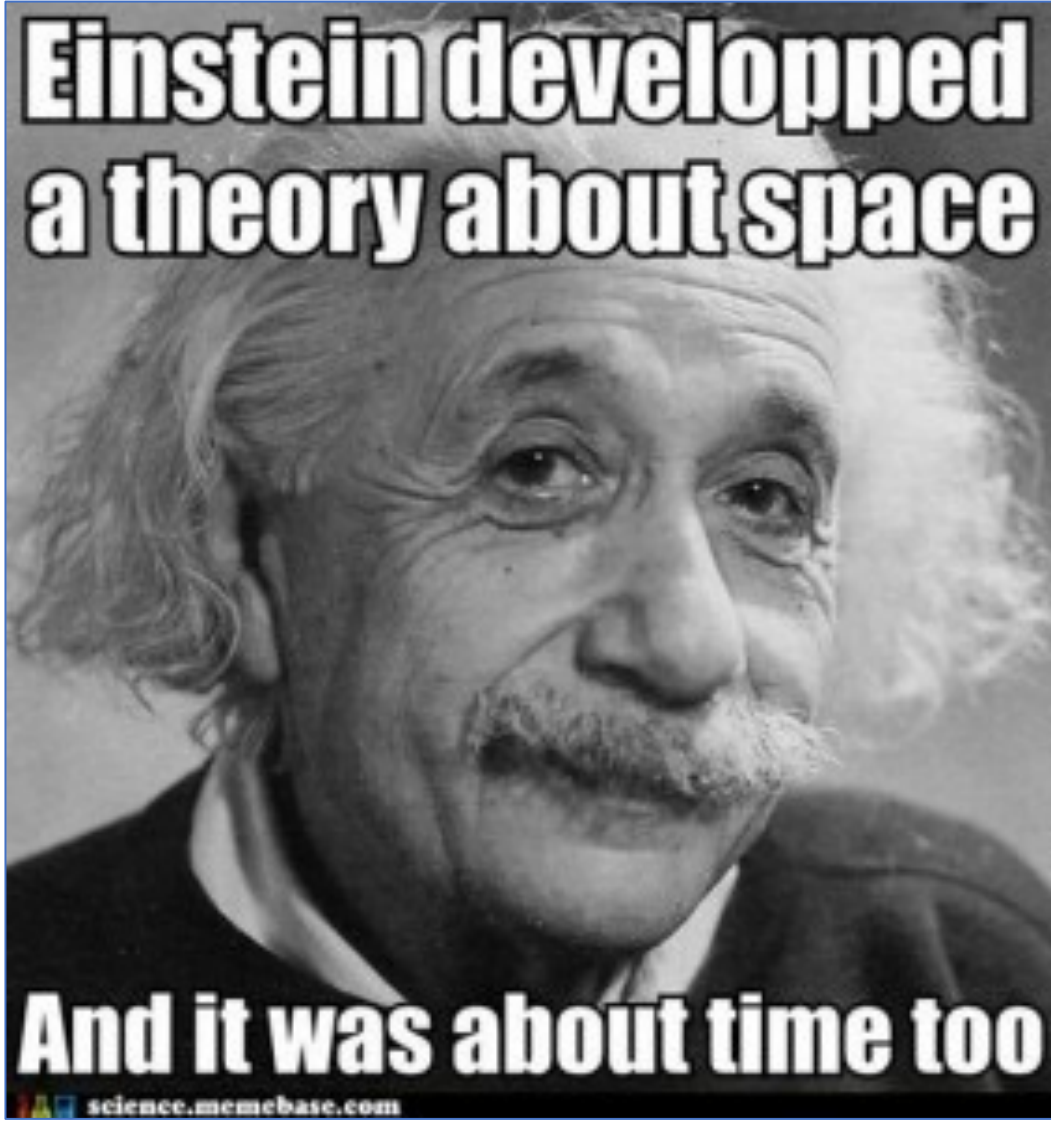
© 2011 Pearson Education, Inc.



Light wants to go straight but space is curved!

Confirmed during solar eclipse on November 10 1919!



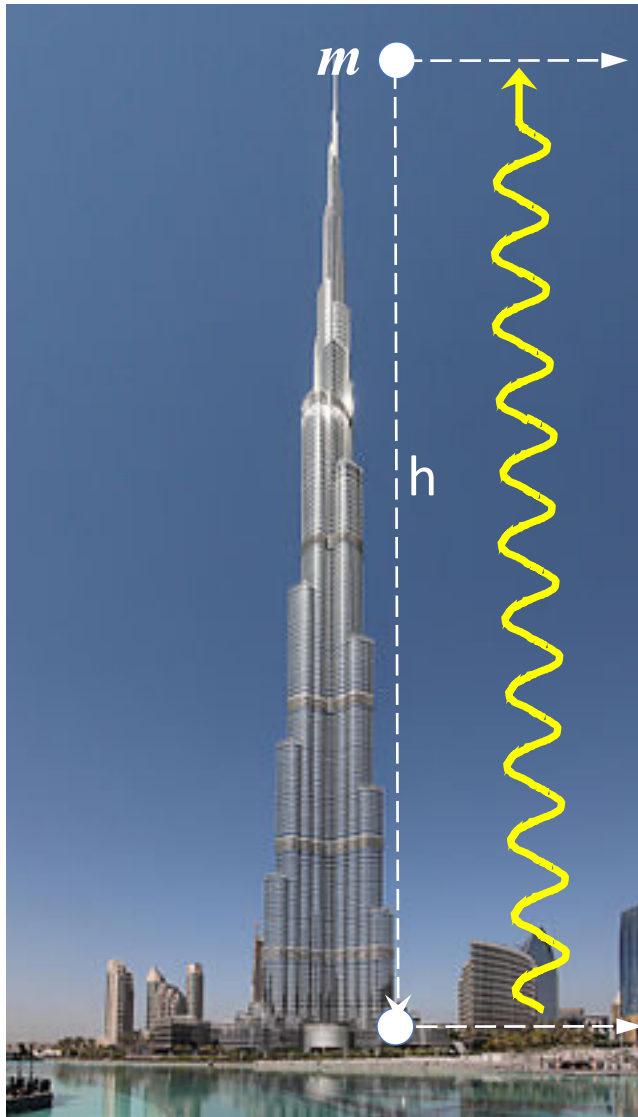


**Einstein developed
a theory about space**

And it was about time too

Einstein's next thought experiment on light

Particle with mass m falling from tower:



$$E = mc^2 \quad \Rightarrow \quad E' > E$$

From quantum mechanics we know that the energy of light is related to frequency (and wavelength): $E = hf = hc/\lambda$

$$E' > E ??$$

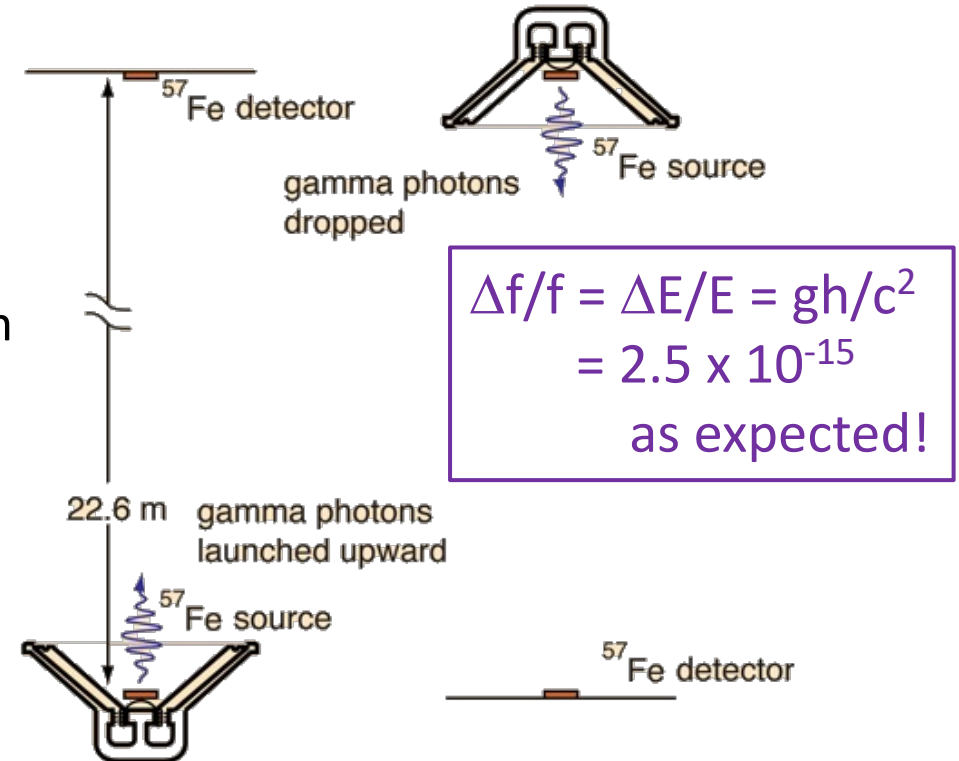
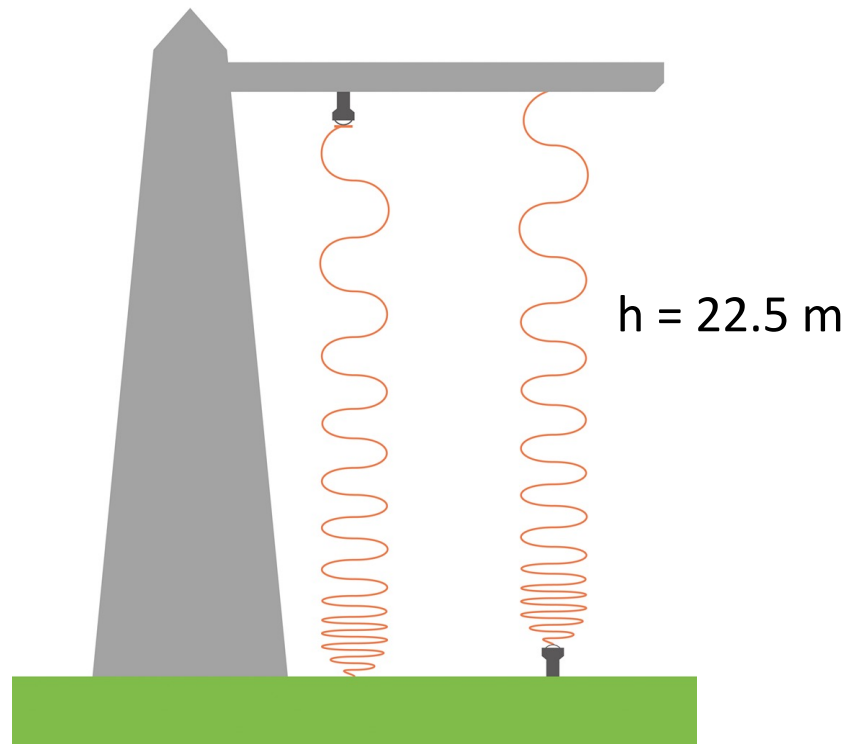
Perpetuum mobile? \rightarrow No!

Photon **loses energy** gh/c^2 as it travels up the gravitational field!
 \rightarrow Wavelength red-shift

$$E' = mc^2 + \frac{1}{2}mv^2 = mc^2 + mgh \quad (E_{kin} = E_{pot})$$
$$= mc^2 (1 + gh/c^2) \quad \Rightarrow \quad E' = hf'$$

The Harvard Tower Experiment

Harvard Tower Experiment (Pound-Rebka)
at Jefferson lab in Harvard (1960):
Measure red-shift of photons in earth
gravitational field.

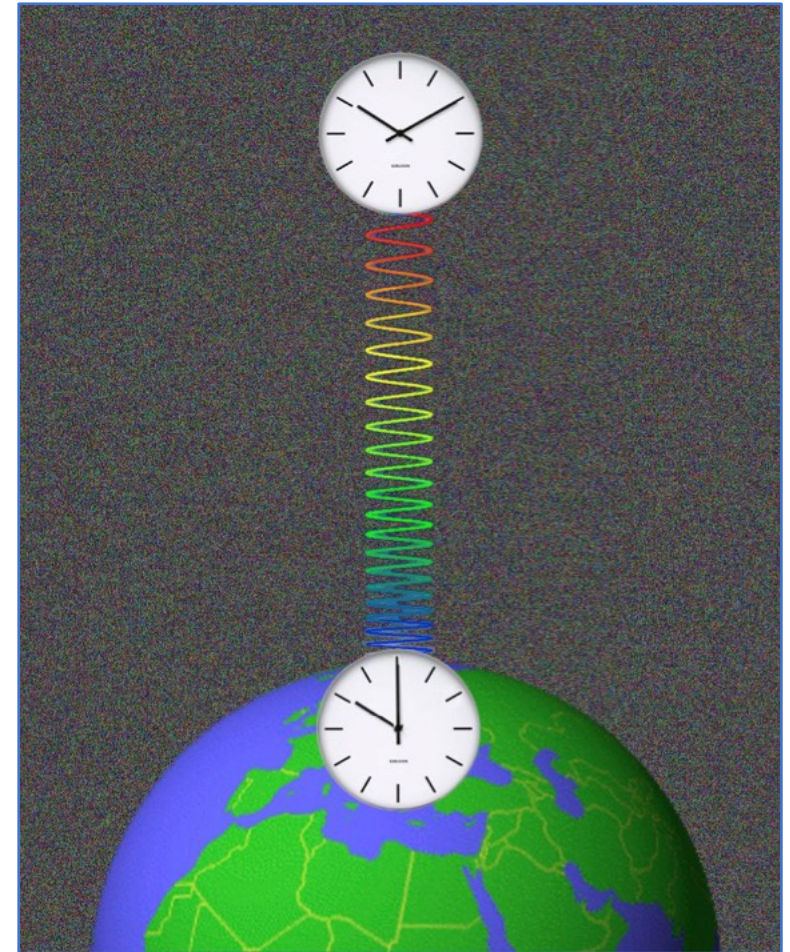
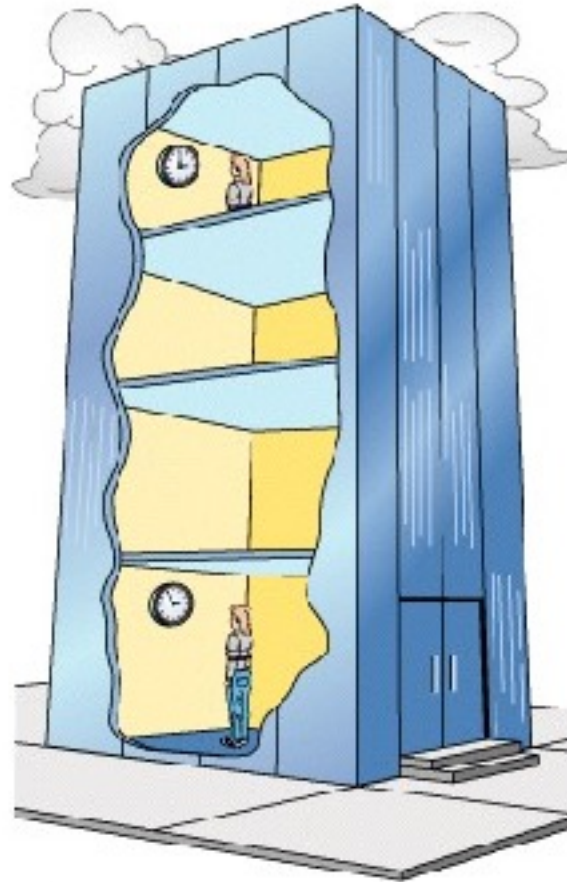
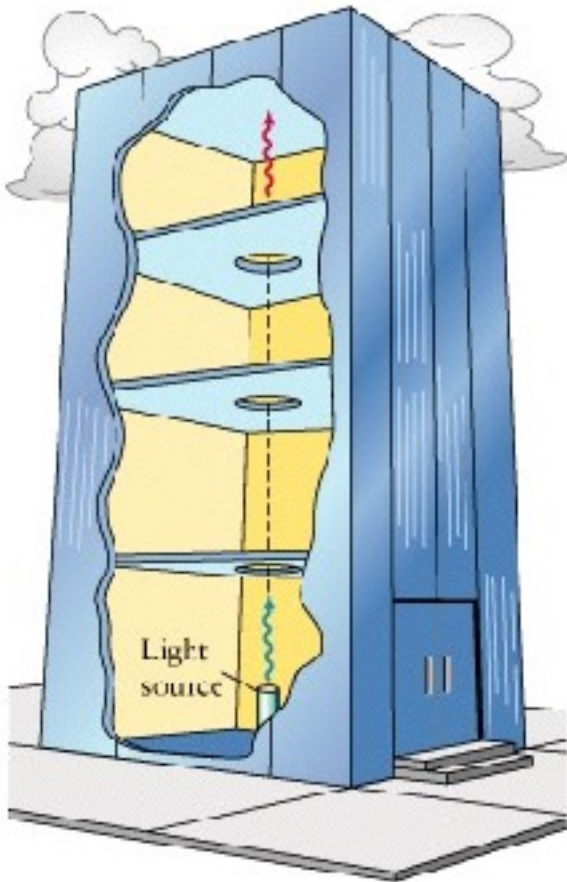


The photon loses energy as it climbs the gravitational field.

Longer wavelength

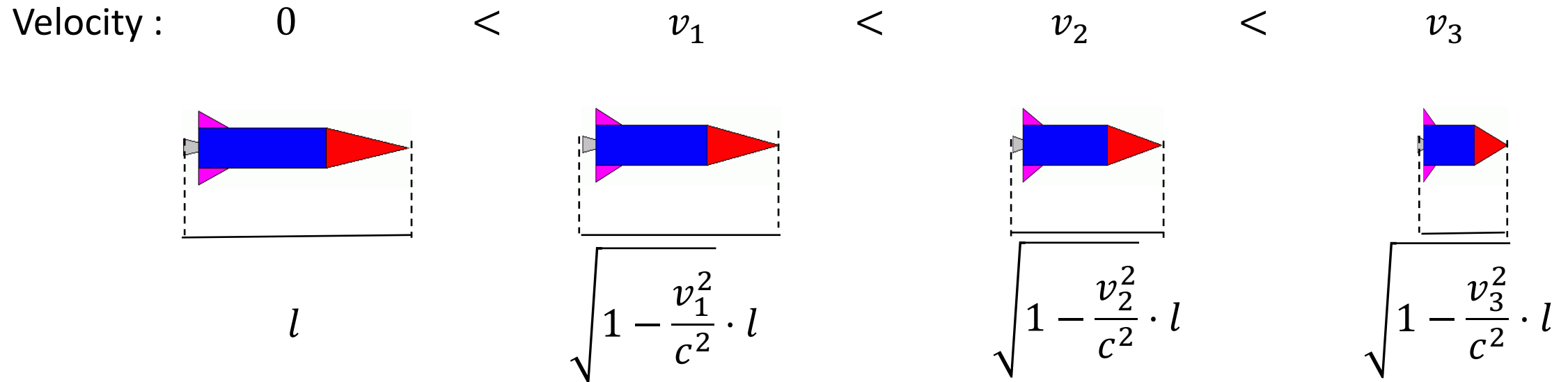
$$c = \lambda f$$

Lower frequency



→ Time ticks faster at higher altitude.

From special relativity we know that space contracts at high velocity



Space is seen to shrink further and further with increasing velocity!

$$1/\gamma = \sqrt{1 - \frac{v^2}{c^2}}$$

$$E_{pot} = G \frac{mM_{\oplus}}{R}$$

↓

$$E_{kin} = E_{pot}$$

$$\frac{1}{2}mv^2 = G \frac{mM_{\oplus}}{R}$$

↓

$$v^2 = 2 \frac{GM_{\oplus}}{R}$$

↓

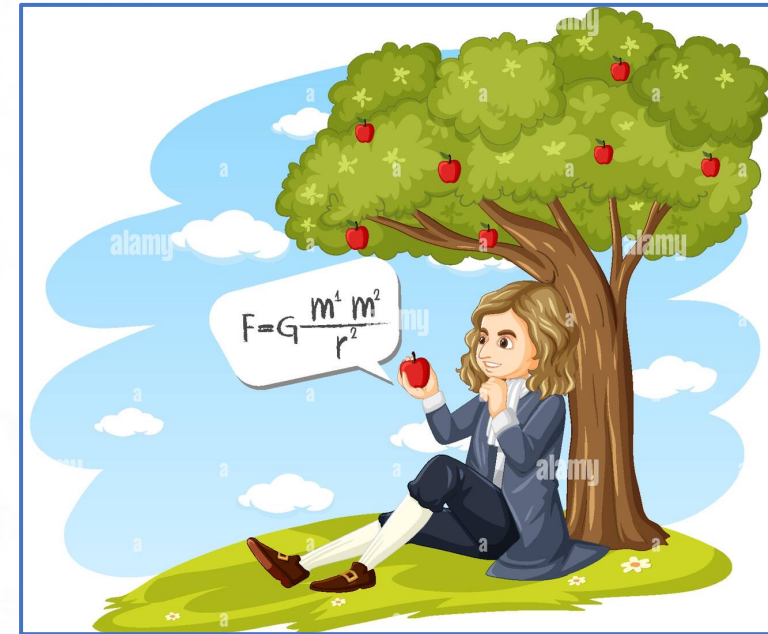
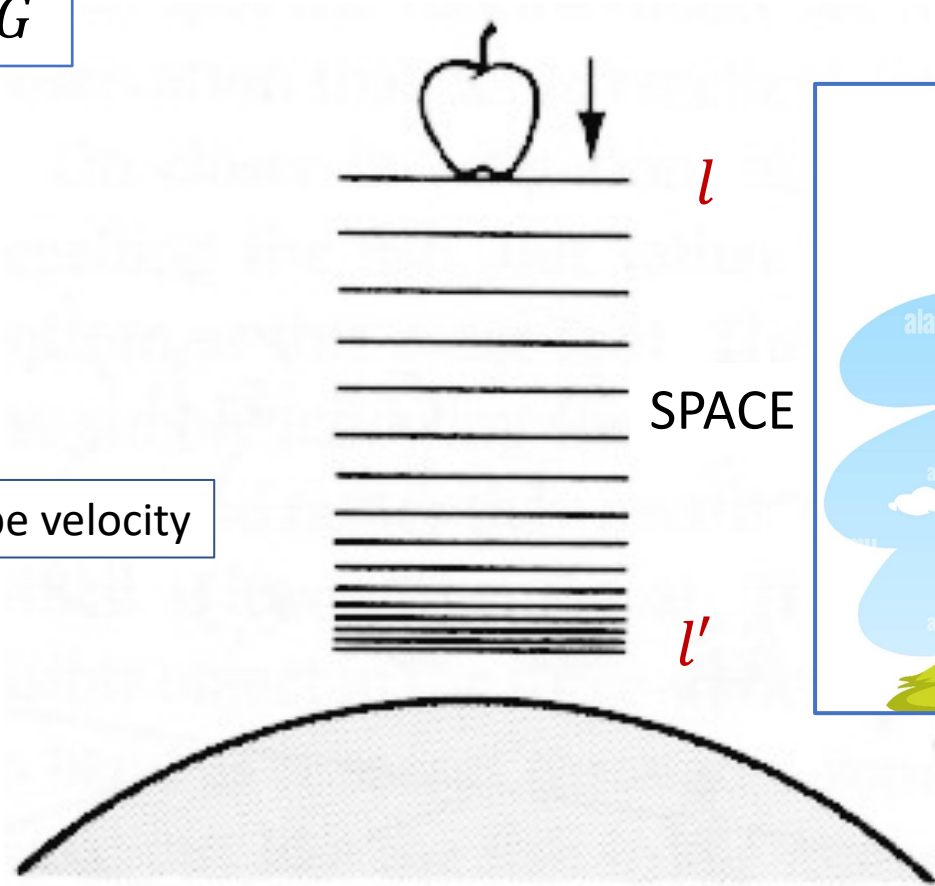
$$\sqrt{1 - \frac{v^2}{c^2}} = \sqrt{1 - 2 \frac{GM_{\oplus}}{Rc^2}}$$

