

#### PHY3004: Nuclear and Particle Physics Marcel Merk, Jacco de Vries



The Standard Model



# <u>Recap</u>: "Seeing the wood for the trees"

#### The Standard Model

- Lecture 1: "Particles"
  - Zooming into constituents of matter
  - Skills: distinguish particle types, Spin
- Lecture 2: "Forces"
  - Exchange of quanta: EM, Weak, QCD
  - Skills: 4-vectors, Feynman diagrams
- Lecture 3: "Waves"
  - Quantum fields and gauge invariance
  - Dirac algebra, Lagrangian, co- & contra variant
- Lecture 4: "Symmetries"
  - Standard Model, Higgs, Discrete Symmetries
  - Skills: Lagrangians, Chirality & Helicity
- Lecture 5: "Scattering"
  - Cross section, decay, perturbation theory
  - Skills: Dirac-delta function, Feynman Calculus
- Lecture 6: "Detectors"
  - Energy loss mechanisms, detection technologies











# Lecture 1: "Particles"

#### **Classification of particles**

- Lepton: fundamental particle
- Hadron: consist of quarks
  - Meson: 1 quark + 1 antiquark ( $\pi^+$ , $B_s^0$ , ...)
  - Baryon: 3 quarks (*p* ,*n* , Λ, ...)
    - Anti-baryon: 3 anti-quarks

#### • Fermion: particle with half-integer spin.

- Antisymmetric wave function: obeys Pauliexclusion principle and Pauli-Dirac statistics
- All fundamental quarks and leptons are spin-1/2
- Baryons (S=1/2, 3/2)
- Boson: particle with integer spin
  - Symmetric wave function: Bose-Einstein statistics
  - Mesons: (S=0, 1), Higgs (S=0)
  - Force carriers: *γ*, *W*, *Z*, *g* (S=1); graviton(S=2)



#### **Standard Model of Elementary Particles**

#### Lecture 2: "Forces"

Griffiths chapter 2



Lecture 3: "W		Probability interpretation (Continuity equation)		
Quantum Mechanics:	$E \rightarrow \hat{E} = i\hbar \frac{\partial}{\partial t}$ ; $p \rightarrow \hat{p}$	$=-i\hbarec{ abla}$		$\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot \vec{j} = 0$
Non-relativistic spin 0: $E = \frac{\vec{p}^2}{2m}$	Schrödinger: $i\hbar \frac{\partial}{\partial t} \psi = -\frac{\hbar^2}{2m} \nabla^2 \psi$	$\psi = Ne^{i(\vec{p}\vec{x} - Et)}$	$\rho \equiv \vec{j} \equiv$	$\equiv \psi^* \psi =  N ^2$ $\equiv \frac{i\hbar}{2m} \left( \psi \vec{\nabla} \psi^* - \psi^* \vec{\nabla} \psi \right) = \frac{ N ^2}{m} \vec{p}$
Relativistic spin 0: $E^2 = p^2c^2 + m^2c^4$ Example: pions	Klein-Gordon: $-\frac{1}{c^2}\frac{\partial^2}{\partial t^2}\phi = -\nabla^2\phi + \frac{m}{d^2}\phi$ $\frac{\partial_\mu}{\partial^\mu}\phi + m^2\phi = 0$	$\frac{\hbar^2 c^2}{\hbar^2} \phi$ $\phi = N e^{i(\vec{p}\vec{x} - Et)}$	$j^{\mu}(p = \vec{\rho} = \vec{j} = \vec{j}$	$\begin{aligned} f(\rho, \vec{j}) &= i[\phi^*(\partial^\mu \phi) - \phi(\partial^\mu \phi^*)] \\ &= 2 N ^2 E \\ &= 2 N ^2 \vec{p} \end{aligned} \qquad j^\mu = 2 N ^2 p^\mu \end{aligned}$
Relativistic spin- ½: $H = (\vec{\alpha} \cdot \vec{p} + \beta m)$ Fundamental quarks and leptons	Dirac: $i \frac{\partial}{\partial t} \psi = (-i \vec{\alpha} \cdot \vec{\nabla} + \beta m) \psi$ $(i \gamma^{\mu} \partial_{\mu} - m) \psi = 0$ $\gamma^{\mu} = (\beta m) \psi$	$\psi = u(p)e^{i(\vec{p}\vec{x}-Et)}$ $u(p) = \left( \begin{array}{c} \cdot \\ \cdot \\ \cdot \\ \cdot \\ \cdot \end{array} \right)$	<b>j</b> 0	$j^{\mu} = \bar{\psi}\gamma^{\mu}\psi$ $= \bar{\psi}\gamma^{0}\psi = \psi^{\dagger}\psi = \sum_{i=1}^{4}  \psi_{i} ^{2}$
Relativistic spin-1: Fundamental force carriers	Proca: $\partial_{\mu}\partial^{\mu}A^{\nu} + m^{2}A^{\nu} = j^{\nu}$	EM: $A^{\mu} = \gamma \rightarrow m = 0$ QCD: $A^{\mu} = g \rightarrow m = 0$ Weak: $A^{\mu} = W, Z \rightarrow m \neq 0$		EM: Maxwell equations for $\vec{E}$ and $\vec{B}$ fields