Newton's Constant G

v = Escape velocity

Lorentz factor

Falling apple:



Space shrinkage ("curvature"):

$$l' = \sqrt{1 - 2 \frac{GM_{\oplus}}{Rc^2}} \cdot l$$

$$E_{pot} = G \frac{mM_{\oplus}}{R}$$

↓

$$E_{kin} = E_{pot}$$

$$\frac{1}{2}mv^2 = G \frac{mM_{\oplus}}{R}$$

↓

$$v^2 = 2 \frac{GM_{\oplus}}{R}$$

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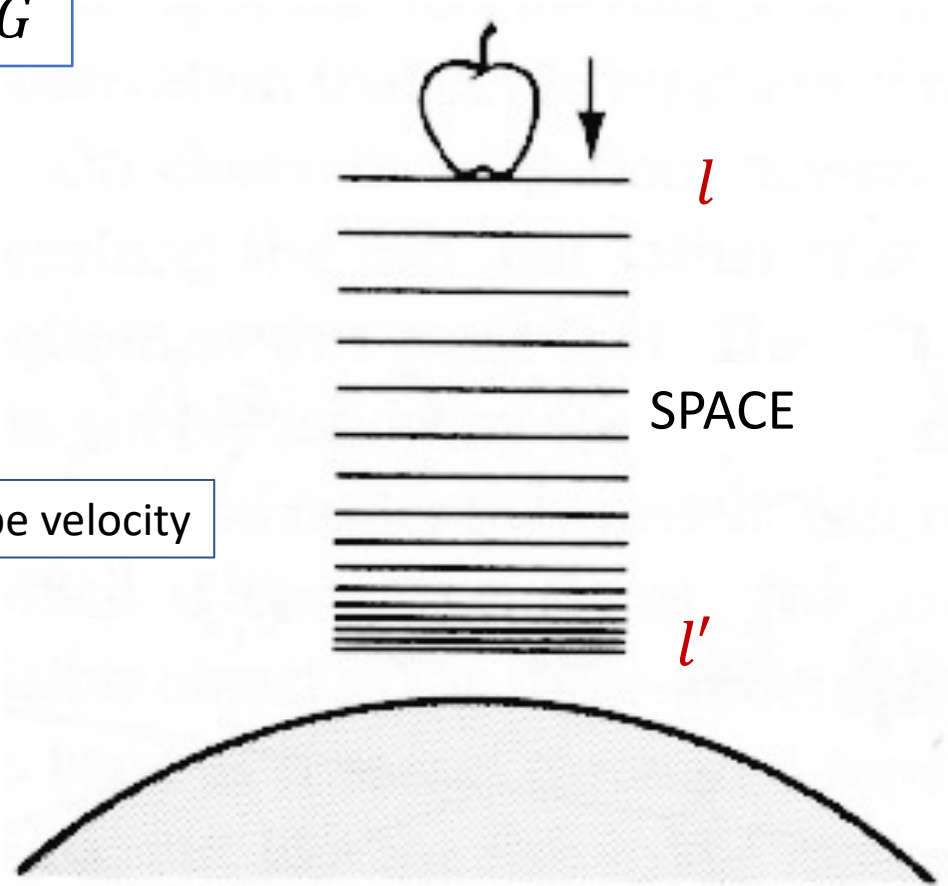
$$\sqrt{1 - \frac{v^2}{c^2}} = \sqrt{1 - 2 \frac{GM_{\oplus}}{Rc^2}}$$

Lorentz factor

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Falling apple:



Compare to accelerating rocket:



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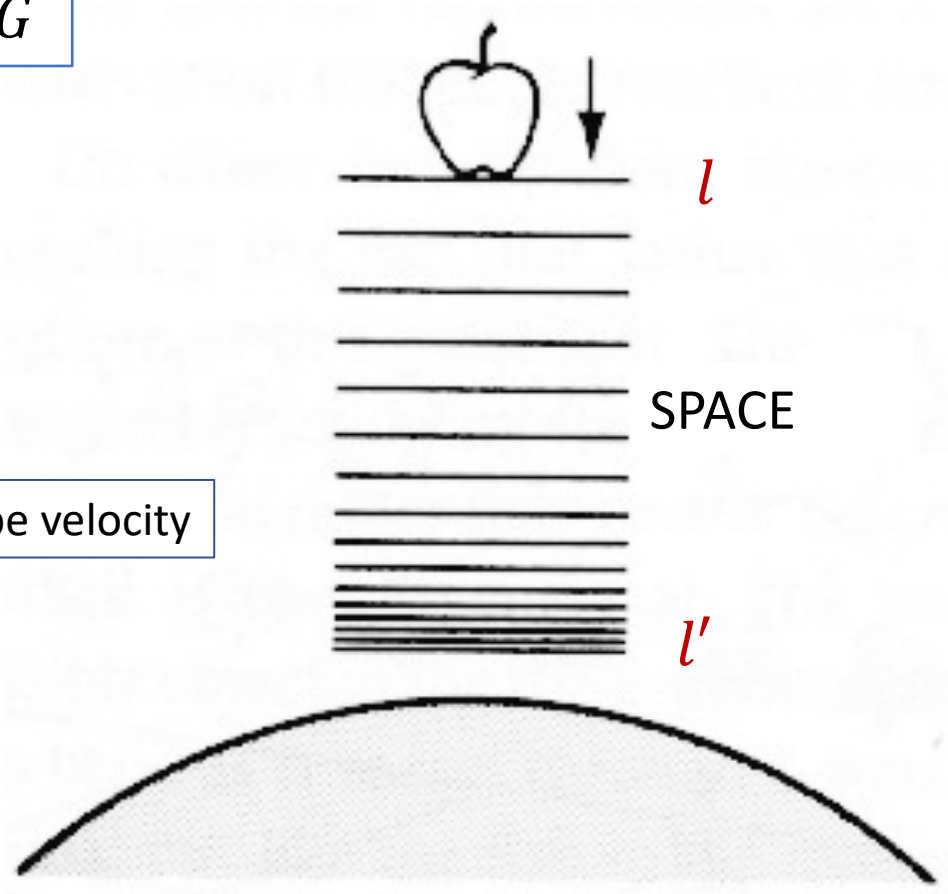
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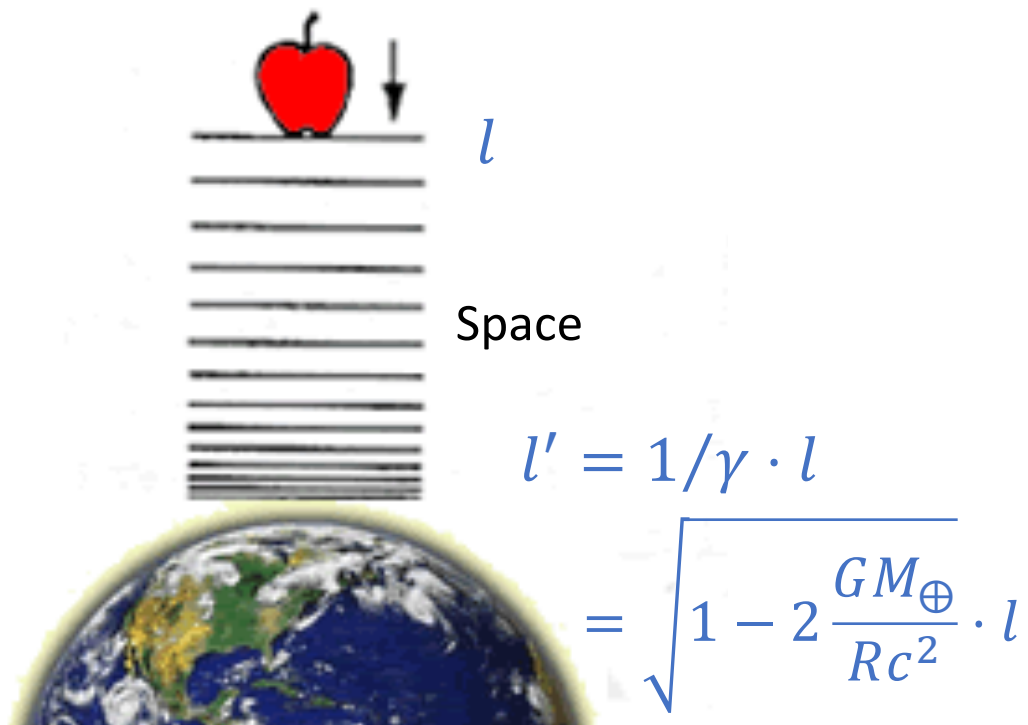
Compare to accelerating rocket:



Space shrinkage ("curvature"):

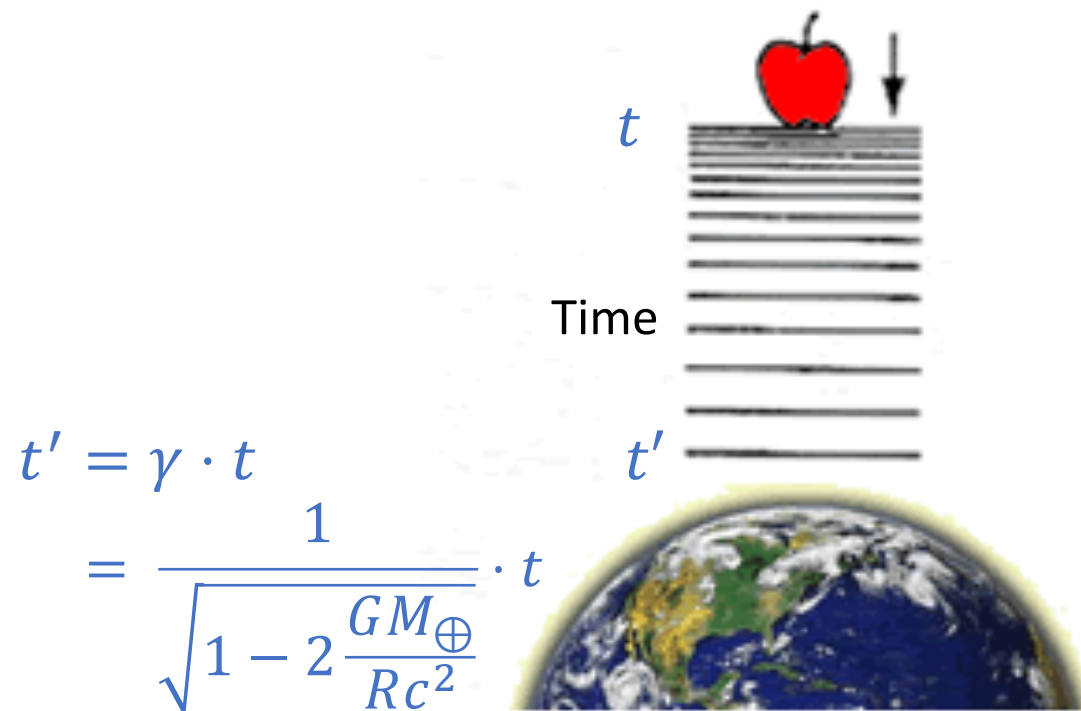
$$l' = \sqrt{1 - 2 \frac{GM_{\oplus}}{Rc^2}} \cdot l = \frac{1}{\gamma} L$$

A falling apple accelerates and units of space get more and more contracted:



Space contracts near mass and dilates away from it.

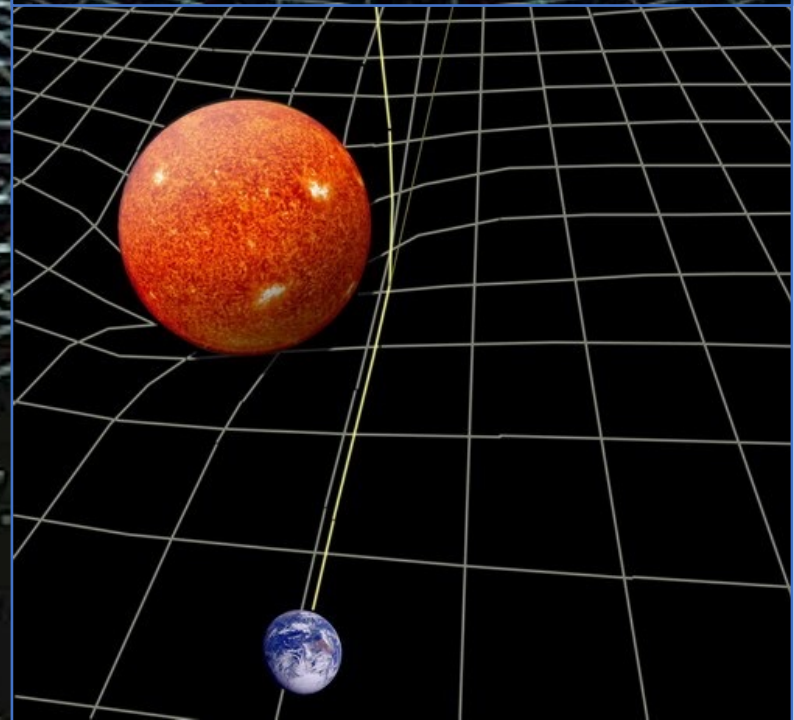
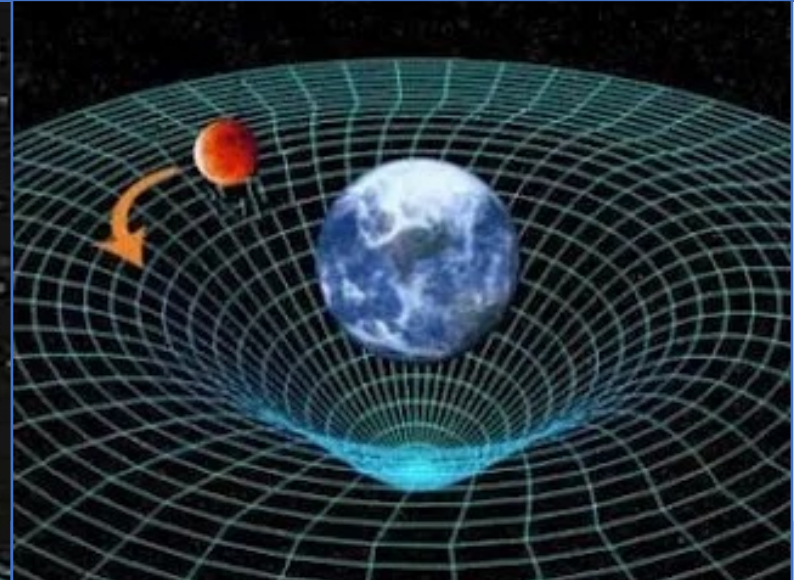
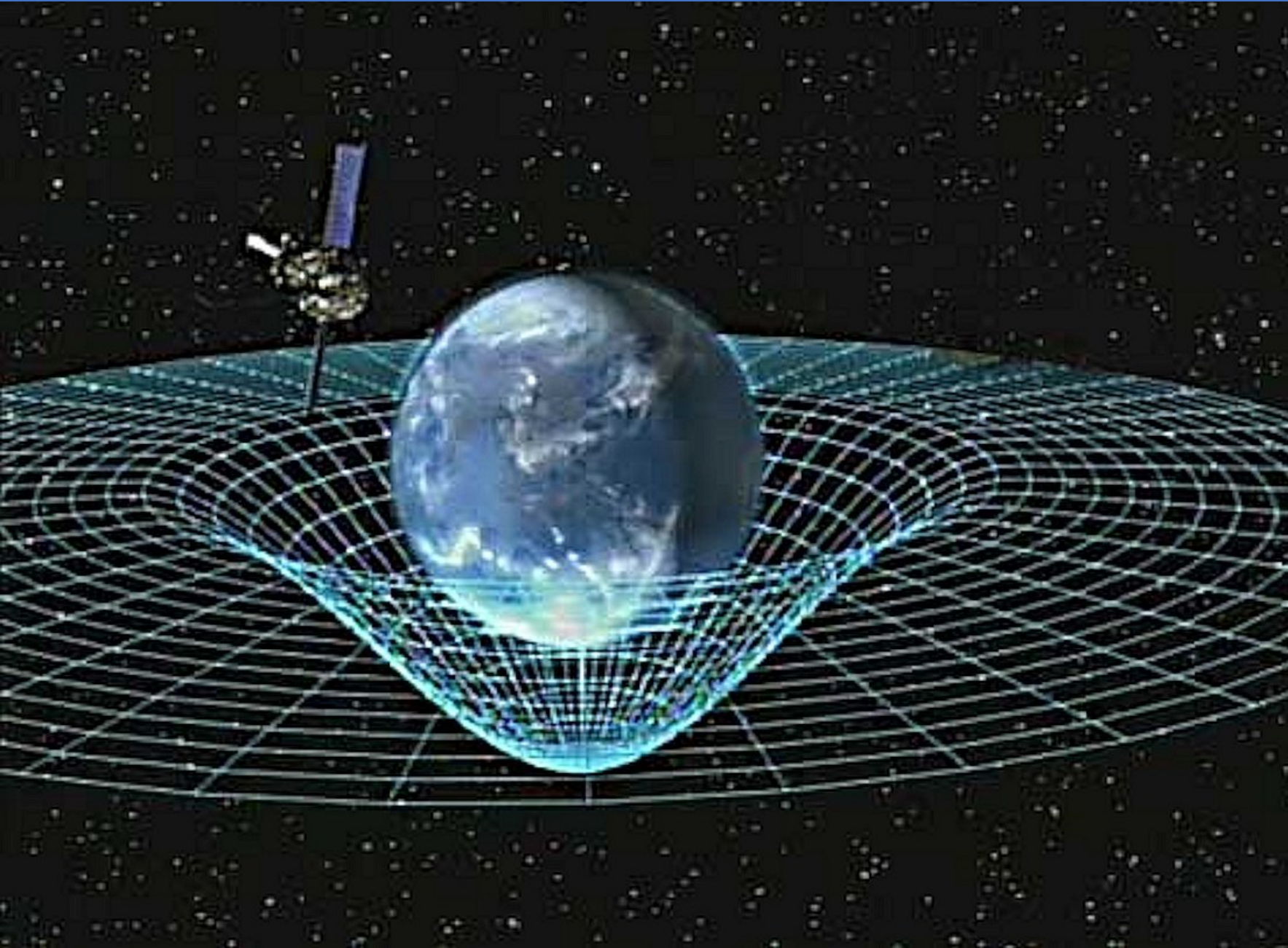
An apple falls into the gravitational field and time runs slower and slower:



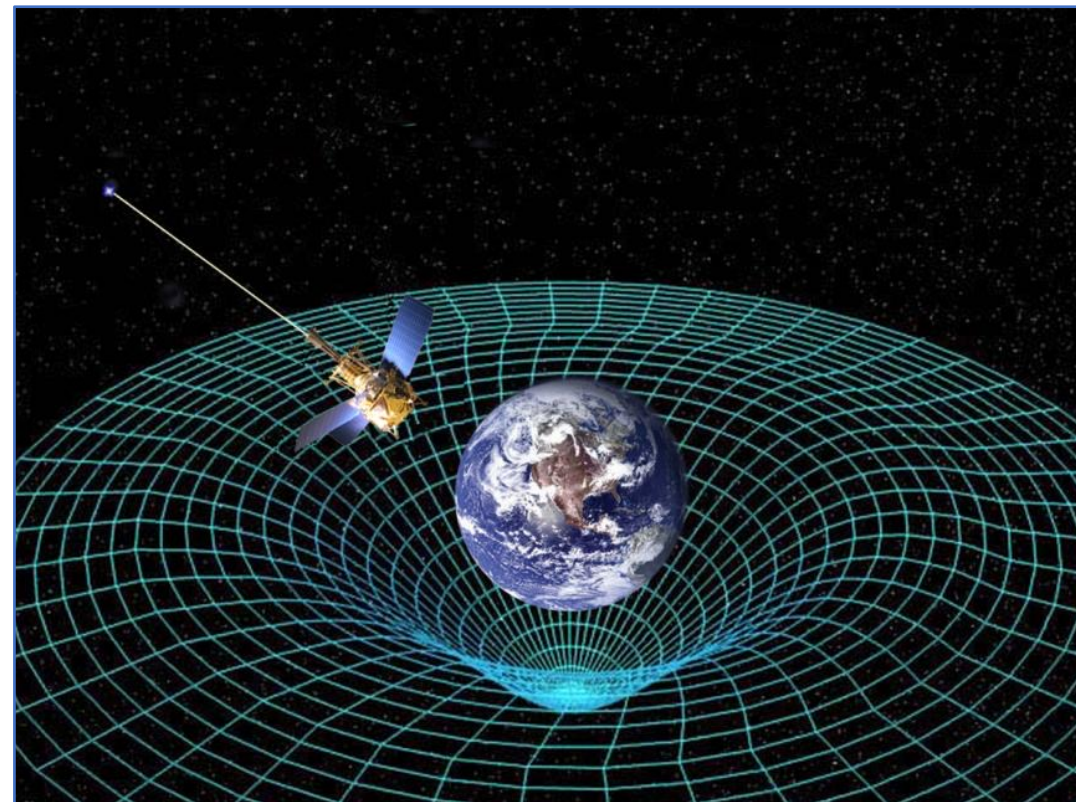
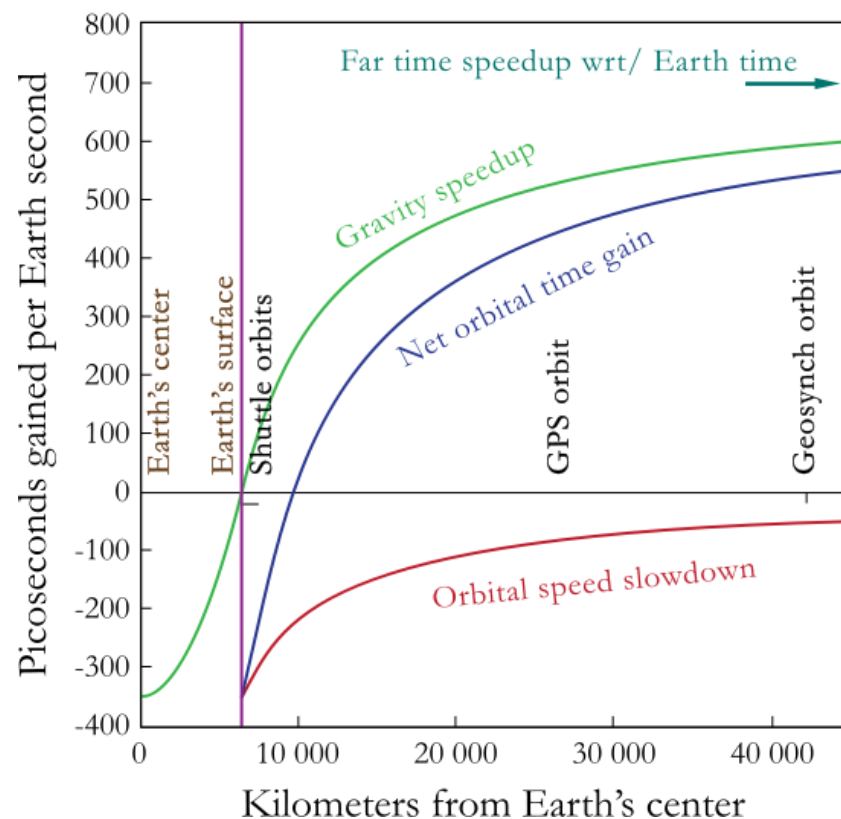
Time slows near mass and speeds up away from it.

Space-time is curved in the presence of mass

Mass causes curvature in space-time



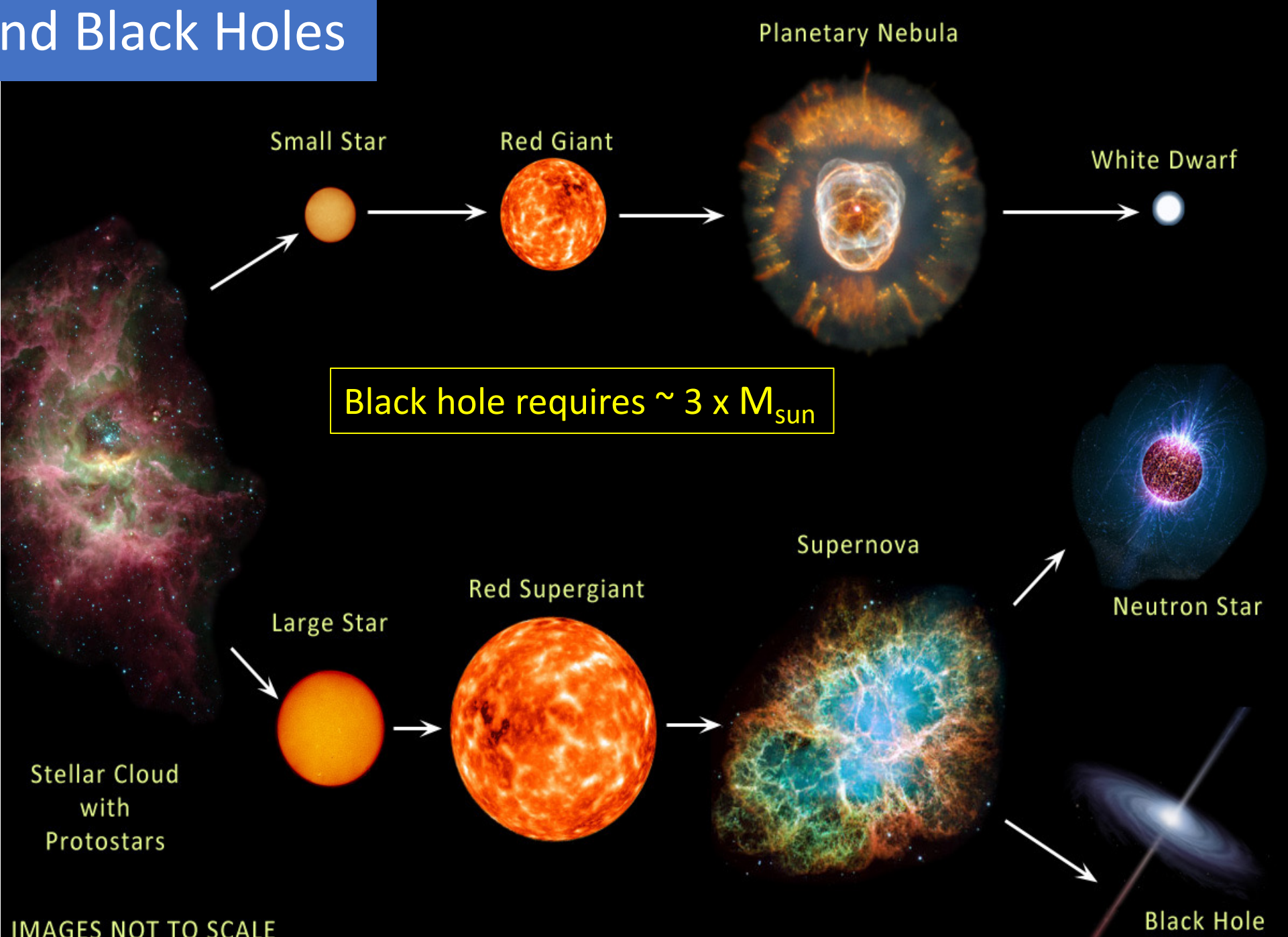
Time Dilation Effects on Earth



Two effects:

- **Time speeds up** at the satellite in comparison to earth surface **due to gravity**
- **Time slows down** at the satellite **due to high velocity** compared to person on earth

➔ **Clocks in satellite and on earth de-synchronize with ~ 40 msec per day!**



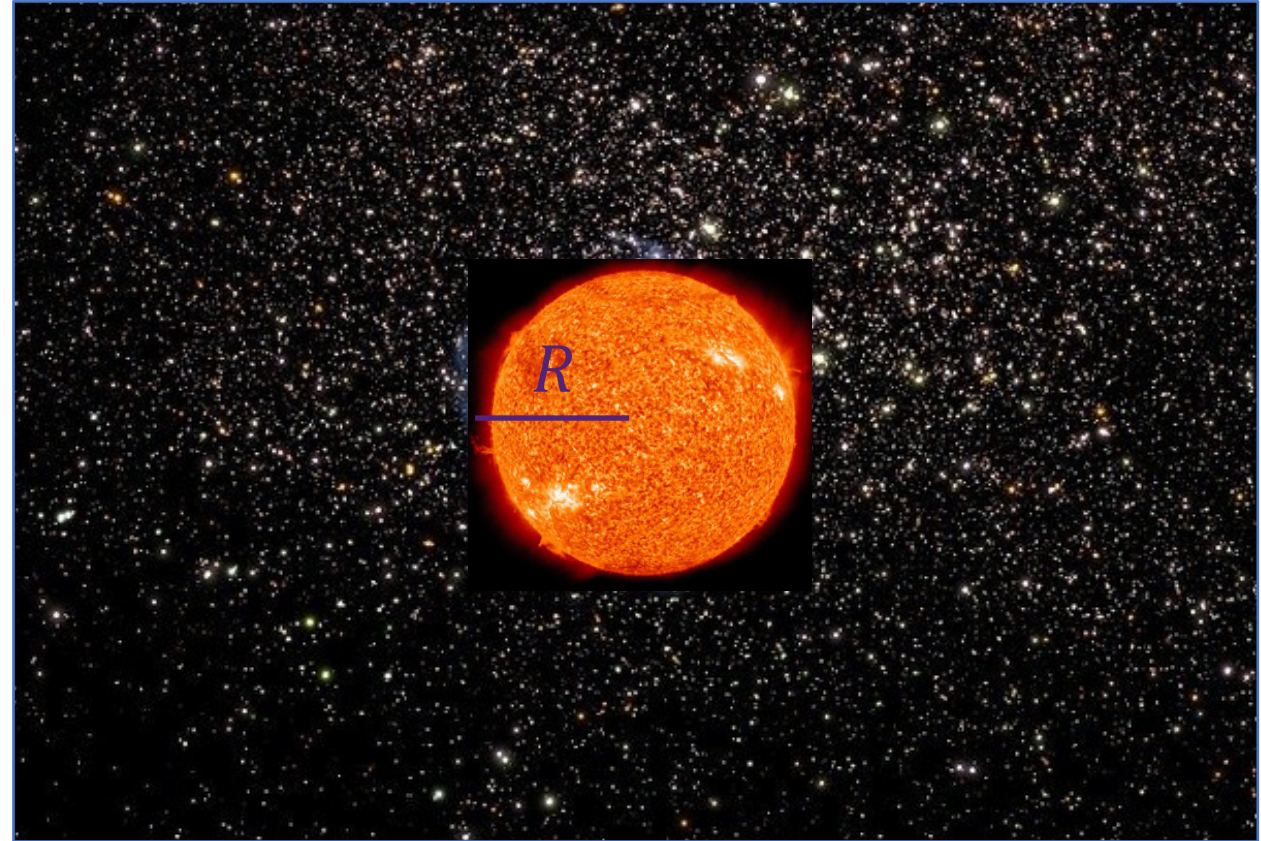
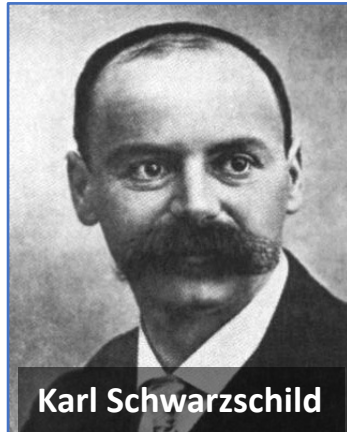
IMAGES NOT TO SCALE

Gravitational time slowdown near a star with mass M :

$$\Delta t' = \Delta t \sqrt{1 - \frac{2GM}{Rc^2}}$$

Schwartzschild radius: $R_s = \frac{2GM}{c^2}$

$$\Delta t' = \Delta t \sqrt{1 - \frac{R_s}{R}}$$

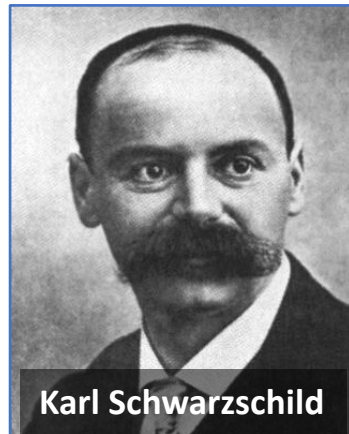


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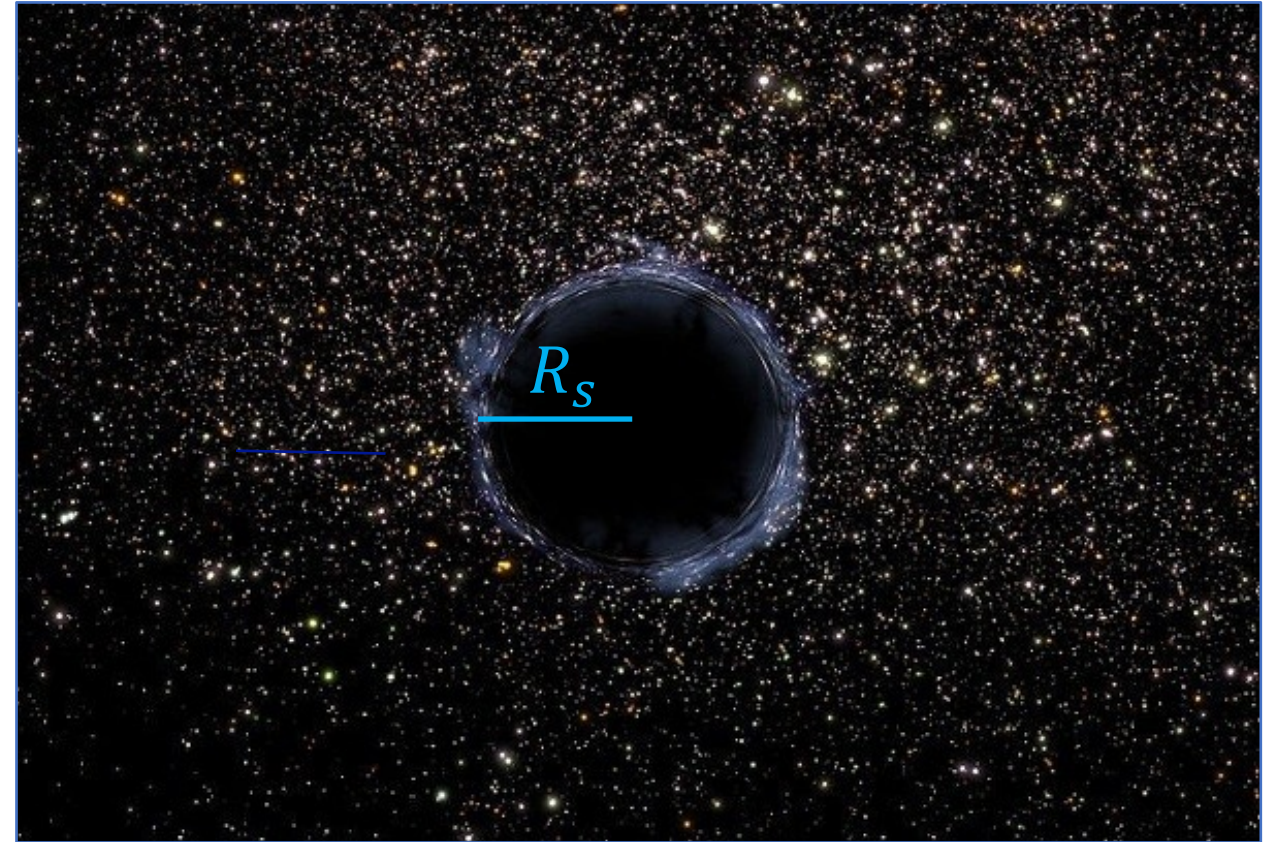
Time stand-still:

If $R = R_s$ then $\Delta t = 0$

Example our sun: $G = 6.67 \times 10^{-11} \text{ m}^3/\text{kg s}^2$ (Newton's gravitation constant)

$$M_{\text{sun}} = 2 \times 10^{30} \text{ kg}$$

→ $R_s = 3 \text{ km}$ for a black hole



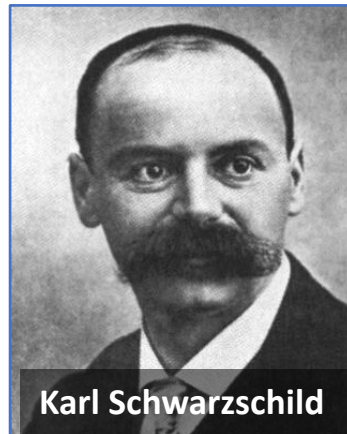
(Time stands still at the horizon of a black-hole)

Gravitational time slowdown near a star with mass M :

$$\Delta t' = \Delta t \sqrt{1 - \frac{2GM}{Rc^2}}$$

Schwartzschild radius: $R_s = \frac{2GM}{c^2}$

$$\Delta t' = \Delta t \sqrt{1 - \frac{R_s}{R}}$$



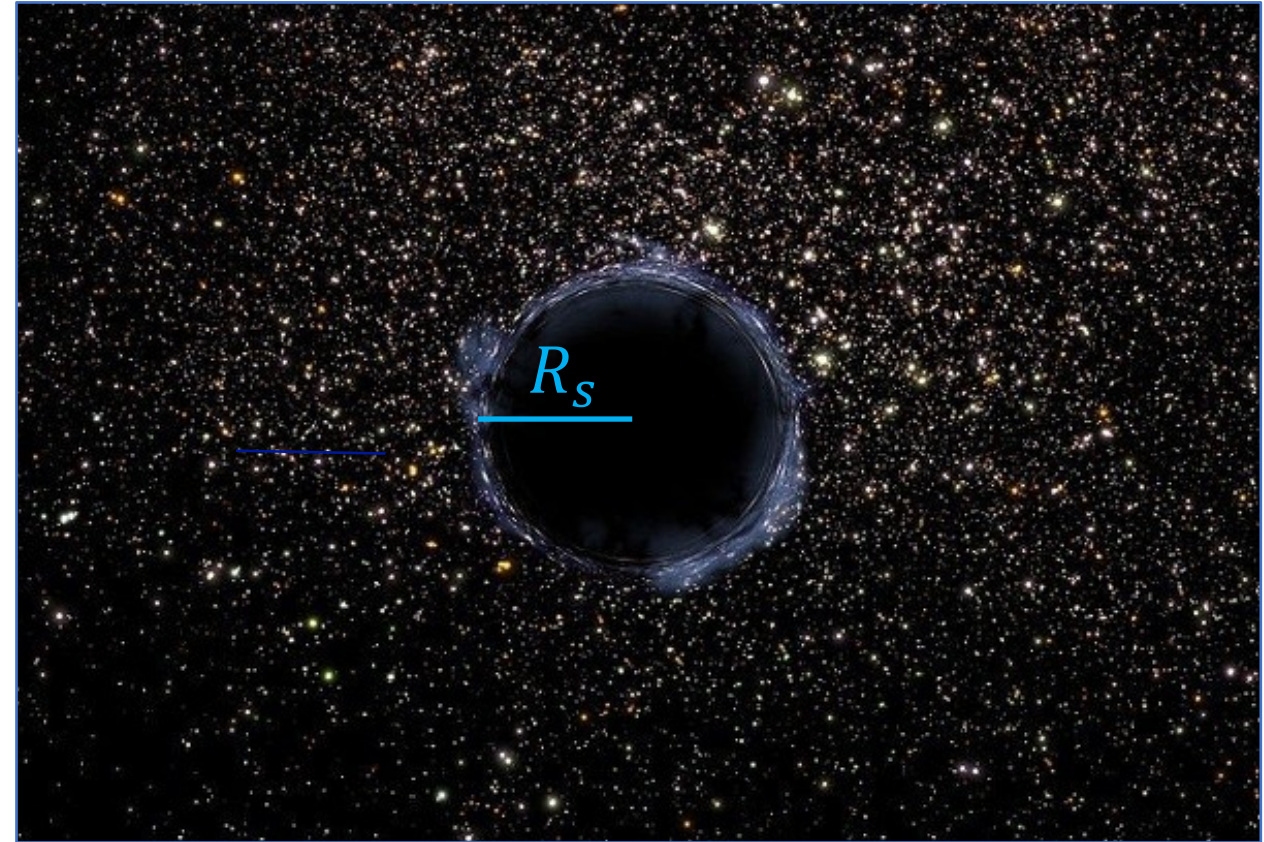
Time stand-still:

If $R = R_s$ then $\Delta t = 0$

Example our earth: $G = 6.67 \times 10^{-11} \text{ m}^3/\text{kg s}^2$ (Newton's gravitation constant)

$$M_{\text{earth}} = 6 \times 10^{24} \text{ kg}$$

→ $R_s = 9 \text{ mm}$ for a black hole



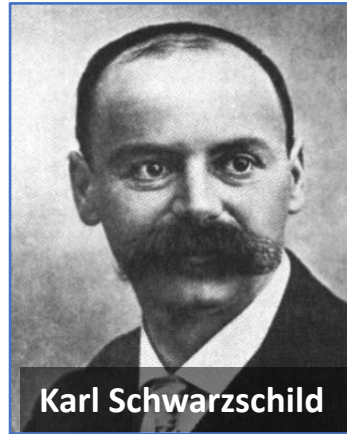
(Time stands still at the horizon of a black-hole)

Gravitational time slowdown near a star with mass M :