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#### Lecture 3: "Waves" – gauge invariance $S = \int d^4x \, \mathcal{L}\big(\phi(x), \partial \phi(x)\big)$ Spin 0 Scalar field: $\mathcal{L} = \frac{1}{2} (\partial_{\mu} \phi) (\partial^{\mu} \phi) - \frac{1}{2} m^2 \phi^2$ Lagrangians: $\delta S = 0$ Spin ½ Dirac fermion $\mathcal{L} = i \bar{\psi} \gamma_{\mu} \partial^{\mu} \psi - m \bar{\psi} \psi$ Spin 1 gauge boson (photon) : $\mathcal{L} = -\frac{1}{4} (\partial^{\mu} A^{\nu} - \partial^{\nu} A^{\mu}) (\partial_{\mu} A_{\nu} - \partial_{\nu} A_{\mu}) - j^{\mu} A_{\mu}$ $\frac{\partial \mathcal{L}}{\partial \phi(x)} = \partial_{\mu} \frac{\partial \mathcal{L}}{\partial \left(\partial_{\mu} \phi(x)\right)}$ (stationary action in field theory) Euler Lagrange lead to the wave equations: All forces result from requiring a symmetry principle: Lagrangian should stay invariant under transformations $\mathcal{L} = i\bar{\psi}\gamma_{\mu}\partial^{\mu}\psi - m\bar{\psi}\psi \longrightarrow \mathcal{L} = i\bar{\psi}\gamma_{\mu}D^{\mu}\psi - m\bar{\psi}\psi$ 1) QED = U(1) symmetry $\psi(x) \rightarrow \psi'(x) = e^{iq\alpha(x)}\psi(x)$ Covariant derivative: $\partial^{\mu} \rightarrow D^{\mu} \equiv \partial^{\mu} + iqA^{\mu}$ $\mathcal{L} = i\overline{\psi}\gamma_{\mu} \,\partial^{\mu}\psi - m\overline{\psi}\psi - q\overline{\psi}\gamma_{\mu}\psi A^{\mu}$ "free" "interaction" $\begin{pmatrix}\psi_{r}\\\psi_{g}\\\psi_{b}\end{pmatrix}$ $\psi(x) \rightarrow \psi'(x) = \exp\left(\frac{i}{2}g_{s}\vec{\lambda}\cdot\vec{\alpha}(x)\right)\begin{pmatrix}\psi_{r}\\\psi_{g}\\\psi_{b}\end{pmatrix}$ $A^{\mu}(x) \rightarrow A'^{\mu}(x) = A^{\mu}(x) - \partial^{\mu}\alpha(x)$ $\rightarrow$ 1 E.M. photon field: $A^{\mu}(x)$ 2) Weak = SU(2) symmetry $\psi = \begin{pmatrix} \psi_u \\ \psi_d \end{pmatrix}$ $\psi(x) \to \psi'(x) = \exp\left(\frac{i}{2}g\vec{\tau}\cdot\vec{\alpha}(x)\right)\begin{pmatrix}\psi_u\\\psi_d\end{pmatrix}$ → 8 colored gluon fields: $g^{\mu}(x)$ $\rightarrow$ 3 weak fields: $W^{\mu+}(x)$ , $W^{\mu-}(x)$ , $Z^{\mu}(x)$

### Lecture 4: "Symmetries" – Standard Model

- The Lagrangian of the Standard Model includes electromagnetic, weak and strong interactions according to the gauge field principle
- Construction of the Lagrangian:  $\mathcal{L} = \mathcal{L}_{\underline{free}} \mathcal{L}_{\underline{interaction}} = \mathcal{L}_{\underline{Dirac}} gJ^{\mu}A_{\mu}$ 
  - With g a coupling constant,  $J^{\mu}$  a current  $(\overline{\psi}O_{i}\psi)$  and  $A_{\mu}$  a force field
  - A. Local U(1) gauge invariance: symmetry under complex phase rotations
    - Conserved quantum number: (hyper-) charge

• Lagrangian: 
$$\mathcal{L} = \bar{\psi}(i\gamma^{\mu}D_{\mu} - m)\psi = \bar{\psi}(i\gamma^{\mu}\partial_{\mu} - m)\psi - q\,\bar{\psi}\gamma^{\mu}\psi\,A_{\mu}$$
  $(\partial_{\mu} \to D_{\mu} \equiv \partial_{\mu} + iqA_{\mu})$   
 $\underbrace{\mathcal{J}_{EM}^{\mu}}$ 

- B. Local SU(2) gauge invariance: symmetry under transformations in isospin doublet space.
  - Conserved quantum number: weak isospin

• Lagrangian: 
$$\mathcal{L} = \overline{\Psi}(i\gamma^{\mu}D_{\mu} - m)\Psi = \overline{\Psi}(i\gamma^{\mu}\partial_{\mu} - m)\Psi - \frac{g}{2}\overline{\Psi}\gamma^{\mu}\vec{\tau}\Psi\vec{b}_{\mu}$$
  $(I\partial_{\mu} \to D_{\mu} = I\partial_{\mu} + igB_{\mu})$   
 $B_{\mu} = \frac{1}{2}\vec{\tau}\cdot\vec{b}_{\mu} = \frac{1}{2}\tau_{1}^{a}b_{\mu}^{a} = \frac{1}{2}\begin{pmatrix}b_{3} & b_{1} - ib_{2}\\b_{1} + ib_{2} & -b_{3}\end{pmatrix}$   $\underbrace{J_{Weak}^{\mu}}$ 

- C. Local SU(3) gauge invariance: symmetry under transformations in colour triplet space
  - Conserved quantum number: color
  - Lagrangian:  $\mathcal{L} = \overline{\Phi}(i\gamma^{\mu}D_{\mu} m)\Phi = \overline{\Phi}(i\gamma^{\mu}\partial_{\mu} m)\Phi \frac{g_s}{2}\overline{\Phi}\gamma^{\mu}\vec{\lambda}\Phi\vec{c}_{\mu}$   $(I\partial_{\mu} \to D_{\mu} = I\partial_{\mu} + ig_sC_{\mu})$  $C_{\mu}$  are 3x3 matrices  $\rightarrow$  gluon fields  $\mathcal{J}_{QCD}^{\mu}$

#### Lecture 4: "Symmetries" – Standard Model



#### Lecture 4: Electroweak Quantum Numbers

For	weak isospin s	some peop	le write $T_3$	while oth	ers write $I_3$	With	$n: Q = T_3 +$	$\frac{1}{2}Y$
	Generation						$: Q = I_3 +$	$\frac{1}{2}Y$
		Ι	II	III	$I_3$	Y	Q	
	Leptons	$\begin{pmatrix} v_e \\ e \\ e \end{pmatrix}_L \\ e_R$	$ \begin{pmatrix} \nu_{\mu} \\ \mu \end{pmatrix}_{L} \\ \mu_{R}$	$\begin{pmatrix} \nu_{\tau} \\ \tau \end{pmatrix}_{L} \\ \tau_{R}$	$+1/2 \\ -1/2 \\ 0$	$-1 \\ -1 \\ -2$	$0 \\ -1 \\ -1$	
	Quarks	$\begin{pmatrix} u \\ d \end{pmatrix}_{L} \\ u_{R} \\ d_{R} \end{pmatrix}$	$\begin{pmatrix} c \\ s \end{pmatrix}_L \\ c_R \\ s_R \end{pmatrix}$	$\begin{pmatrix} t \\ b \end{pmatrix}_{L} \\ t_{R} \\ b_{R}$	$+1/2 \\ -1/2 \\ 0 \\ 0$	+1/3 +1/3 +4/3 -2/3	+2/3 -1/3 +2/3 -1/3	