$$\Delta t' = \Delta t \sqrt{1 - \frac{2GM}{Rc^2}}$$

Schwartzschild radius: $R_s = \frac{2GM}{c^2}$

$$\Delta t' = \Delta t \sqrt{1 - \frac{R_s}{R}}$$



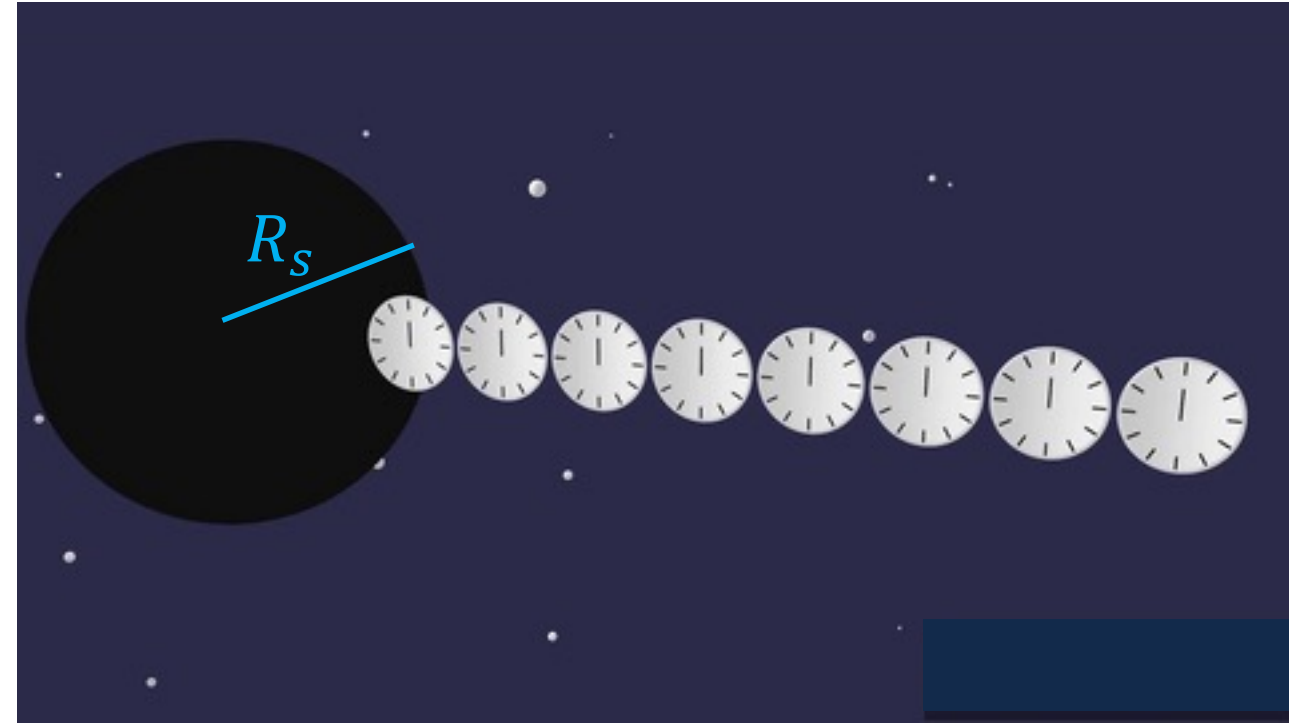
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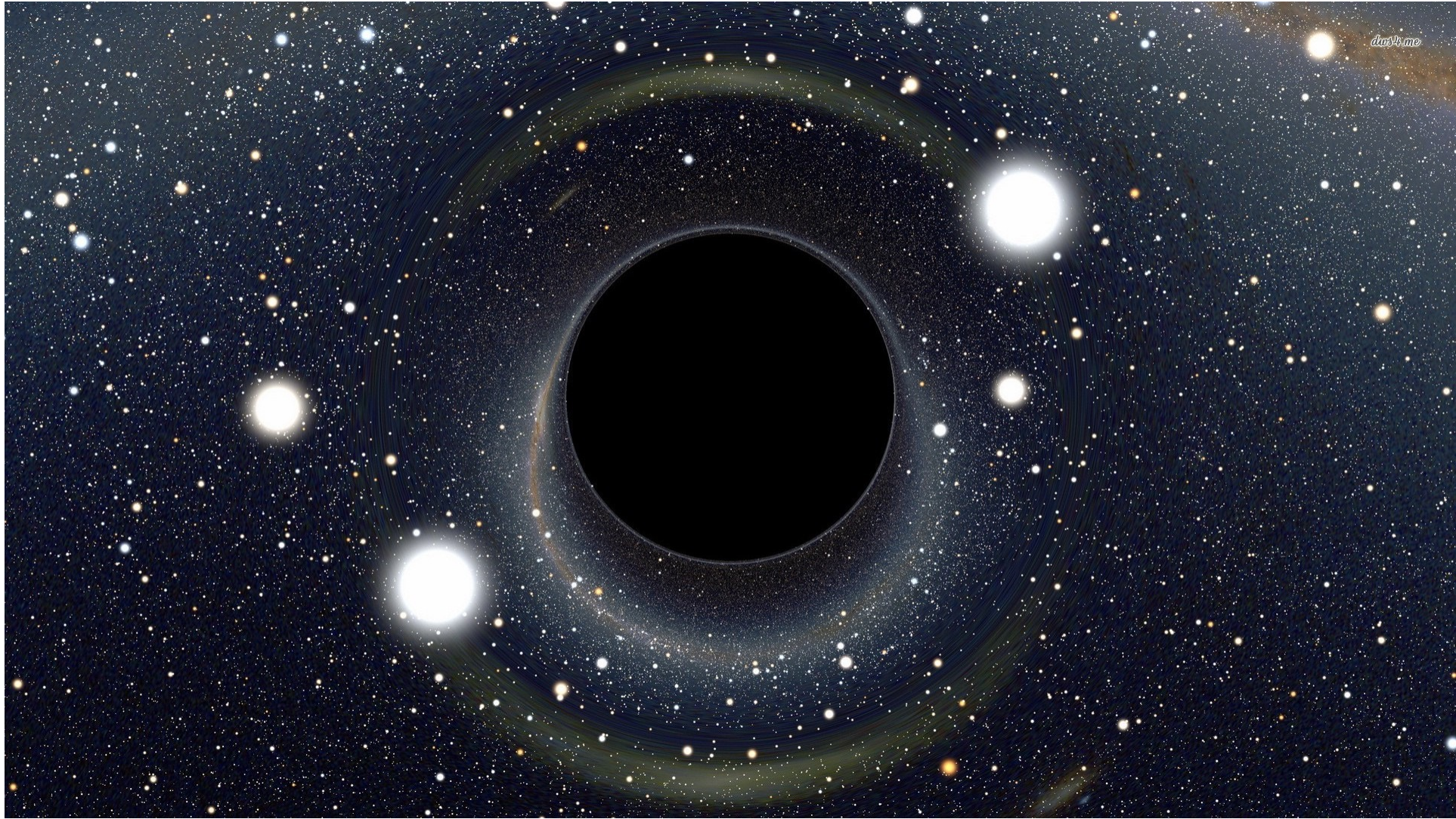
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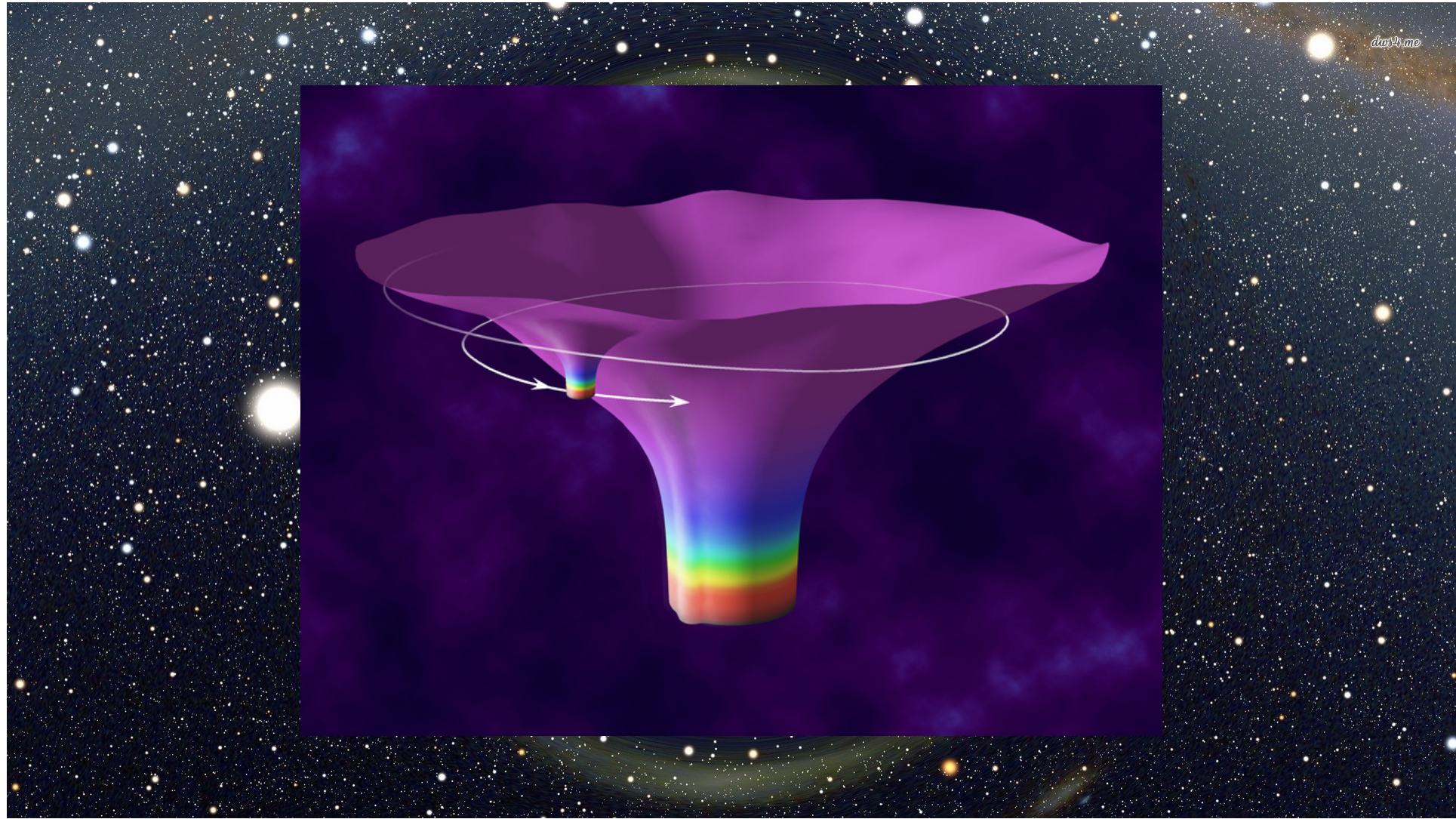


(Time stands still at the horizon of a black-hole)



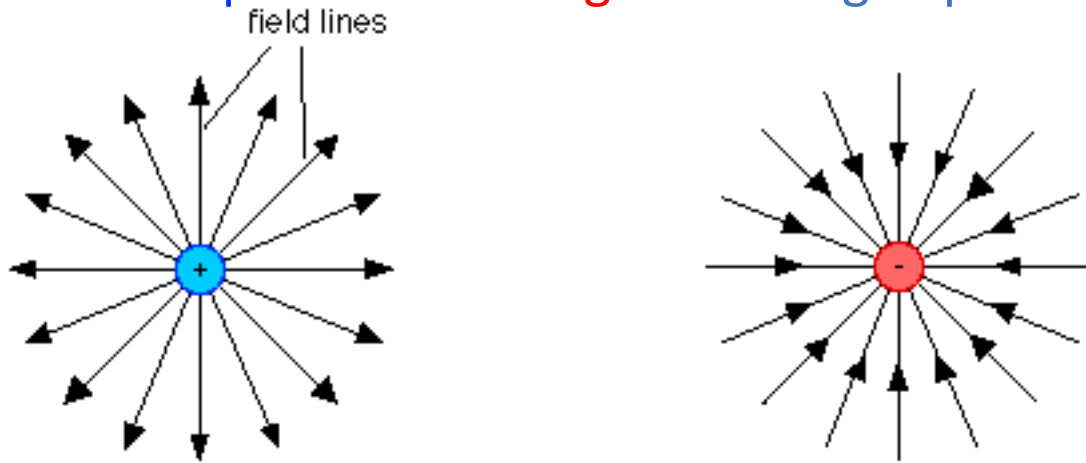
Purely curved space-time!

What *is* a black hole?

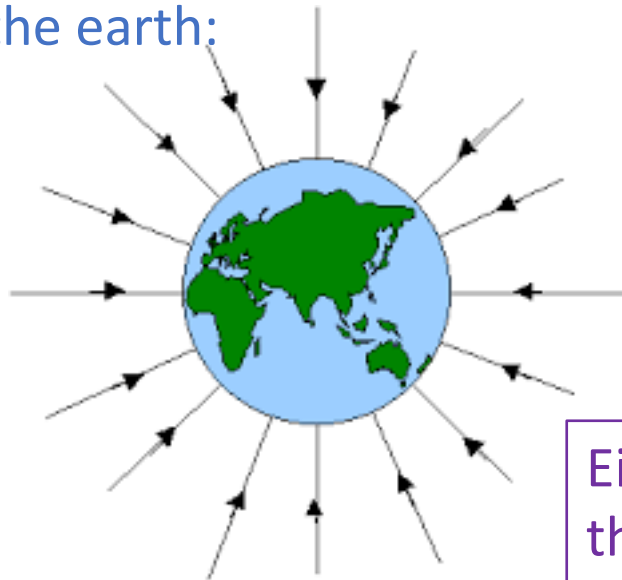


What happens when two black holes meet?

Electric field of **positive** and **negative** charged particle:



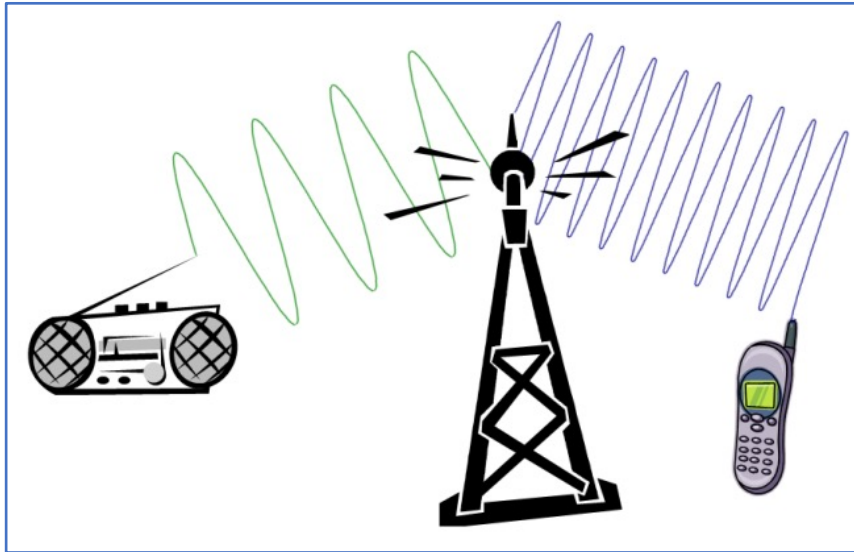
Gravitational field of the earth:



$$F_e = q E = \frac{1}{4\pi\epsilon_0} \frac{qQ}{r^2}$$

$$F_g = m g = -G \frac{mM}{R^2}$$

Einstein spent most of his life looking for a unified theory of electromagnetism and general relativity.



Electromagnetic waves:

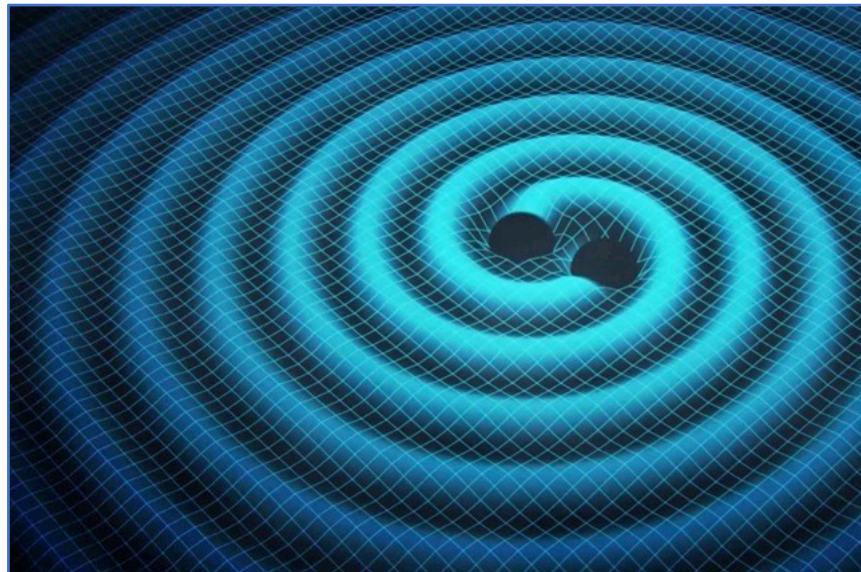
Caused by accelerating electric particles (electrons)
eg.: radio-emission

Maxwell equations:

$$\partial_{\mu} F^{\mu\nu} = \frac{4\pi}{c} J^{\nu}$$

E and B fields

Electric charge
and currents



Gravitational Waves:

Caused by moving masses.

Requires very heavy masses → **black holes.**

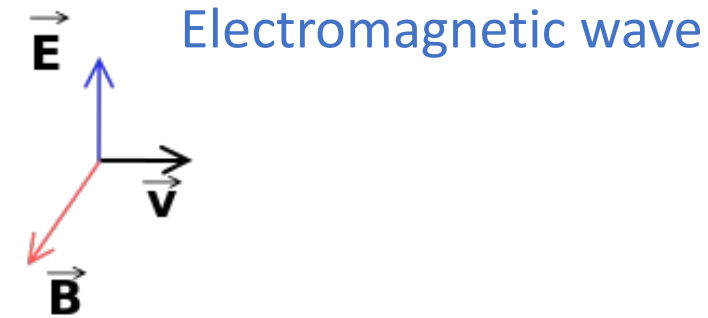
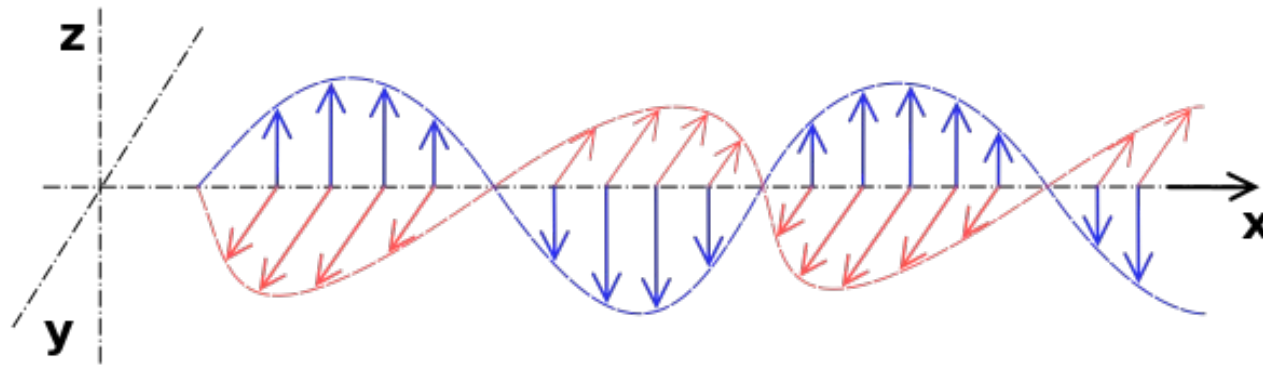
(Einstein thought these couldn't be observed)

Einstein equations:

$$G^{\mu\nu} + \Lambda g^{\mu\nu} = \frac{8\pi G}{c^4} T^{\mu\nu}$$

Space-time fields

Mass and mass-flow



Electromagnetic wave:

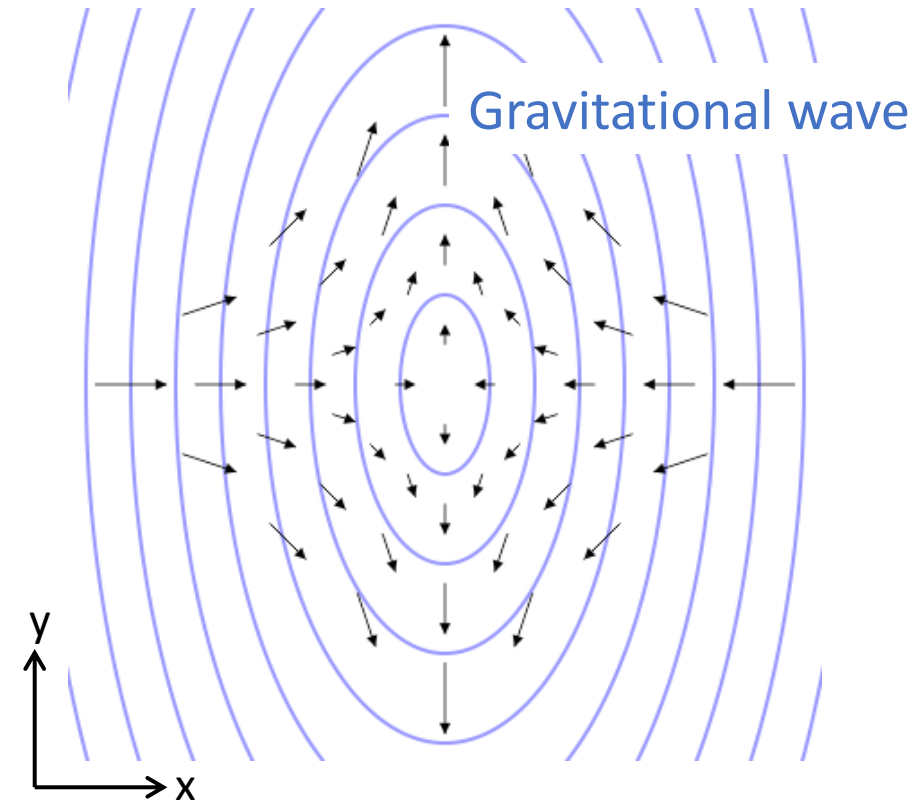
Changing electric and magnetic field propagating through space.

Caused by moving (accelerating!) electric charges.

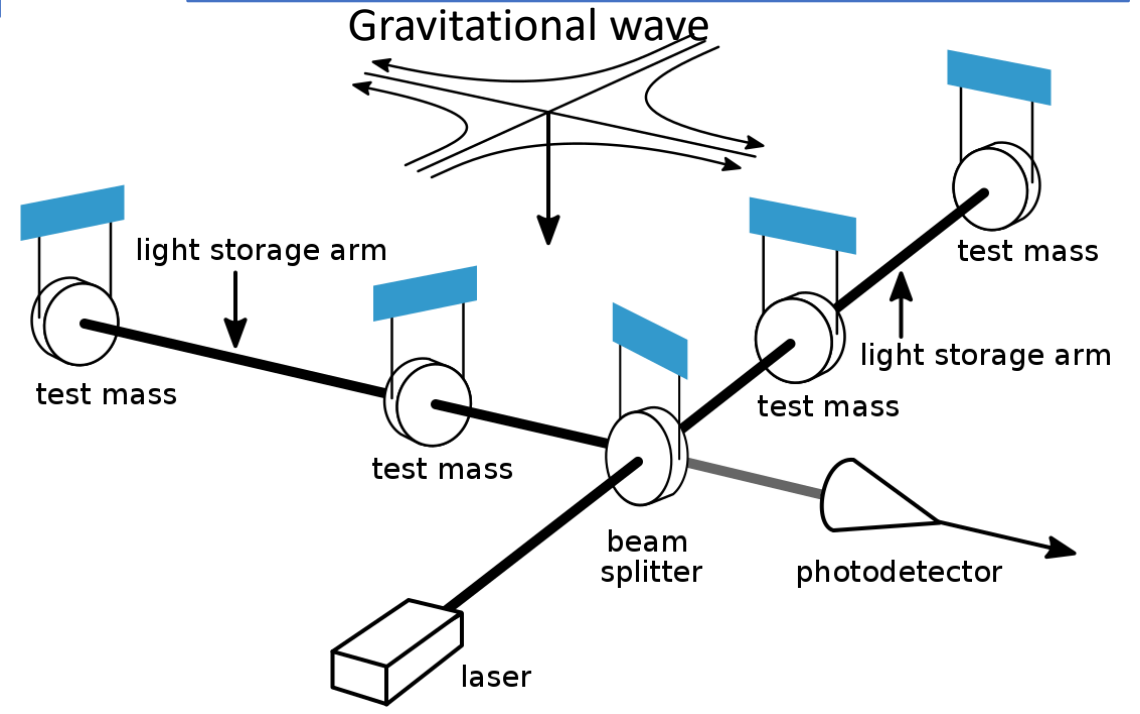
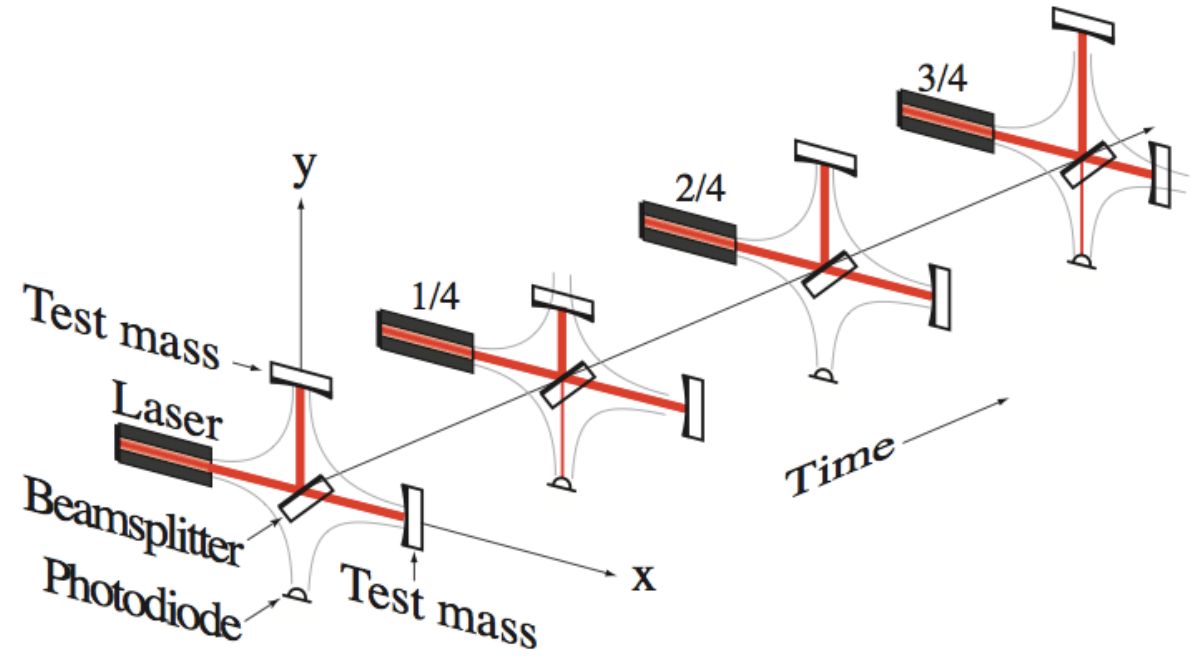
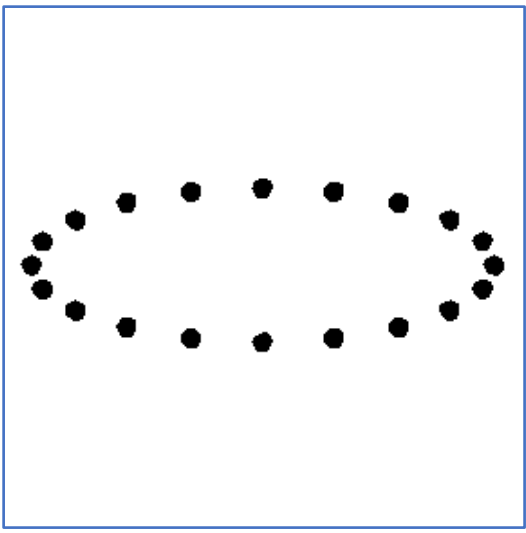
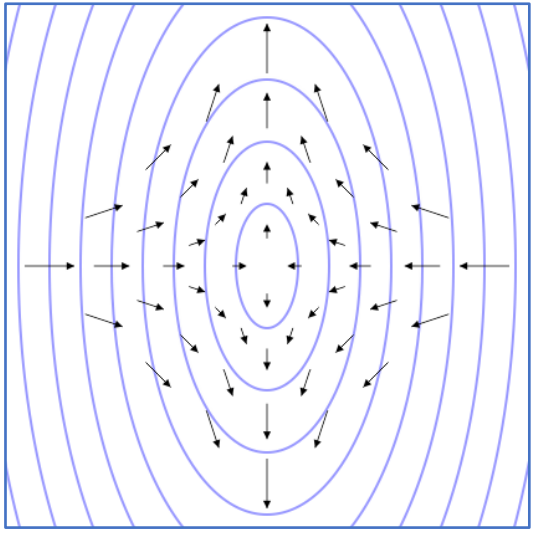
Gravitational wave:

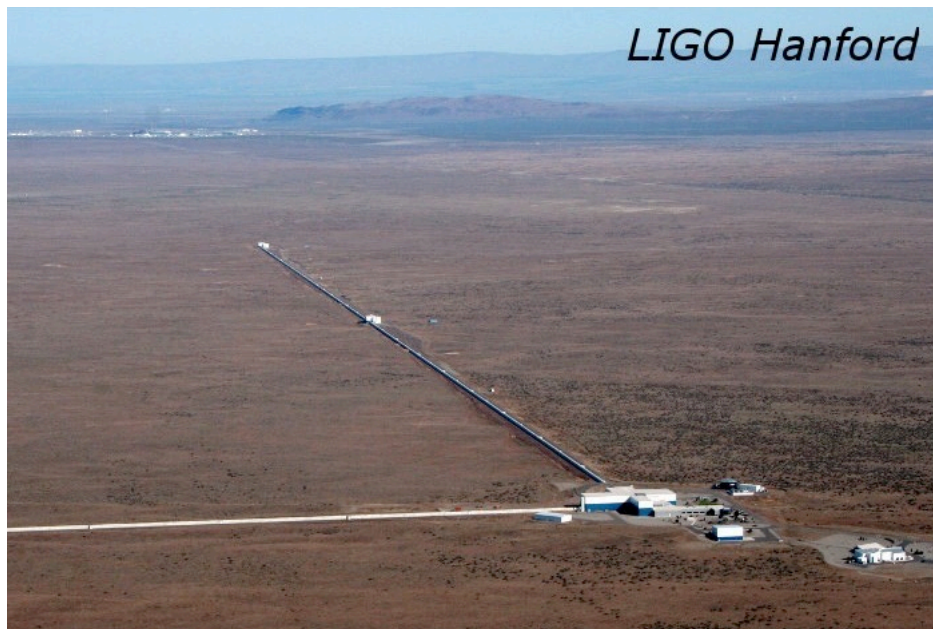
Changing space-time field.

Caused by moving (accelerating!) masses.

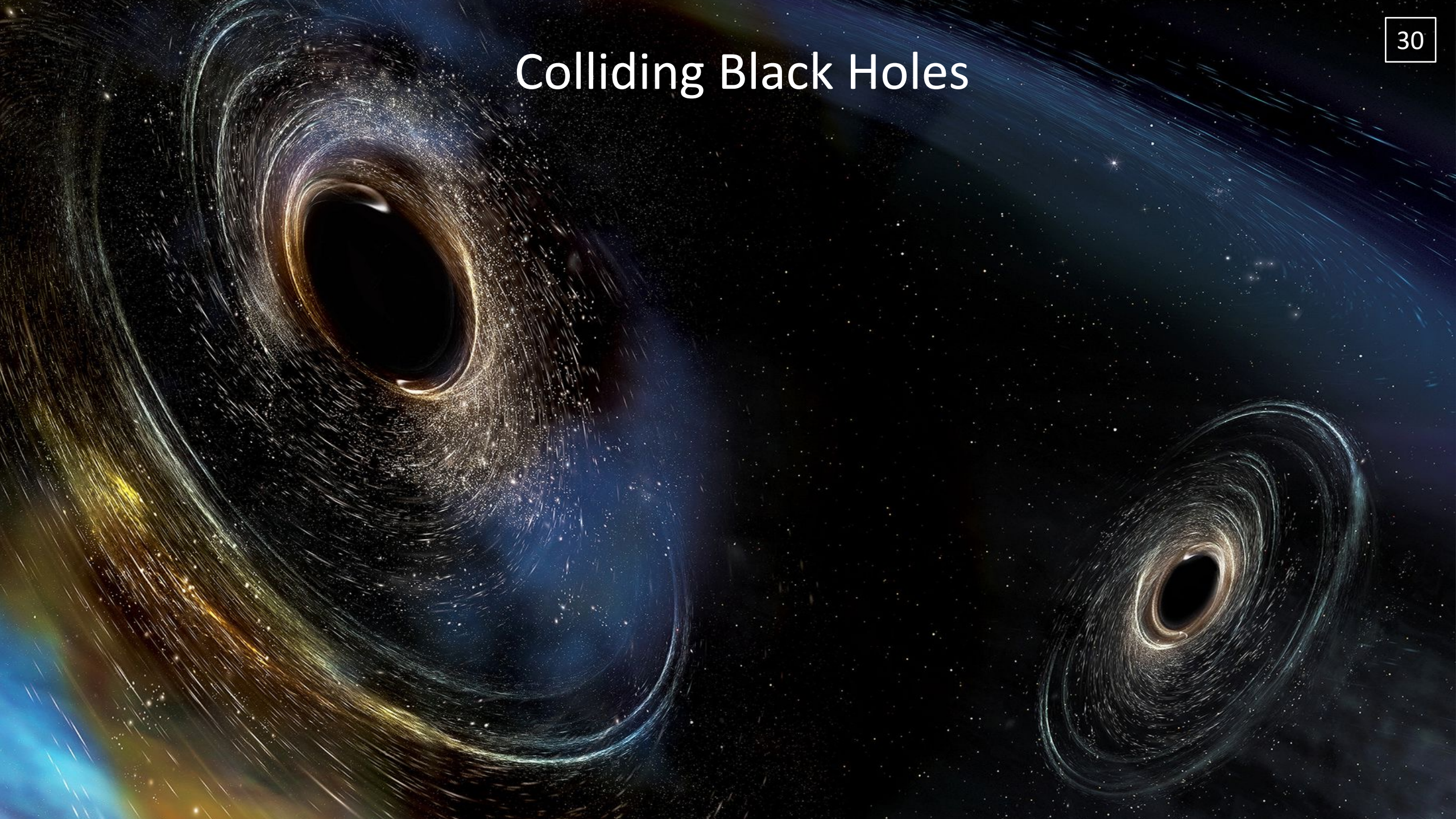


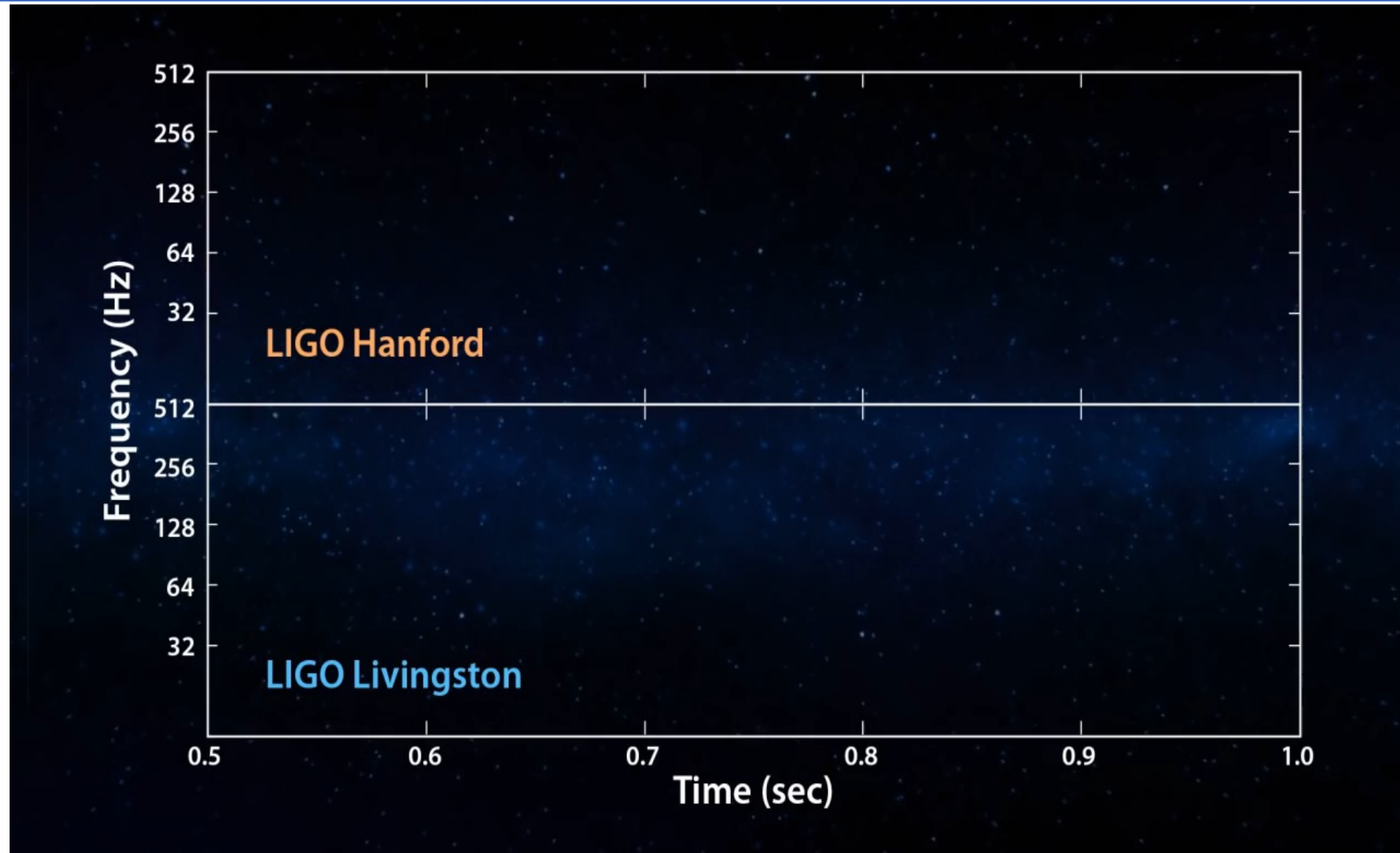
Remember the interferometer!





Colliding Black Holes



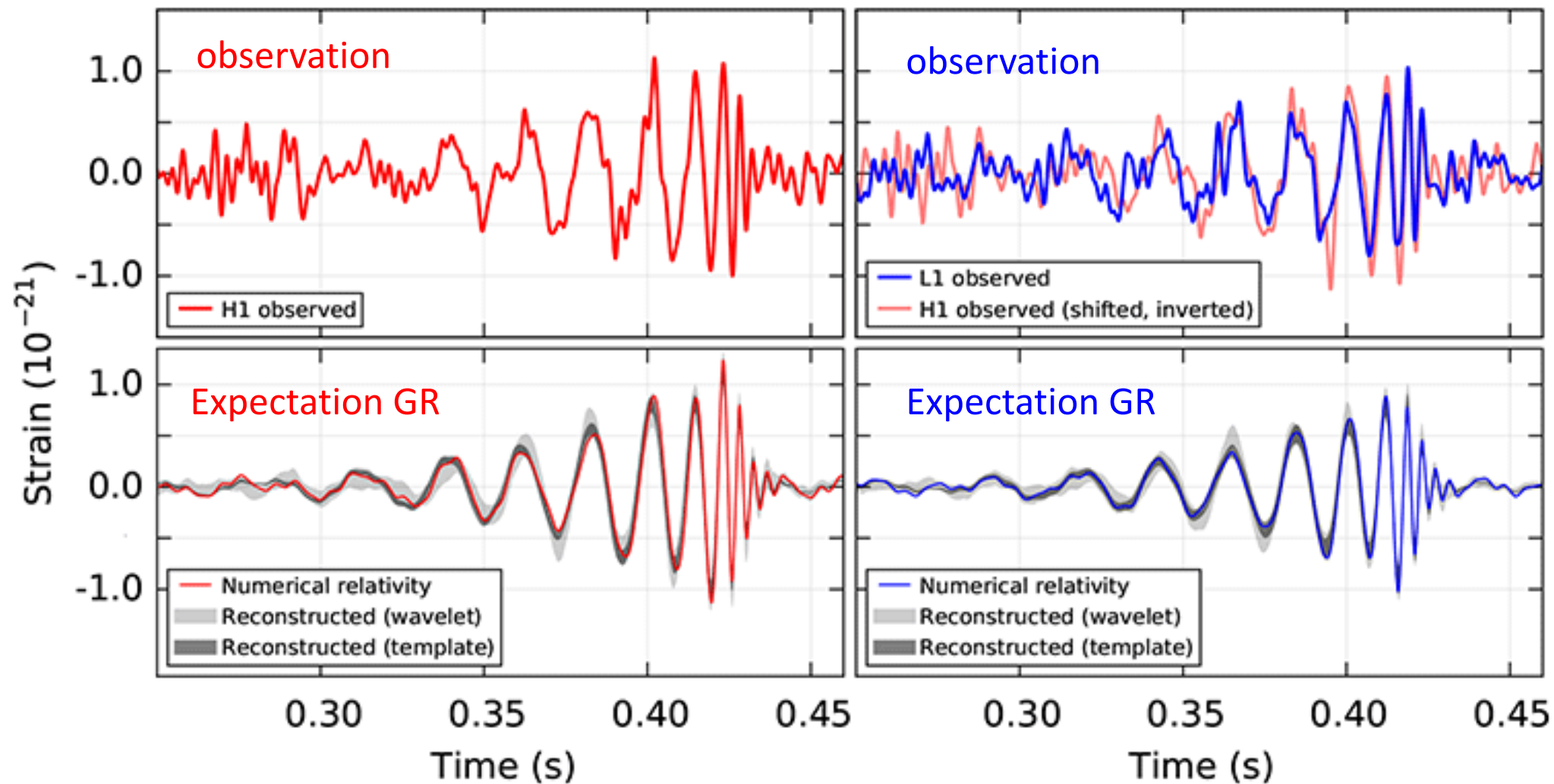


“Chirp” of colliding black holes at 1.3 billion lightyears distance

(GW150914)

Hanford, Washington

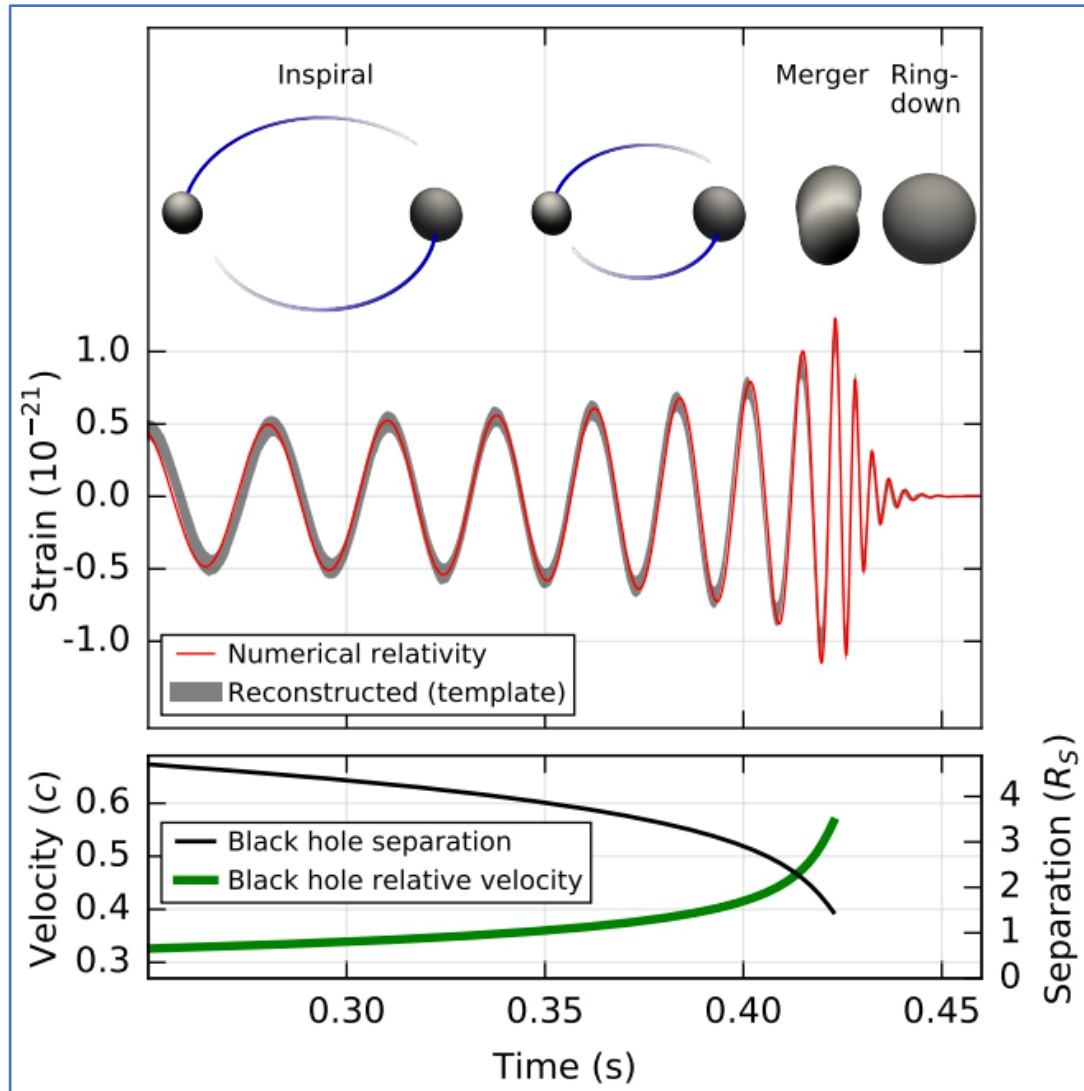
Livingston, Louisiana





Two Merging Black Holes

Two massive colliding/merging black holes:



Distance: 1.3 billion lightyears

B.H.1 = 36 x mass of the sun

B.H.2 = 29 x mass of the sun

New BH: 62 solar masses

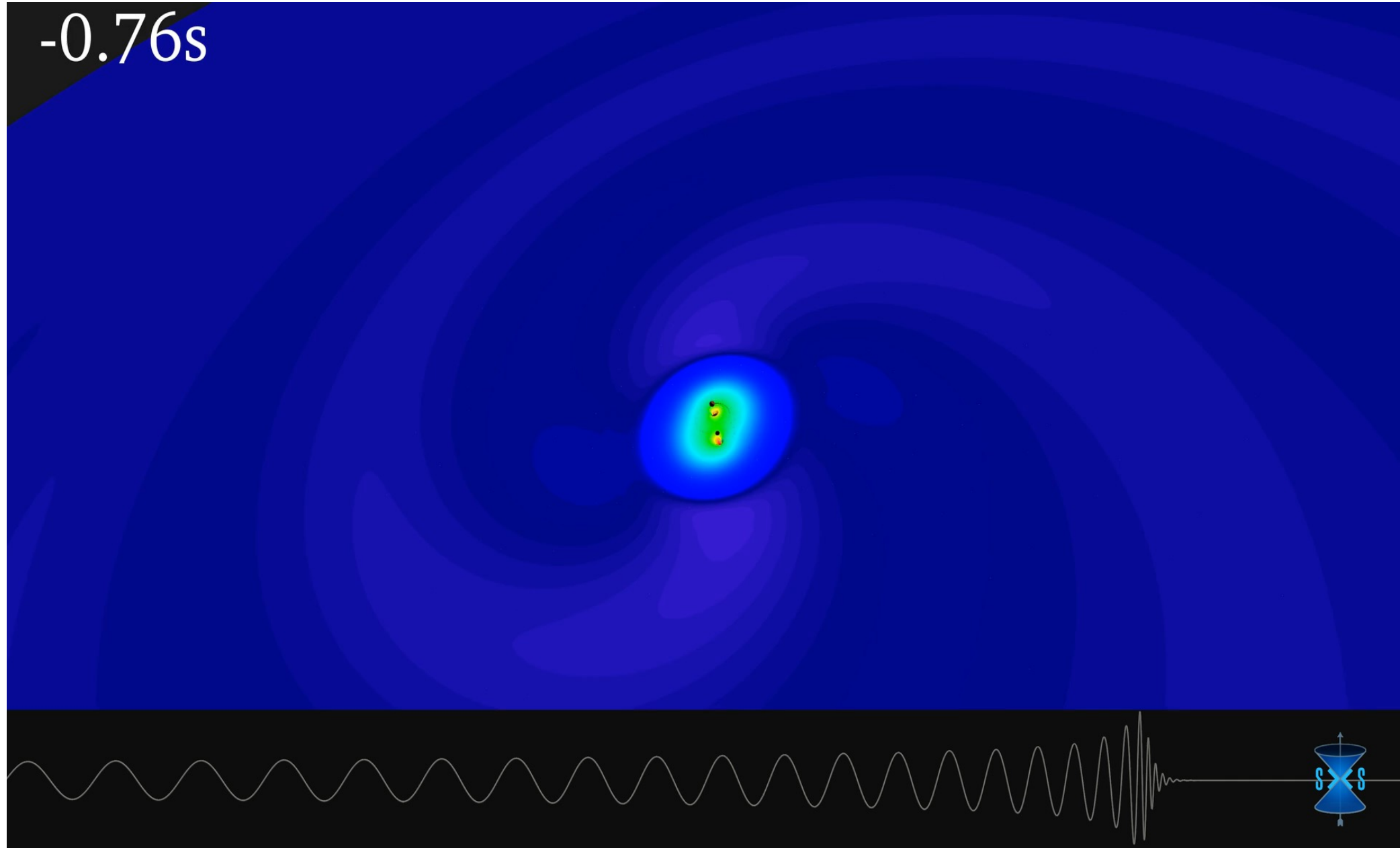
→ 3 solar masses of energy ($E=mc^2$)
radiated into space

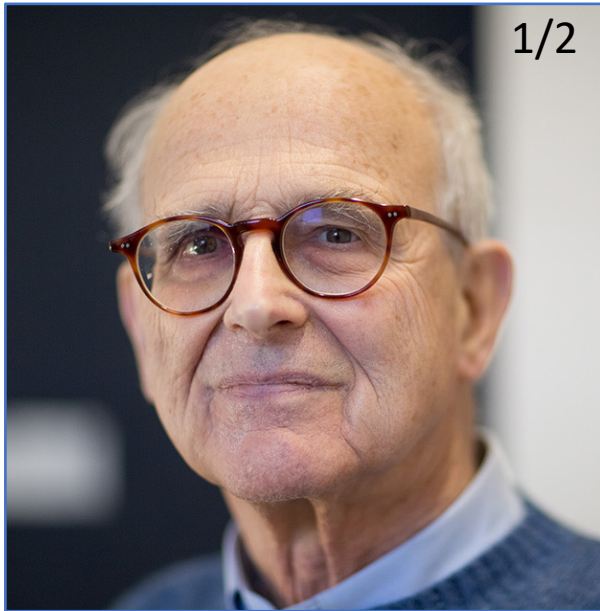
Relative change of space (strain)

0.000000000000000000000001%

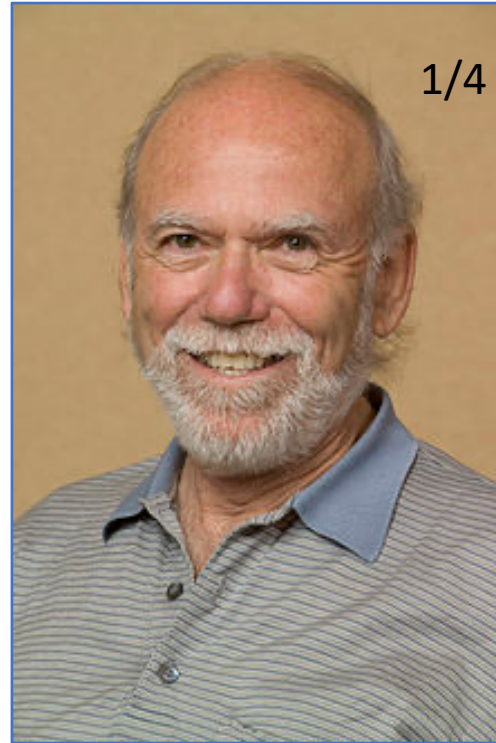
Rotation speed increasing to half the light speed!