# Symmetry breaking with a *real* field $\phi$

• Explicit mass terms violate the symmetry:  $m^2 A_{\mu} A^{\mu} \rightarrow m^2 \left(A_{\mu} + \frac{1}{2} \partial_{\mu} \alpha\right) \left(A^{\mu} + \frac{1}{2} \partial^{\mu} \alpha\right) \neq m^2 A_{\mu} A^{\mu}$ Note that  $\mathcal{L}$  is symmetric under  $\phi \rightarrow -\phi$  and



minimum in the Lagrangian. We can investigate  $\mu^2 < 0$ :  $\mathcal{L} = \frac{1}{2} \left( \partial_{\mu} \phi \right)^2 - V(\phi) = \frac{1}{2} \left( g_{\mu} \phi \right)^2 \int -\mu \frac{1}{2} \mu^2 \phi \partial \sin \frac{1}{4} \psi \phi \partial r \operatorname{nega} \right)^2 \cdot \frac{1}{2} \left( \partial_{\mu} \phi \right)^2 \int -\frac{1}{2} \left$ 

Massive Klein-Gordon term (Spin 0, mass =
$$\mu$$
)  $0$  term  $Tee pa$ 

The Lagrangian has a minimum for  $\phi_0 = \sqrt{\frac{\mu^2}{\lambda}} = v$  or  $\mu^2 = -\lambda v^2$ 

#### **Conclusion:**

- The symmetry of the Lagrang (10) adding a symmetric potential  $\phi$  has not been destroyed
- The vacuum is no longer in a symmetric position

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 $V(\phi)$ 

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a

Griffiths §10.9

## Symmetry breaking with a *complex* field $\phi$

- Introduce a complex scalar field:  $\phi = \frac{1}{\sqrt{2}}(\phi_1 + i\phi_2)$
- The Lagrangian term is:  $\mathcal{L} = (\partial_{\mu}\phi)^*(\partial^{\mu}\phi) V(\phi)$ , with  $V(\phi) = \mu^2(\phi^*\phi) + \lambda (\phi^*\phi)^2$
- Lagrangian:

$$\begin{aligned} \mathcal{L}(\phi_1,\phi_2) &= \frac{1}{2} \left( \partial_\mu \phi_1 \right)^2 + \frac{1}{2} \left( \partial_\mu \phi_2 \right)^2 \\ &- \frac{1}{2} \mu^2 (\phi_1^2 + \phi_2^2) - \frac{1}{4} \lambda (\phi_1^2 + \phi_2^2) \end{aligned}$$

#### **Conclusion:**

- The symmetry of the Lagrangian by adding a symmetric potential  $\phi$  has not been destroyed
- The *vacuum is no longer* in a symmetric position

The real case includes a complex (isospin doublet) field  $\phi$ -  $\phi$  degrees of freedom lead to mass terms for the  $W^+, W^-, Z^0$ -  $\phi$  can also couple to fermions  $\rightarrow$  particle masses



Griffiths §10.9

### Lecture 5 : "Scattering" – non-Rel. 1) Scattering in external potential







### Lecture 5 : "Scattering" – non-Relativistic







### Lecture 5 : "Scattering" - Relativistic

Cross section:



Example diagram:

From ABC  $\rightarrow$  Standard Model theory:

More complicated rules and spin objects

A

 $p_2$ 

 $A + A \rightarrow B + B$ 

B

В

How to determine  $\mathcal{M}$ ?  $\rightarrow$  Feynman rules (depend on actual theory/interaction):

Feynman rules (ABC theory):

- 1. Diagram: see sketch
- 2. Labels: see sketch
- 3. Two vertices:  $(-ig)^2 = -g^2$
- 4. Propagators: one internal line:  $\frac{i}{q^2 m_c^2} \rightarrow \text{Master level education}$
- 5. Conservation of *E*,  $\vec{p}$  twice:  $(2\pi)^4 \, \delta^4(\vec{p}_1 p_3 q)$  and  $(2\pi)^4 \, \delta^4(p_2 + q p_4)$
- 6. Integrate: one integral:  $1/(2\pi)^4 d^4 q$
- 7. Erase delta-function and multiply by *i* to find:  $\mathcal{M} = \frac{g^2}{(n_i n_i)^2 m^2}$