More energy was emitted in gravitational waves than all the visible (EM) energy of all stars in the universe!





Rainer Weiss



Barry C. Barish



Kip S. Thorne

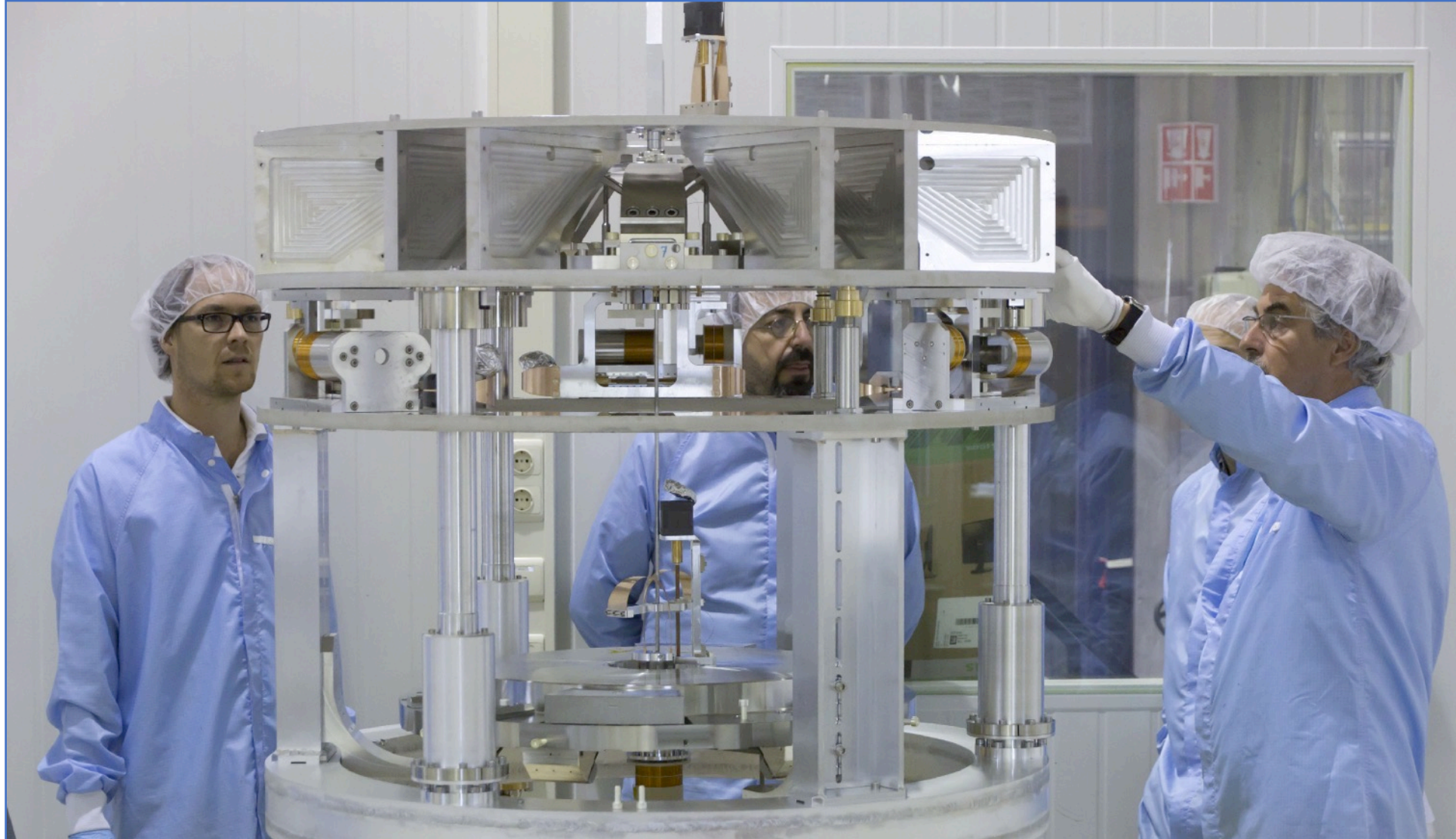
“For decisive contributions to the LIGO detector and the observation of gravitational waves”

The Virgo Experiment in Pisa

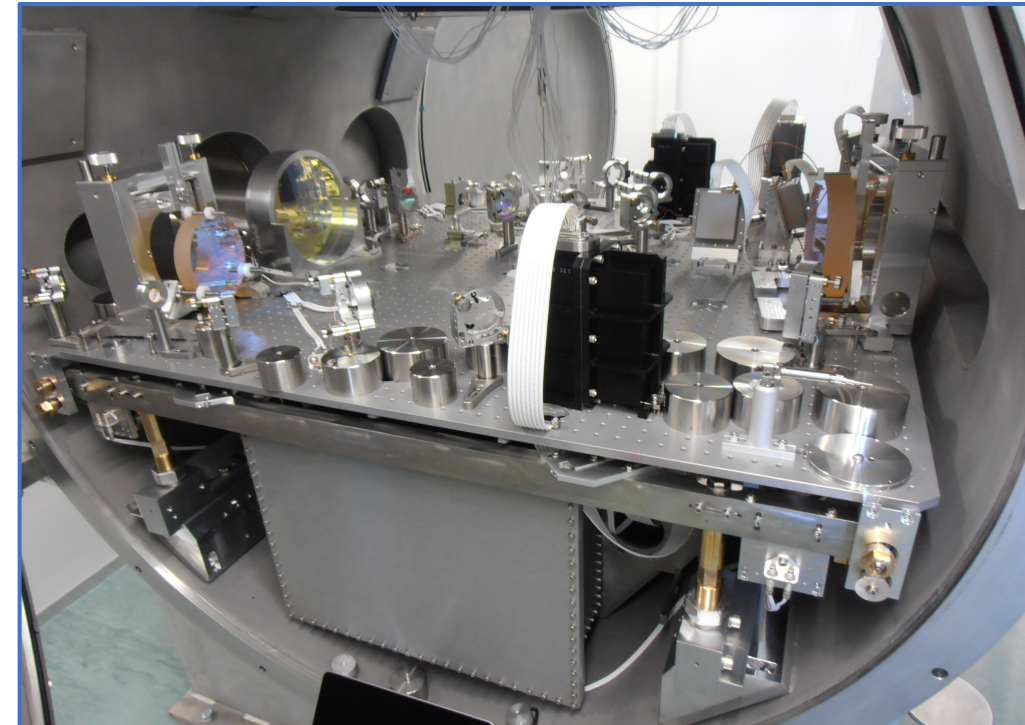
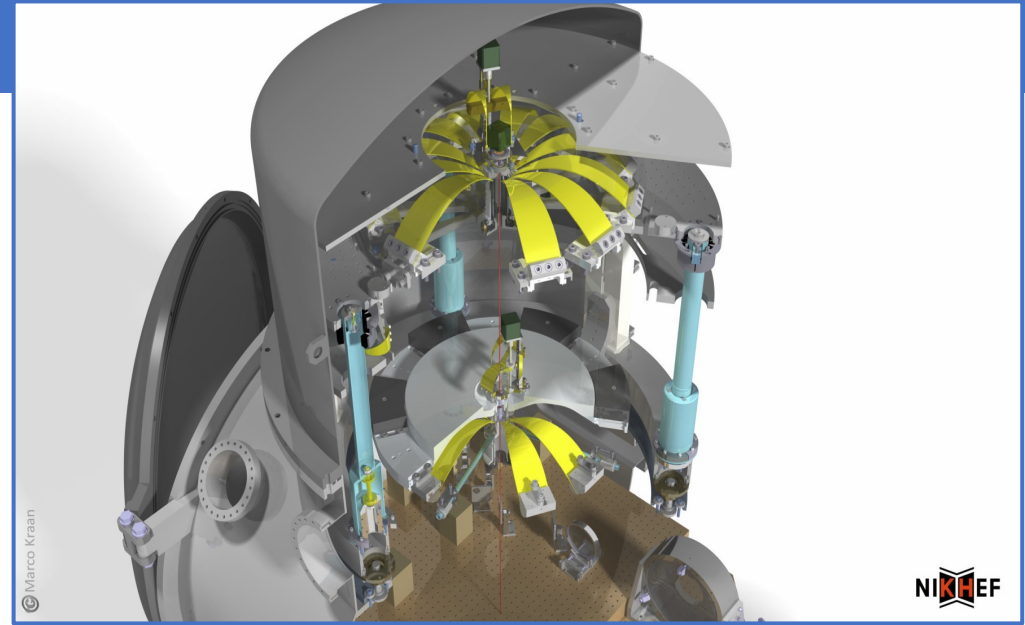
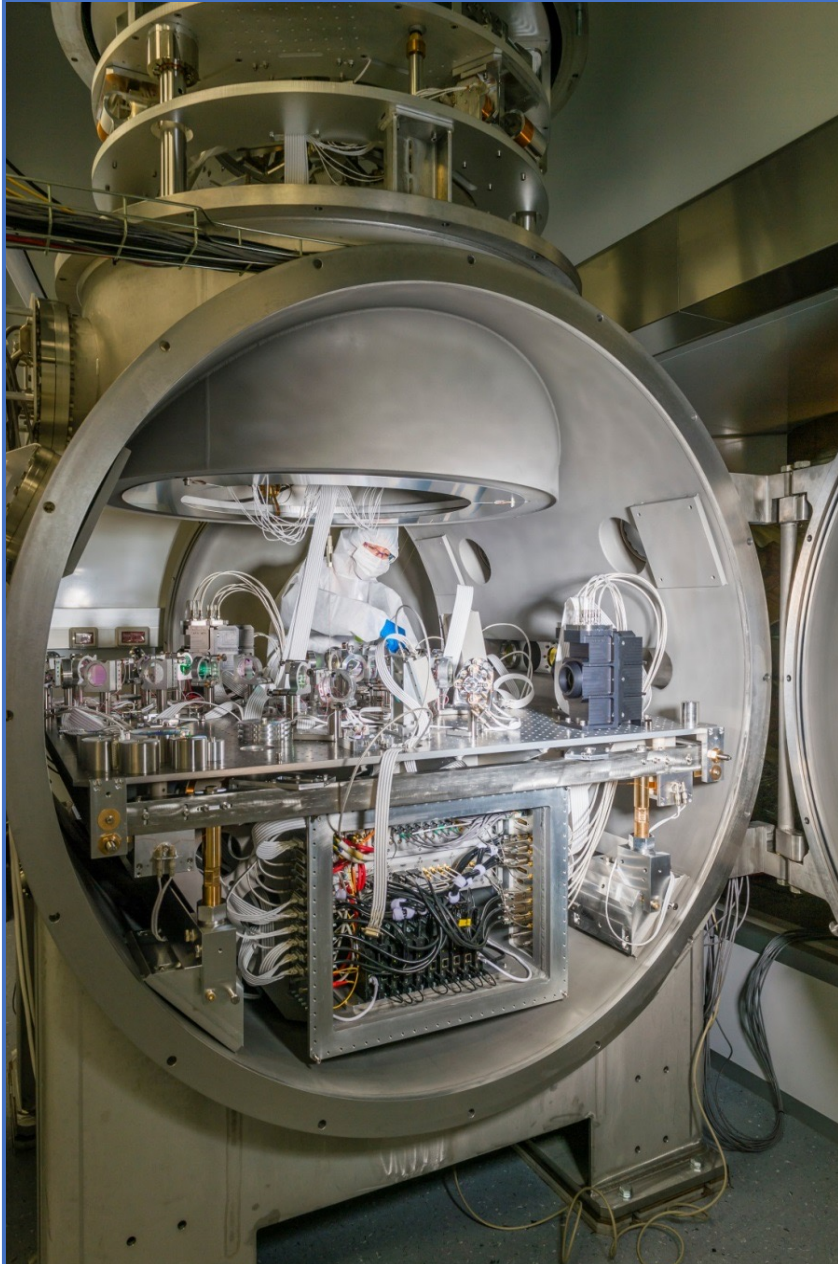