#### $A + A \rightarrow B + B$ Scattering: $d\sigma/d\Omega$ $p_4$ $A p_1$ $p_2 \quad A$ After Look at the matrix element and assume that Look at the matrix element and assume that $m_A = m_B = m \text{ and } m_C = 0 \text{ (eg. a photon):}$ $\mathcal{M} = \frac{g^2}{(p_2 - p_2)^2} + \frac{g^2}{(p_4 - p_2)^2}$ $(p_4 - p_2)^2 - m_c^2 = p_4^2 + p_2^2 - 2p_2 \cdot p_4$ Note that for 4-vectors: $= m_4^2 + m_2^2 - 2p_2 \cdot p_4$ $p_i \cdot p_j = p_{i_{\mu}} p_j^{\mu} = E_i E_j - \vec{p}_i \cdot \vec{p}_j$ $= 2m^2 - 2E_2E_4 + 2(\vec{p}_2 \cdot \vec{p}_4)$ and that $p^2 = p_{\mu}p^{\mu} = E^2 - \vec{p}^2 = m^2$ $= 2m^{2} - 2\left(\sqrt{m^{2} + \vec{p}^{2}}\right)\left(\sqrt{m^{2} + \vec{p}^{2}}\right) + 2\vec{p}^{2}\cos\theta$ (Invariant mass) $= -2\vec{p}^2(1-\cos\theta)$ • Plug in: (S = 1/2) $(p_3 - p_2)^2 - m_c^2 = -2\vec{p}^2(1 + \cos\theta)$ $\frac{d\sigma}{d\Omega} = \left(\frac{1}{8\pi}\right)^2 \frac{S|\mathcal{M}|^2}{(E_1 + E_2)^2} \frac{\left|\vec{p}_f\right|}{\left|\vec{p}_i\right|}$ $\mathcal{M} = \frac{g^2}{-2\vec{p}^2(1-\cos\theta)} + \frac{g^2}{-2\vec{p}^2(1+\cos\theta)} = -\frac{g^2}{2\vec{p}^2\sin^2\theta}$ $\frac{d\sigma}{d\Omega} = \frac{1}{2} \left( \frac{g^2}{16\pi E \vec{p}^2 \sin^2 \theta} \right)^{-1}$

#### Lecture 5 : Towards Experimental

- Consider beam of particles on a target
  - Luminosity  $\mathcal{L}$  is number of particles per unit time, per unit area.
  - Number of particles passing through area  $d\sigma$ :  $dN = \mathcal{L} d\sigma$
  - Number of particles scattering into solid angle  $d\Omega : dN = \mathcal{L} d\sigma = \mathcal{L} D(\theta) d\Omega$
  - By counting one can measure the *differential cross section*:  $\frac{d\sigma}{d\Omega} = D(\theta) = \frac{dN}{\mathcal{L} d\Omega}$
  - Alternatively the total cross section:  $N = \mathcal{L} \sigma$

These aspects are needed when you Compare theory with experiments.

- Experimental particle physics:
  - Measure number of events N and the luminosity  $\mathcal{L}$  to find cross section  $\sigma = N/\mathcal{L}$

dΩ-

• Compare with theoretical calculation of  $\sigma$  (or  $\frac{d\sigma}{d\Omega}$ ) using e.g. Standard Model

#### Lecture 6 : "Detectors" - Technologies







#### **Cherenkov Detectors**







# Lecture 6 : "Detectors" - Combination

- Charged particle tracking:
  - Position & Momentum measurement
- Particle Identification:
  - Energy, Time-of-Flight, Cherenkov



#### Skill: Four vectors & co- and contra-variance

- Four vector:  $x^{\mu} = (x^0, x^1, x^2, x^3)$  with  $x^0 = ct \Rightarrow x^0 = t$  (since  $c \equiv 1$ )
- We call this a *contravariant* vector and:  $x^{\mu} = (x^0, \vec{x})$
- Lorentz transformation:

Lorentz transformation:  

$$x^{\mu'} = \Lambda^{\mu}_{\nu} x^{\nu} ; \Lambda^{\mu}_{\nu} = \begin{pmatrix} \gamma & -\beta & 0 & 0 \\ -\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \qquad x^{1'} = \gamma(x^{1} - \beta x^{0}) \qquad g_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}$$