Largest vacuüm vessel in Europe: Pressure $\sim 10^{-10}$ mbar

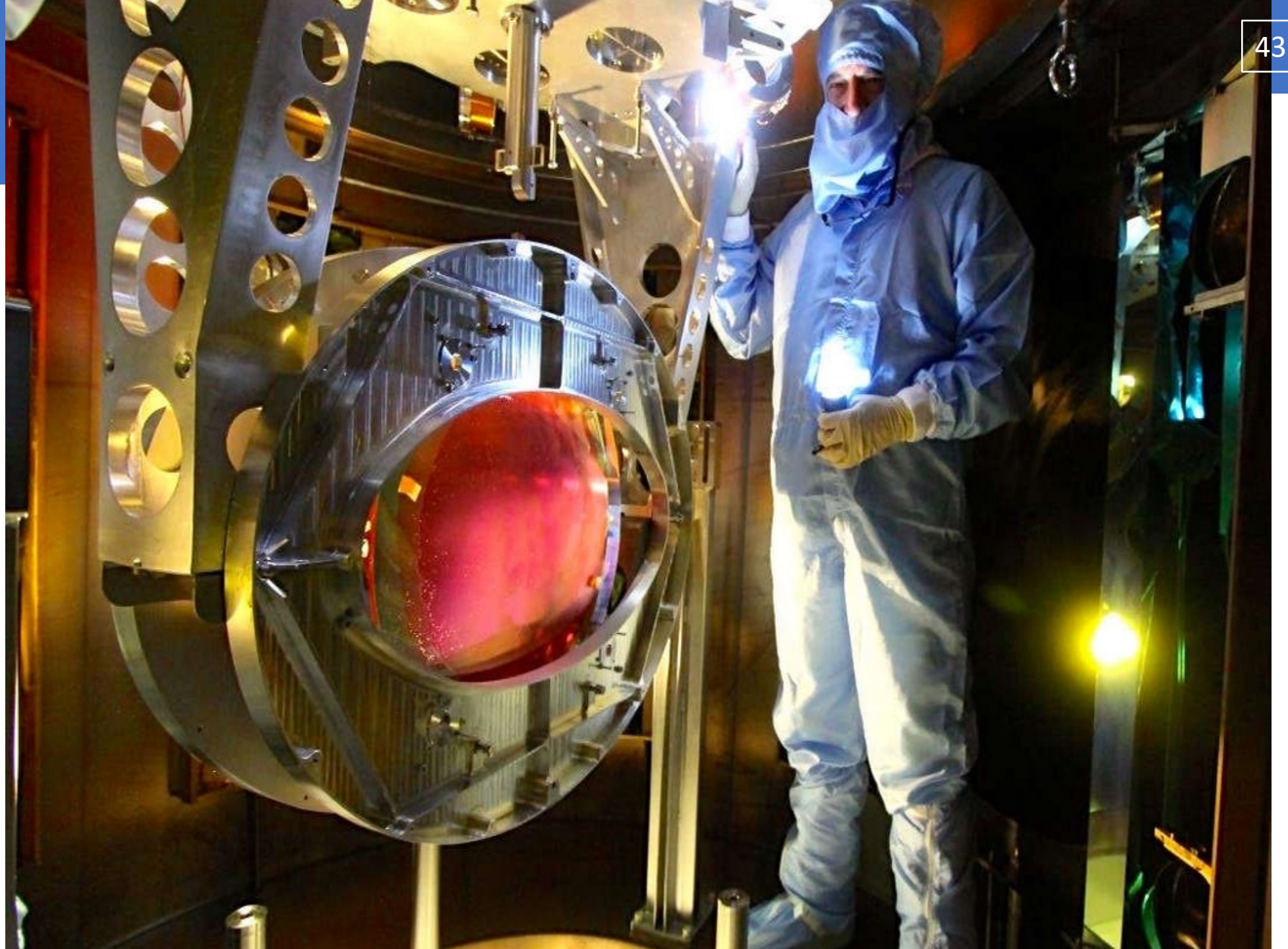


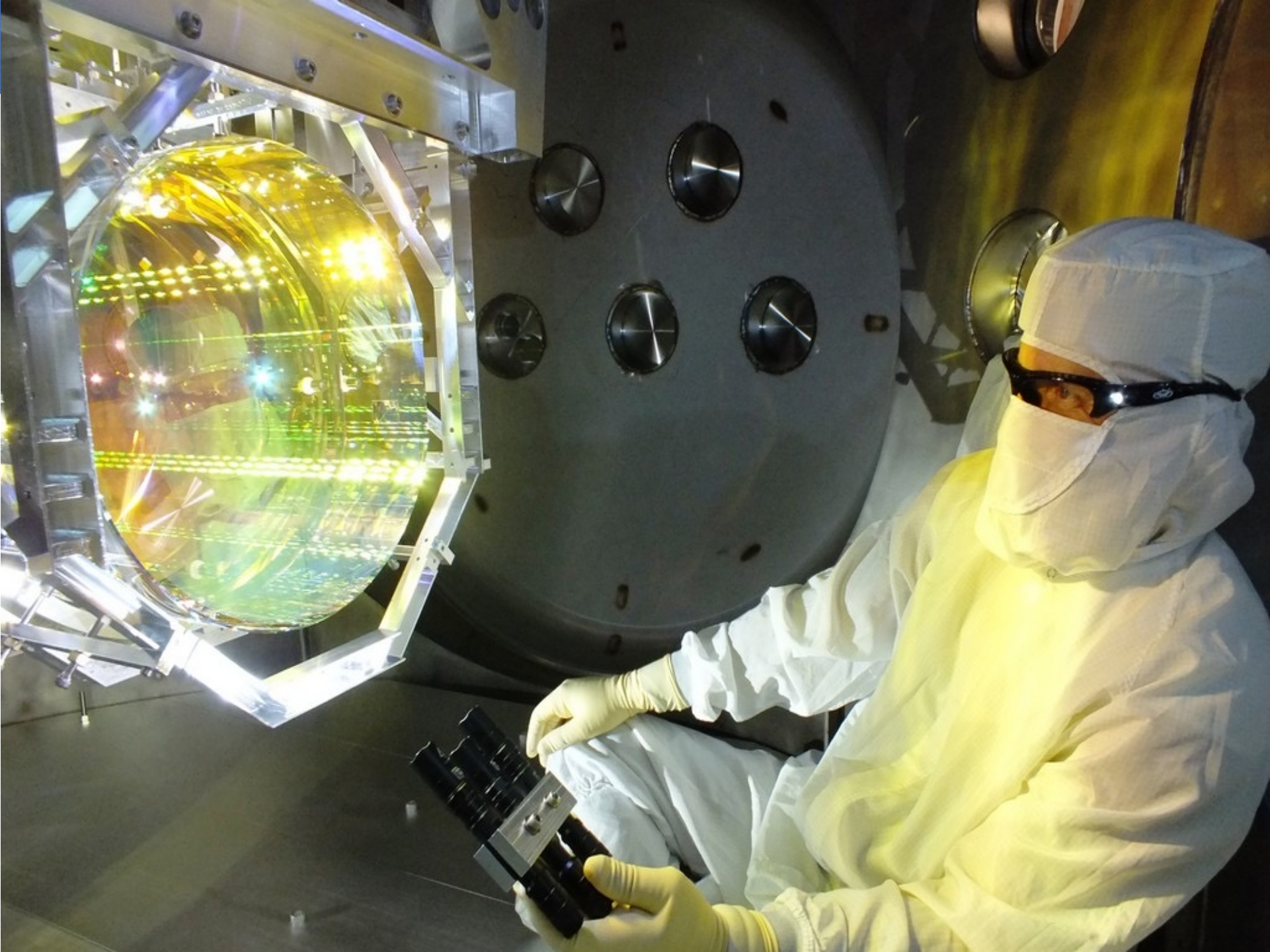
Seismic Damping Table

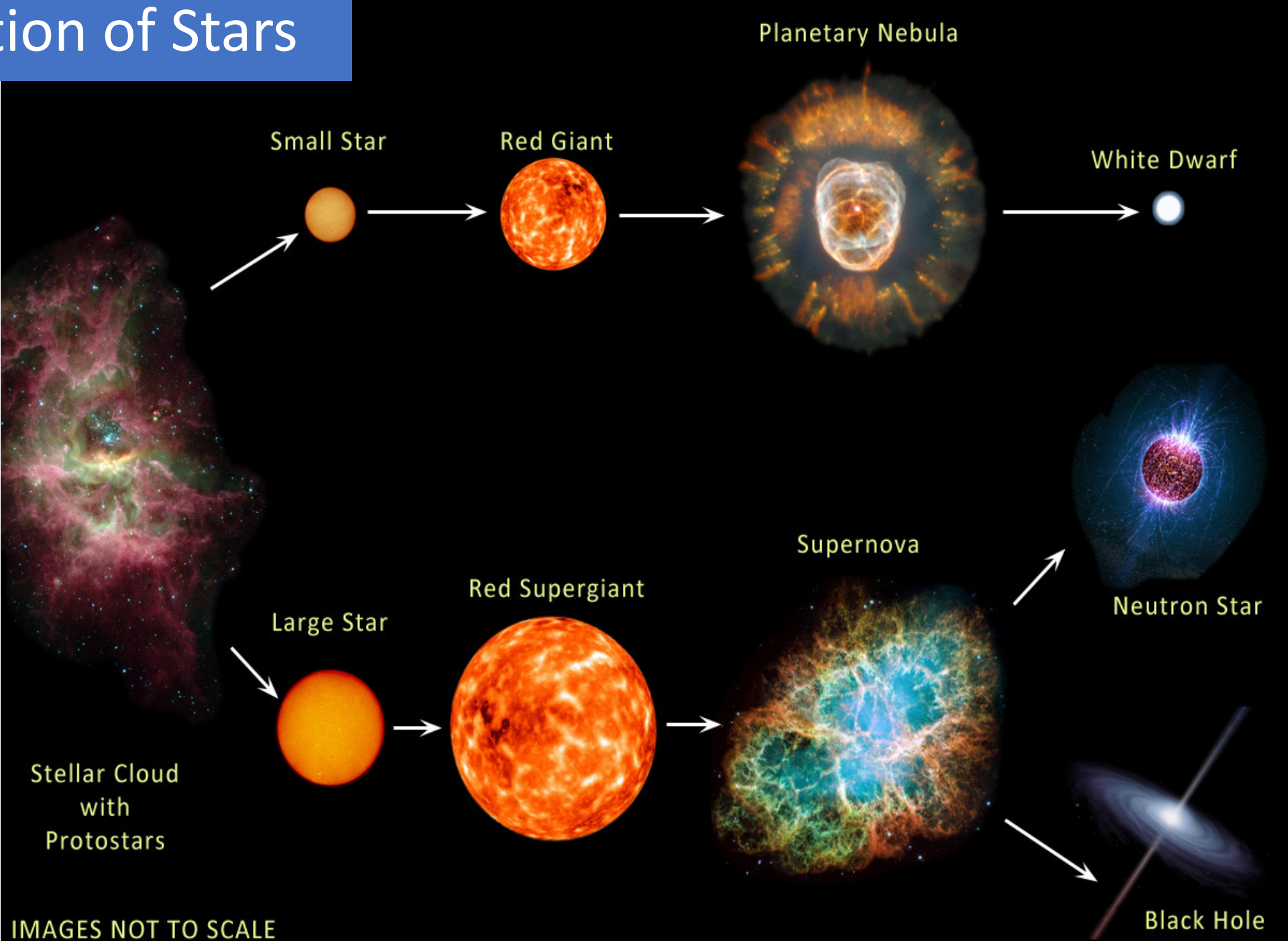




Beam Splitter



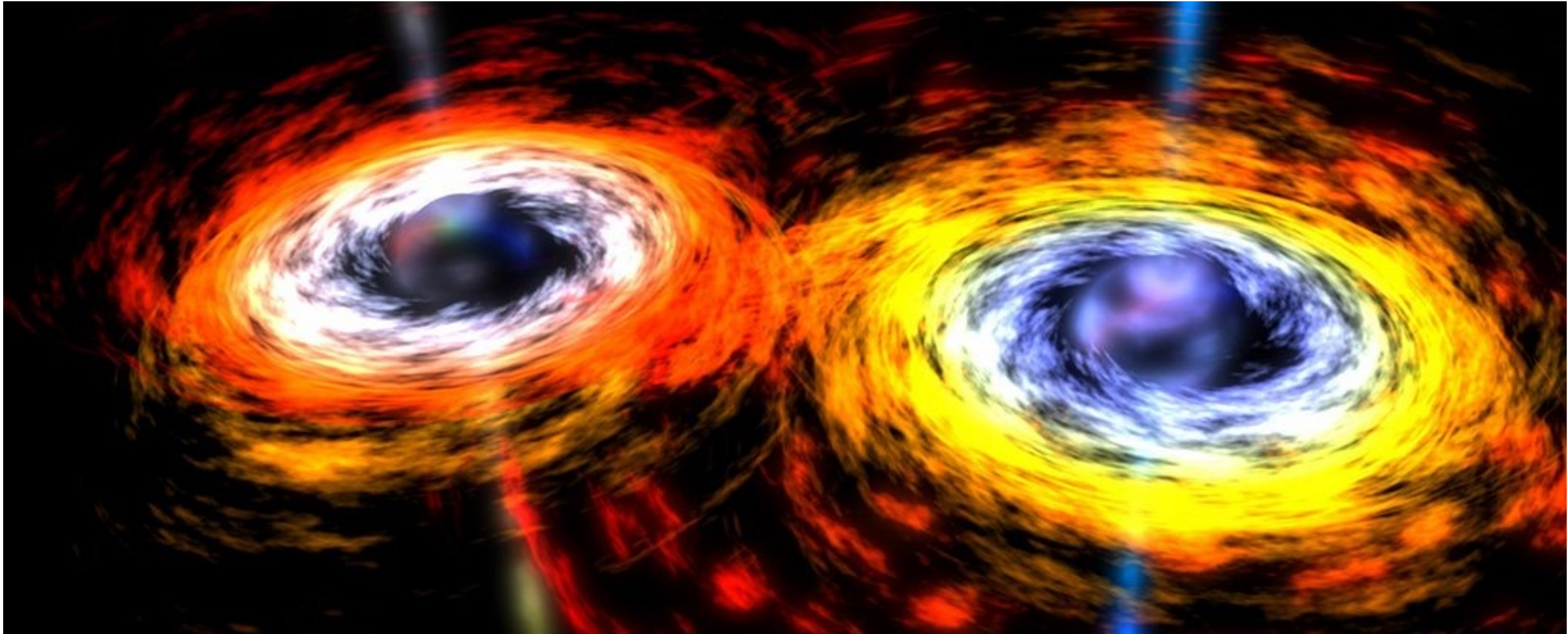




IMAGES NOT TO SCALE



Gravitational Waves and ...



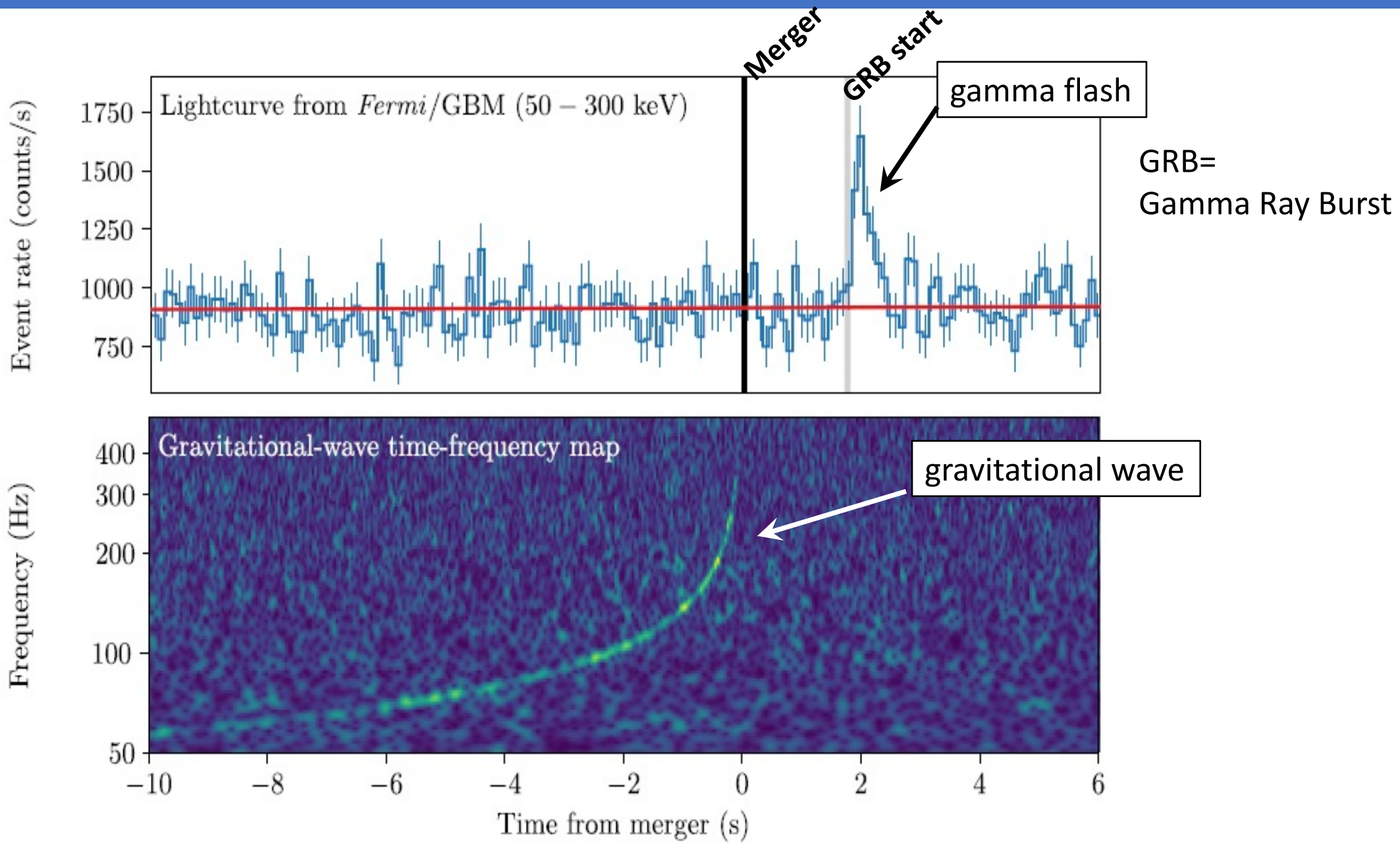


Gamma flash 1.7 sec later...



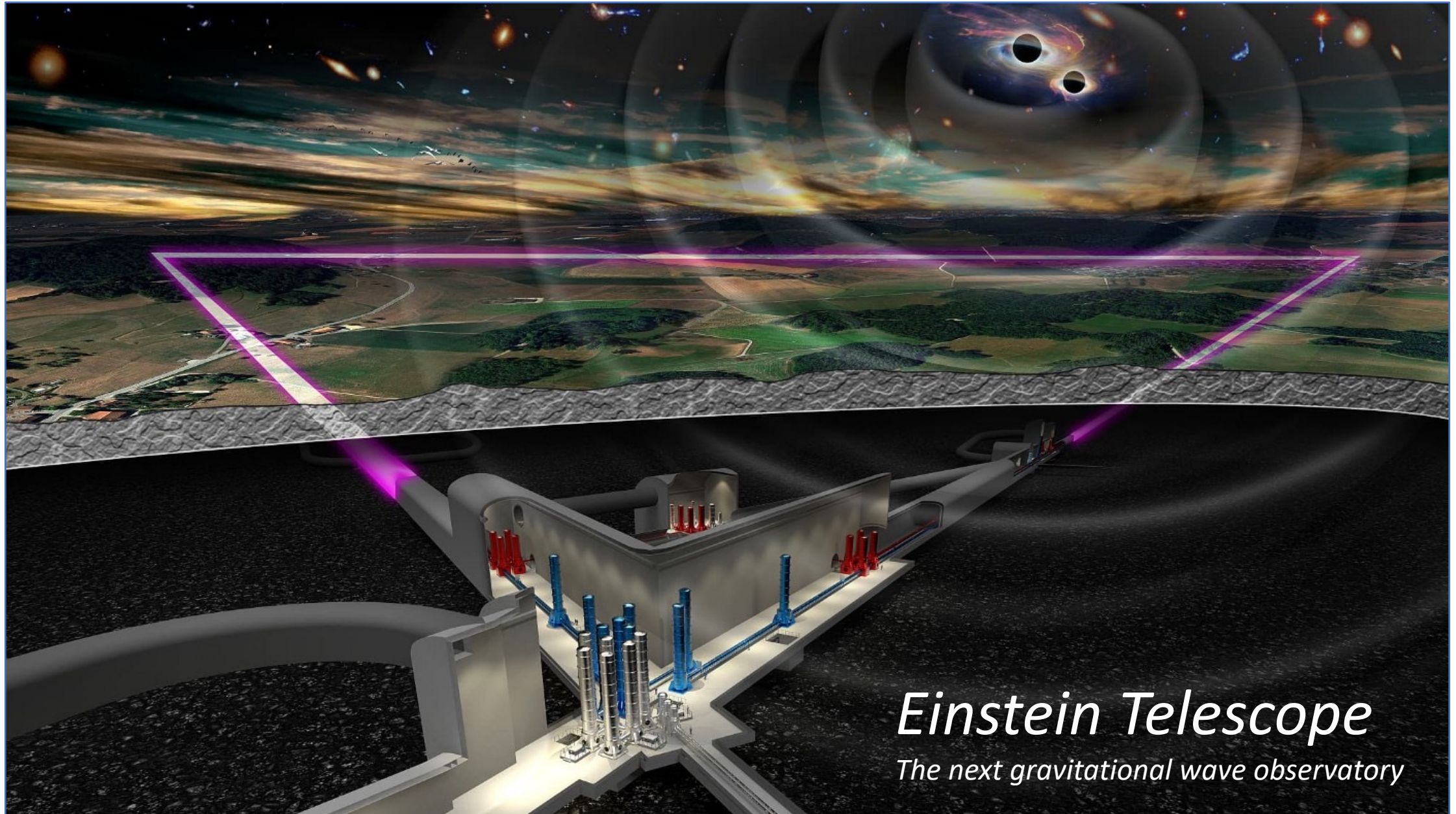
INTEGRAL

Fermi Space Telescope

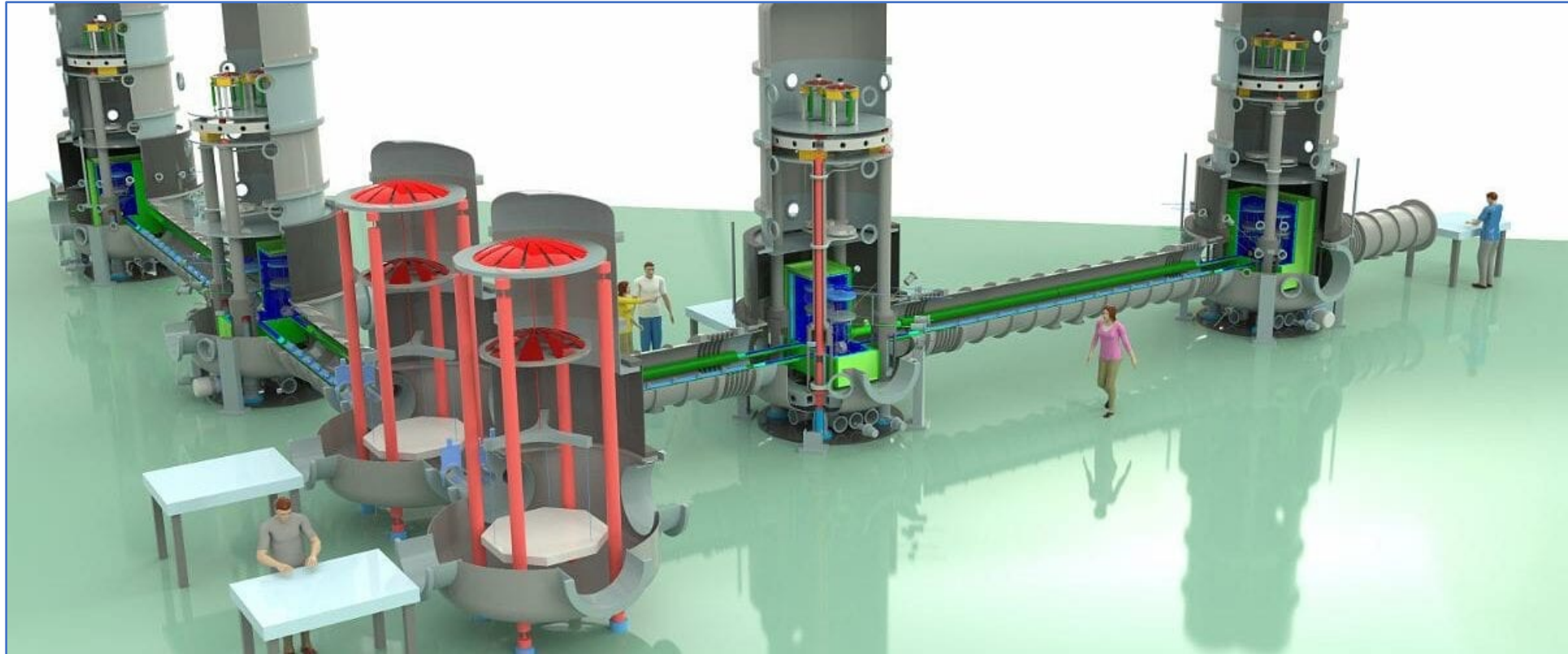




Kilonova: production heavy elements: gold, platinum etc.



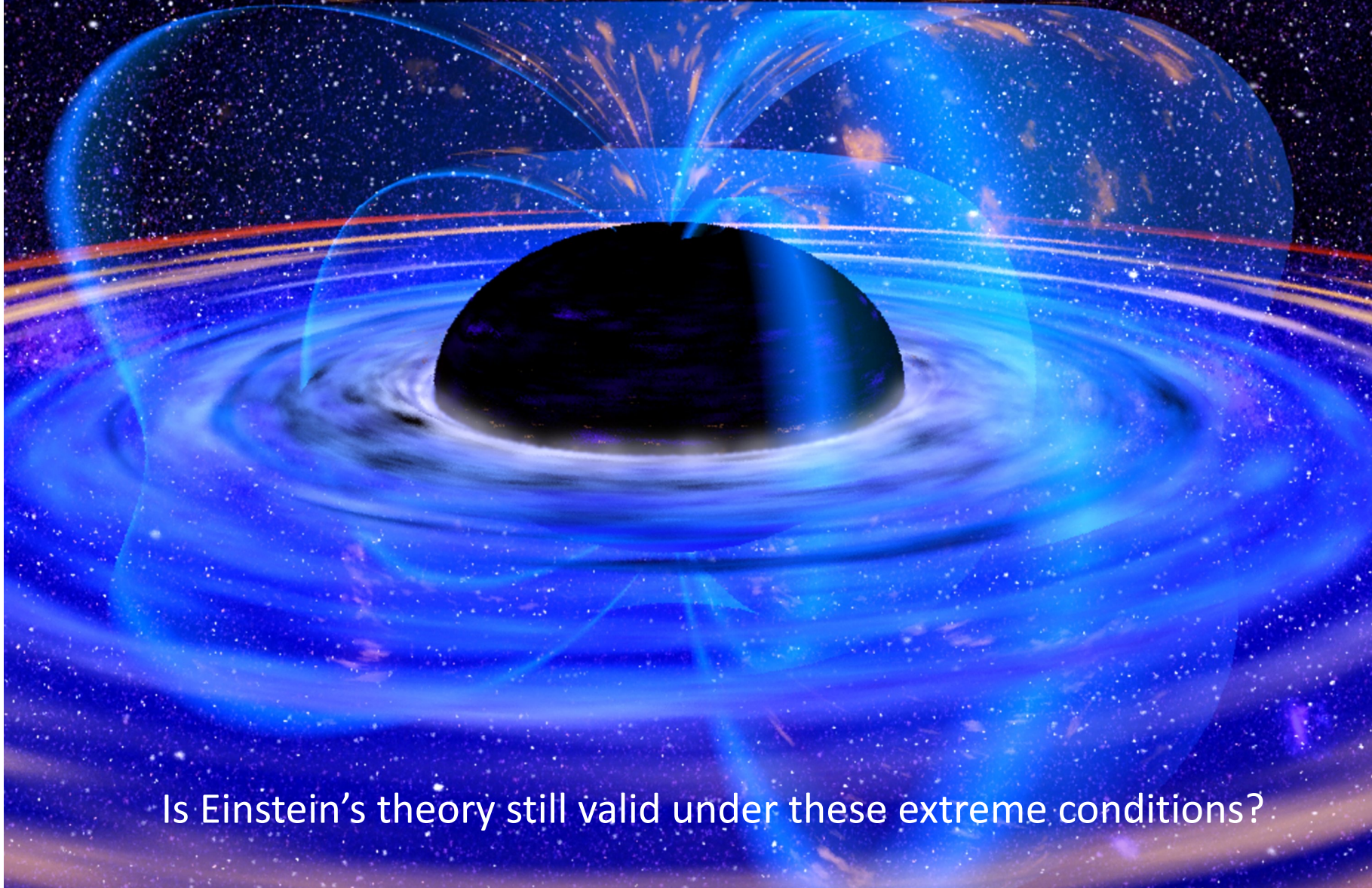




Nikhef, RWTH Aachen, UCL Louvain, Hasselt, Ghent, Antwerp, VUB Brussels, ULB Brussels, Liege, Radboud University Nijmegen, TU Eindhoven and Hamburg

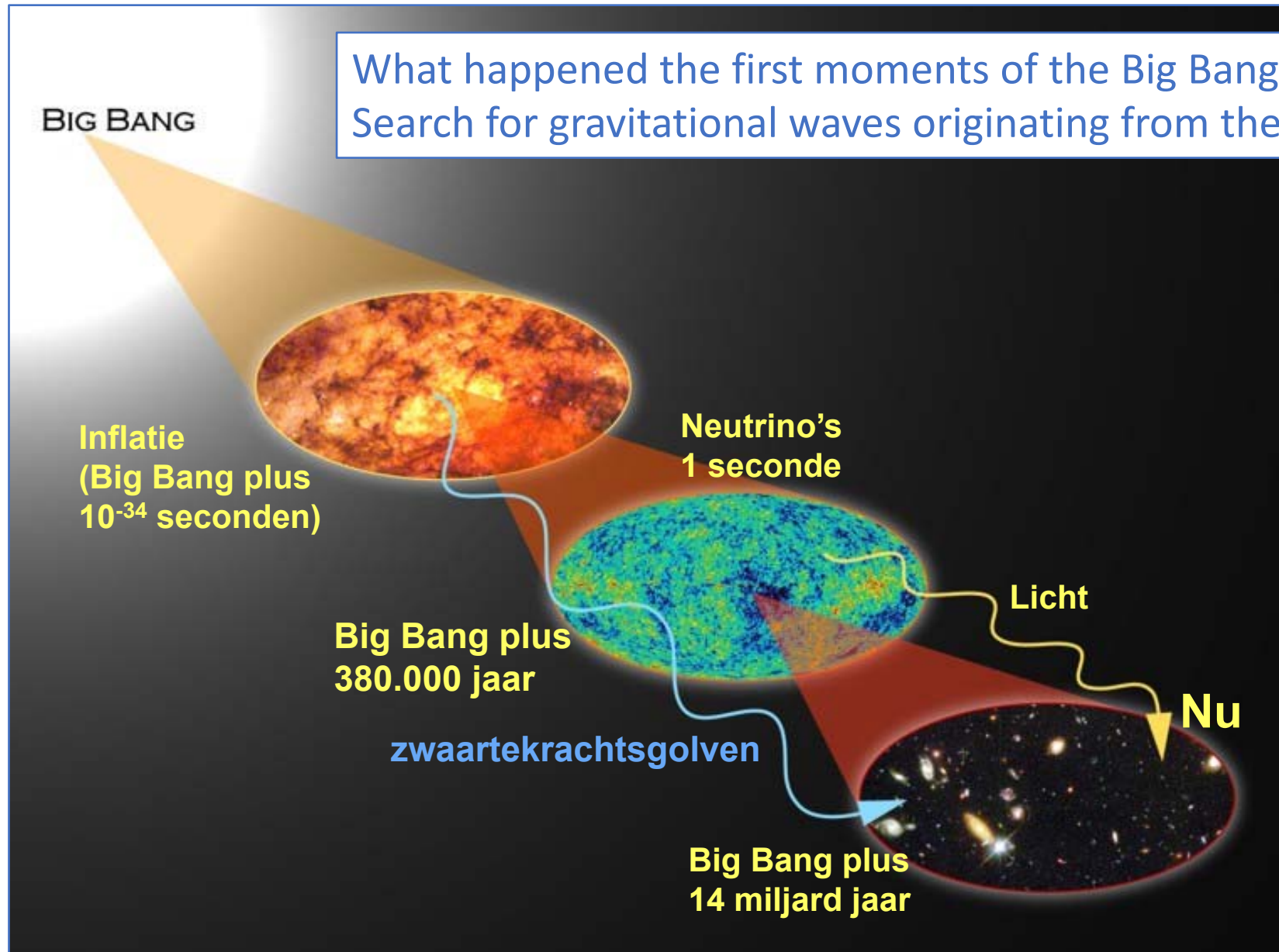


Wat happens at the edge of a black hole (quantum effects)?