- Lorentz transformation leaves the following invariant:  $|x|^2 = x^{0^2} |\vec{x}|^2$ •  $|x|^2 = x^{0^2} - |\vec{x}|^2 = (ct)^2 - |\vec{x}|^2 = (ct')^2 - |\vec{x}'|^2$
- Introduce *covariant* vectors  $x_{\mu} = \sum_{\mu} g_{\mu\nu} x^{\nu} = g_{\mu\nu} x^{\nu}$

Note the Einstein summation convention

- Inproduct invariants:  $I = a_{\mu}b^{\mu} = a \cdot b = a'_{\mu}b'^{\mu}$  for any Lorentz 4-vectors  $a^{\mu}$  and  $b^{\mu}$ 
  - Example invariant mass:  $E^2 = \vec{p}^2 c^2 + m^2 c^4 \Rightarrow p^{\mu} = (E, \vec{p}) \Rightarrow p_{\mu} p^{\mu} = E^2 \vec{p}^2 = m^2$

#### Skill: Four vectors & co- and contra-variance



 $x^{\mu} = (ct, \vec{x})$  $p^{\mu} = (E, \vec{p})$ 

But covariant derivative:

$$\partial^{\mu} = \left(\frac{1}{c}\frac{\partial}{\partial t}, -\vec{\nabla}\right)$$

• Covariant vector:

$$x_{\mu} = (ct, -\vec{x})$$
$$p_{\mu} = (E, -\vec{p})$$

But covariant derivative:

$$\partial_{\mu} = \left(\frac{1}{c}\frac{\partial}{\partial t}, \vec{\nabla}\right)$$

Note that the minus sign is "opposite" to the case of the coordinate four-vectors.

• Use cases:

$$\partial_{\mu}A^{\mu} = \partial^{\mu}A_{\mu} = \frac{\partial A^{0}}{\partial t} + \frac{\partial A^{1}}{\partial x} + \frac{\partial A^{2}}{\partial y} + \frac{\partial A^{3}}{\partial z} \qquad \qquad \partial_{\mu}\partial^{\mu}\phi = \frac{\partial\phi}{\partial t} - \frac{\partial\phi}{\partial x} - \frac{\partial\phi}{\partial y} - \frac{\partial\phi}{\partial z}$$

### Skill: Dirac Gamma Matrices

- Dirac  $\gamma$  matrices:  $\{\gamma^{\mu}, \gamma^{\nu}\} = 2g^{\mu\nu}$  in 4x4 matrices.
  - We will use the Dirac-Pauli representation

$$\gamma^{0} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \qquad \gamma^{1} = \begin{pmatrix} 0 & 0 & 0 & 1 \\ 0 & 0 & 1 & 0 \\ 0 & -1 & 0 & 0 \\ -1 & 0 & 0 & 0 \end{pmatrix}$$
$$\gamma^{2} = \begin{pmatrix} 0 & 0 & 0 & -i \\ 0 & 0 & i & 0 \\ 0 & i & 0 & 0 \\ -i & 0 & 0 & 0 \end{pmatrix} \qquad \gamma^{3} = \begin{pmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & -1 \\ -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \end{pmatrix}$$

Note the indices: (confusing!)

 $\mu, \nu = 0, 1, 2, 3$  are the *Lorentz indices in space-time*:

Dirac matrix indices: 1,2,3,4 Have to do with the row and column indices of the matrix (and spinors)

Or:  $\gamma^0 = \beta = \begin{pmatrix} \mathbb{1}_2 & 0 \\ 0 & -\mathbb{1}_2 \end{pmatrix}$  and  $\gamma^k = \beta \alpha_k = \begin{pmatrix} 0 & \sigma_k \\ -\sigma_k & 0 \end{pmatrix}$  with Pauli matrices  $\sigma_k$ 

Define also the chirality matrix:  $\gamma^5 = i\gamma^0\gamma^1\gamma^2\gamma^3$ 

• Note: although the gamma indices are Lorentz-indices ("space-time", the gamma-matrices are not 4-vectors! (They are simply constants.)

# Skills: Dirac delta function



• For a function f(x) we have:  $f(x)\delta(x) = f(0)\delta(x)$ 

...and therefore: 
$$\int_{-\infty}^{\infty} f(x)\delta(x) \, dx = f(0) \int_{-\infty}^{\infty} \delta(x) \, dx = f(0)$$

• An important representation of the Dirac delta function is:

$$\delta(x) = \frac{1}{2\pi} \int_{-\infty}^{\infty} e^{ikx} \, \mathrm{d}k$$