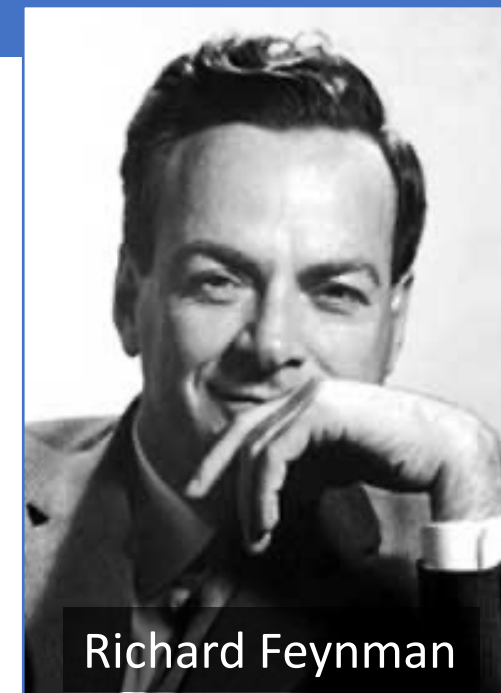
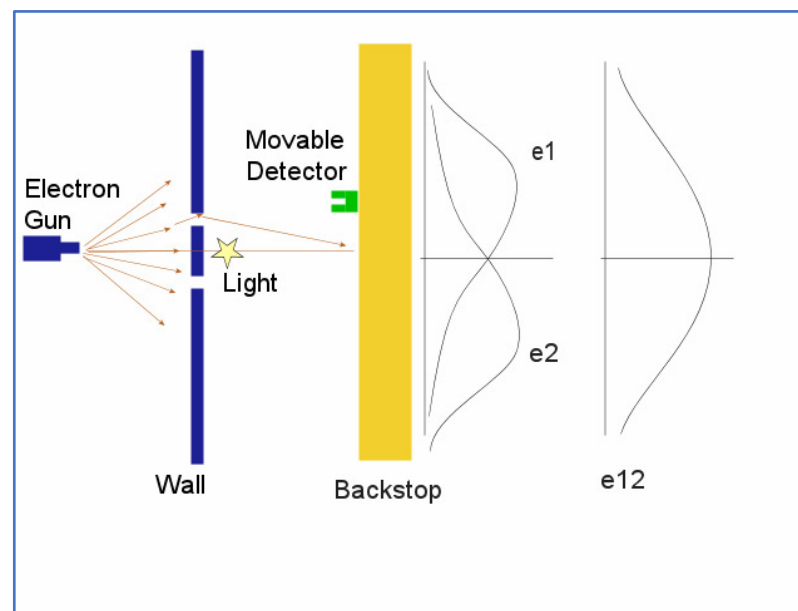
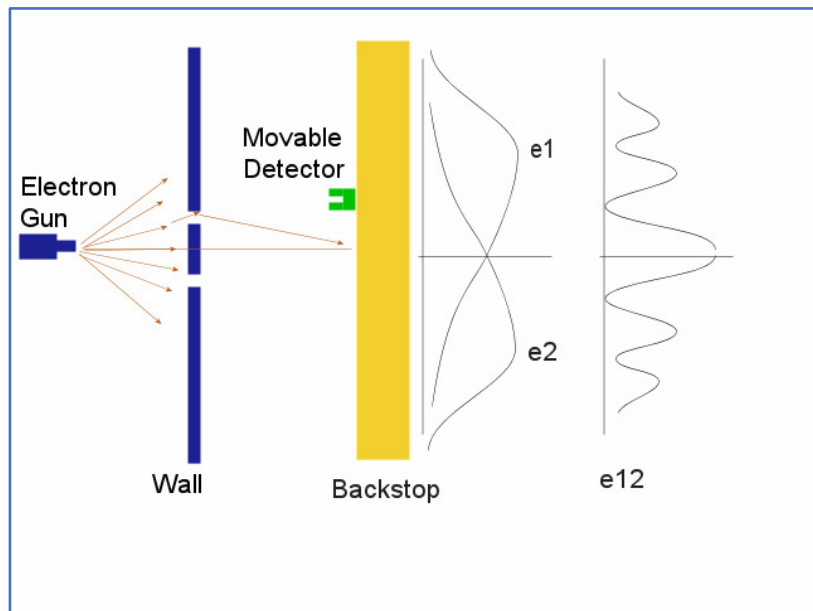


Is Einstein's theory still valid under these extreme conditions?

What happened the first moments of the Big Bang?
Search for gravitational waves originating from the very first moment!



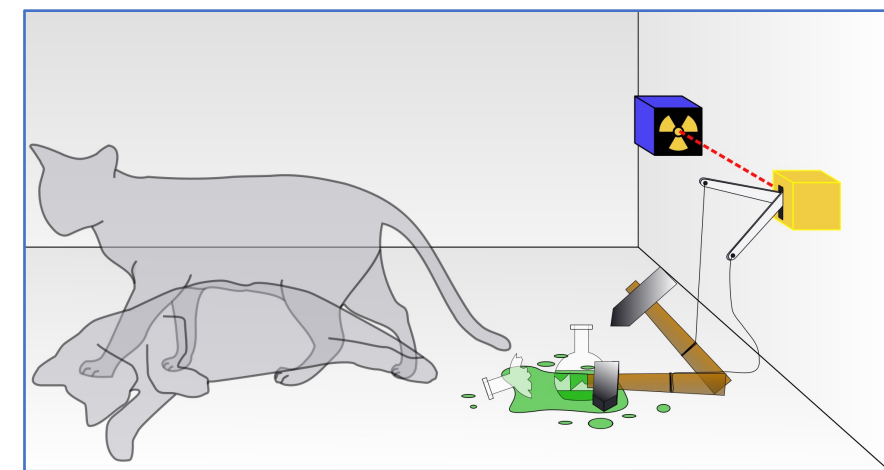
Next week: Quantum Mechanics



Richard Feynman

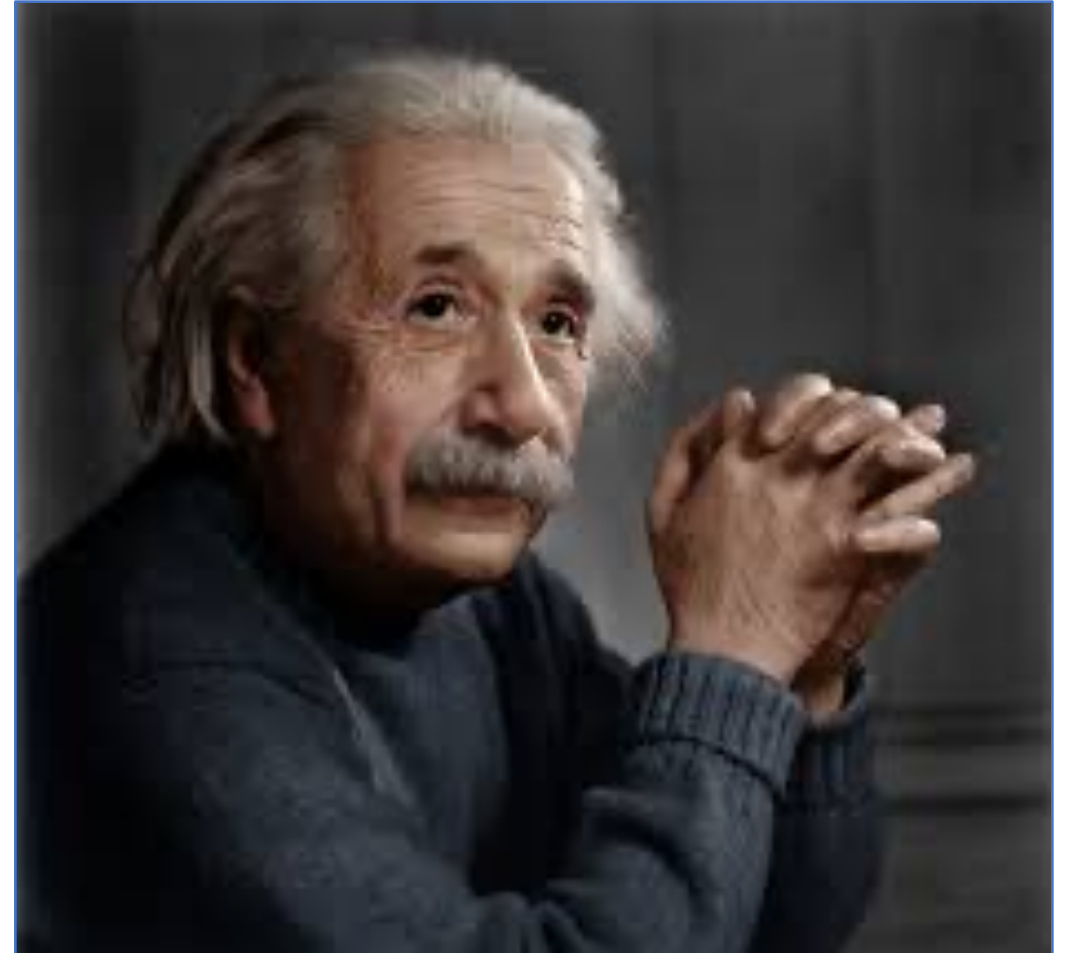
Quantum mechanics developed by Bohr and Heisenberg leads to "absurd" thought experiments of Feynman and Wheeler. Einstein and Schrödinger did not like it.

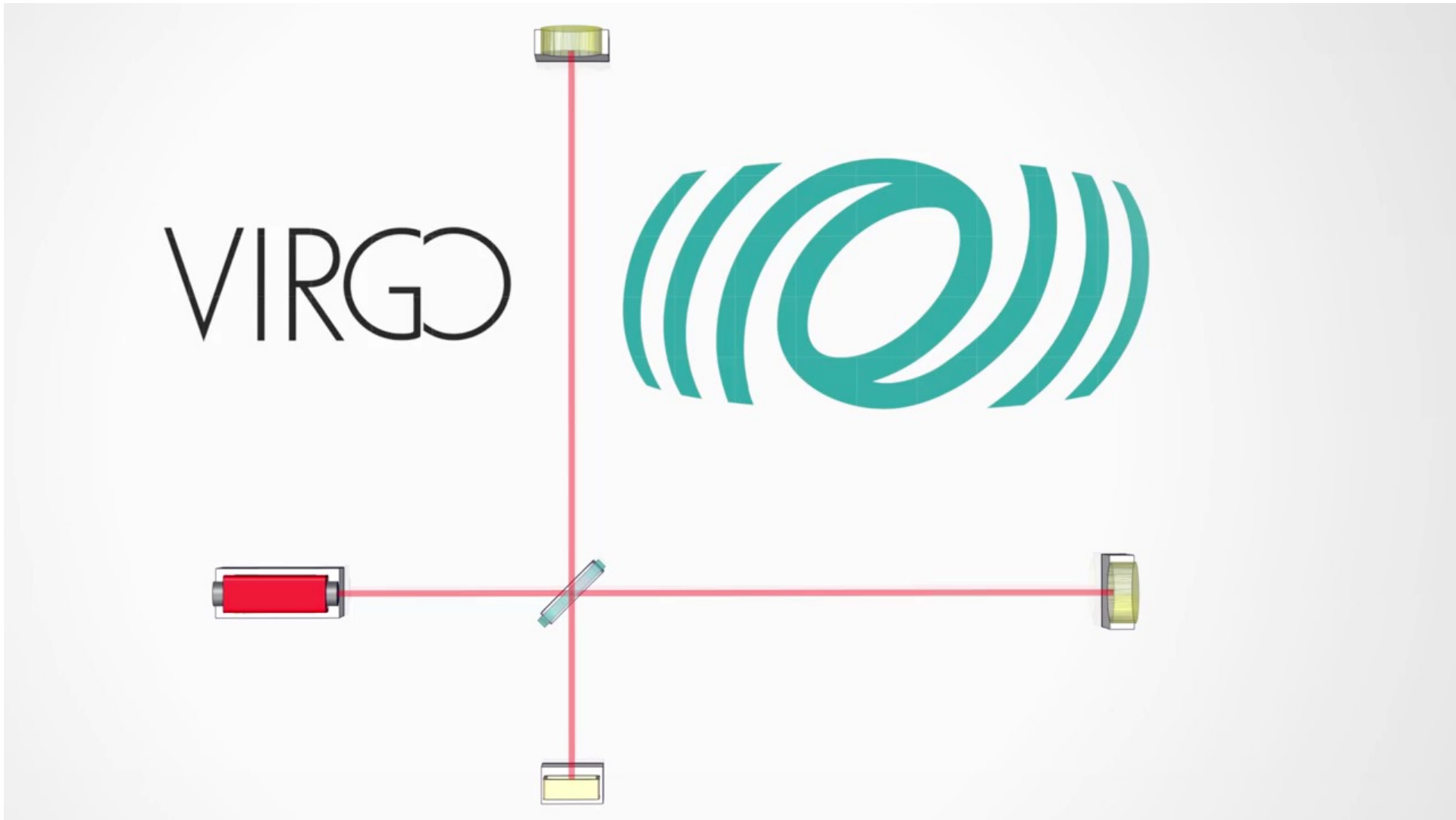
Even today people are debating its interpretation....



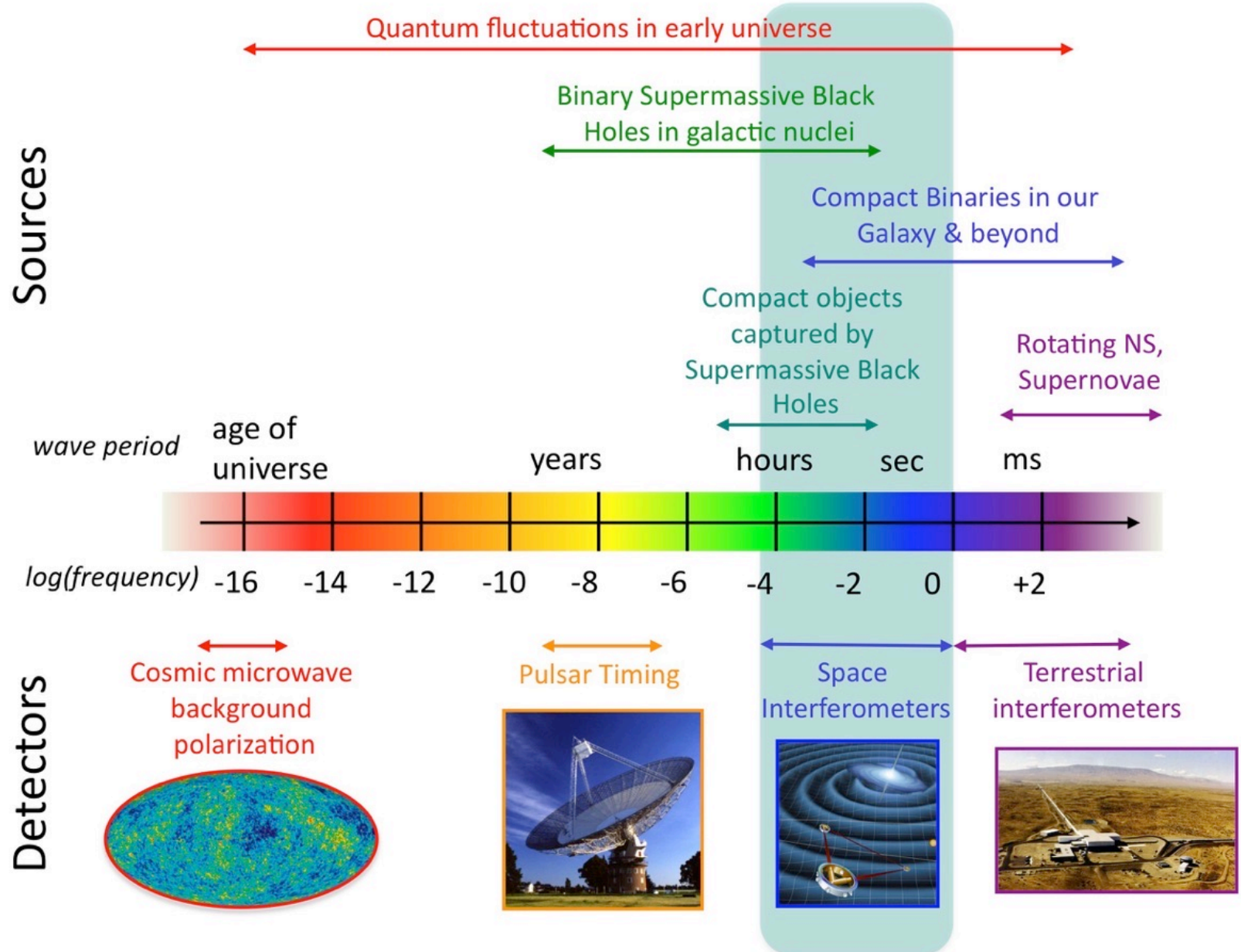
Extra Slides

- “Imagination is more important than knowledge”
- “Education is what remains after one has forgotten what one has learned at school.”
- “I fear the day that technology will surpass our human interaction. The world will have a generation of idiots.”
- “A person who never made a mistake never tried anything new.”





The Gravitational Wave Spectrum



$\frac{1}{2}$: Roger Penrose



For the discovery that black hole formation is a robust prediction of the general theory of relativity.

$\frac{1}{4}$: Reinhard Genzel



For the discovery of a supermassive compact object at the centre of our galaxy.

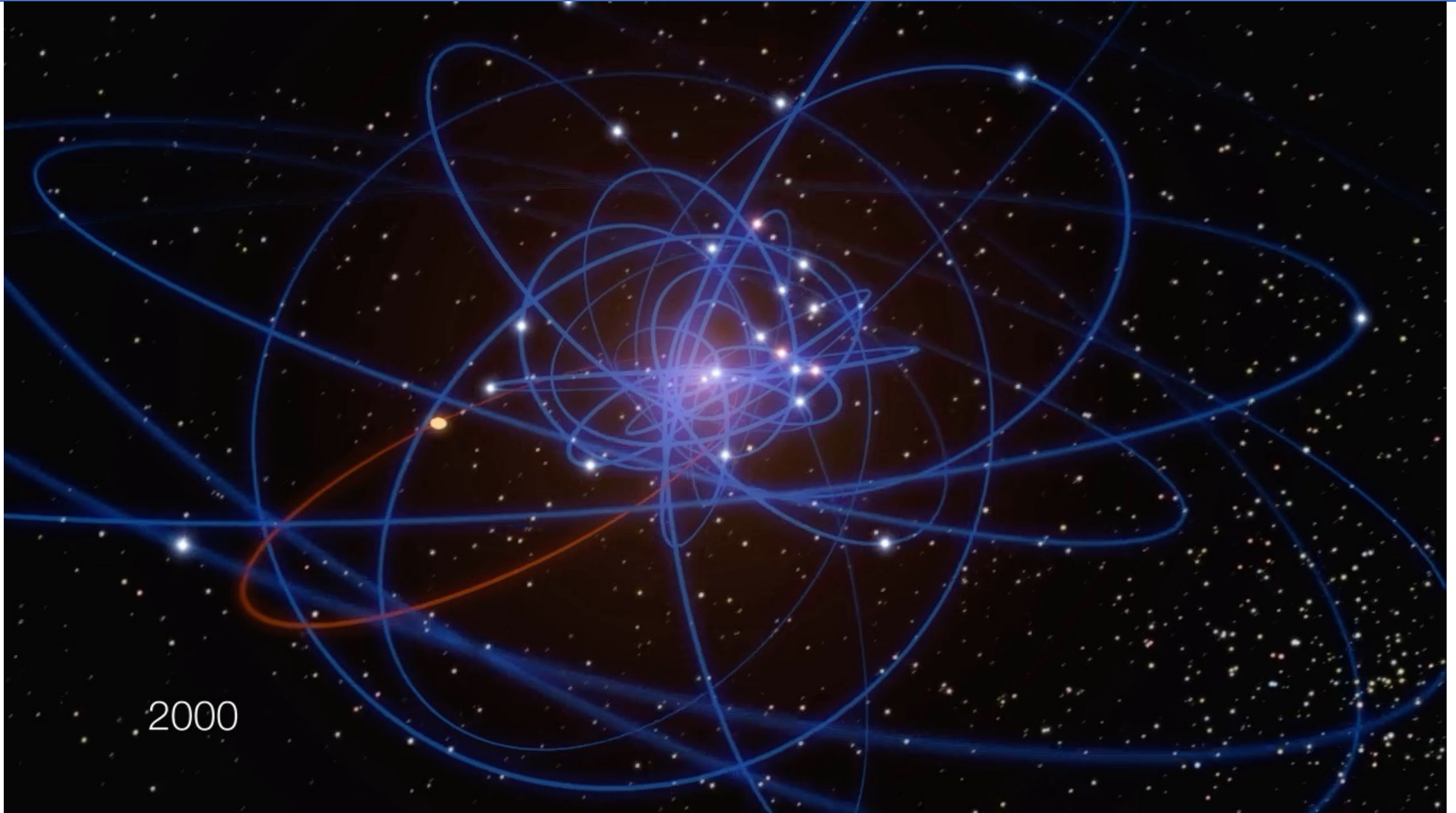
$\frac{1}{4}$: Andrea Ghez



Center of our Milky Way Galaxy



Supermassive Black Hole in the center of our Galaxy



2000